

The Magnetic Vortex during a Solar Eclipse

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Abstract

- Researchers are intensively working to understand the phenomena that remain inadequately explained in connection
- with solar eclipses, including wind fluctuations, gravity waves, and oscillations in the Foucault pendulum. Despite
- the numerous theories that have been put up to elucidate these changes, substantial evidence remains necessary to
- substantiate any of them. Studies indicate that these alterations transpire more frequently at the core of the shadow
- and are observable throughout all atmospheric strata. Nevertheless, no research has concentrated on thoroughly
- examining the complete lunar umbra cone, encompassing the influence of the moon's wake phenomena. The lunar
- umbra cone is a common structure at both ends. This work highlights the parallels between changes in wind and
- gravity waves during solar eclipses and interplanetary magnetic field (IMF) lines. Building on our analysis, we
- propose a novel concept that offers a more cohesive and effective explanation of these phenomena during eclipses.
- This explanation is based on the formation of the magnetic vortex generated by the IMF lines extending from behind
- the moon to the Earth.

1 Introduction

- Solar eclipses arise when the moon intersects the line between the sun and the earth, blocking out sunlight in
- particular areas of the planet's surface depending on its path. This celestial occurrence gives a singular chance for
- researchers to observe and examine the consequences of this phenomenon. Several mysterious phenomena appear
- during this event, for which we still lack a full explanation. Among these phenomena is the Allais (1957) effect,
- which shows oscillations of the Foucault pendulum during a solar eclipse. This oscillation was regarded as a novel
- domain distinct from classical gravitational theory. This phenomenon generated global scientific attention,
- prompting the suggestion of various theories, including the concept that wind and atmospheric pressure alterations
- may affect the pendulum's oscillations. The hypothesis that the supersonic movement of the moon's shadow could
- affect the pendulum's oscillations was dismissed. "Eclipse winds" are another poorly understood phenomenon that
- occurs during a solar eclipse, manifesting as light tornadoes (Aplin and Harrison, 2003; Eaton et al., 1997; Gray and
- Harrison, 2012) and are distinguished by their unstable wind patterns. During a solar eclipse, ionospheric layers
- exhibit upward movements, especially in regions of maximum obscuration, accompanied by a significant drift in
- total electron content (TEC), Furthermore, numerous other variations.
- During a solar eclipse, the moon's shadow moves across the Earth's atmosphere, affecting different layers of it (Anderson, 1999). Chimonas & Hins (1970) proposed that the moon's shadow movement, which exceeds the speed of sound has a cooling impact and disrupts the equilibrium of the atmosphere, creating gravity waves. Since then, these waves have received expert attention, and studies have intensified to determine their source. Reports indicate the emergence of gravity waves due to a solar eclipse in various atmospheric areas, derived from disparate observational evidence. Experimental findings demonstrate that gravity waves travel from the stratosphere to the ionosphere, which exhibits low ozone concentrations due to inadequate solar radiation (Fritts and Luo, 1993). Some studies (Altadill et al., 2001; Gerasopoulos et al., 2008; Jakowski et al., 2008; Manju et al., 2012; Šauli et al., 2006) suggest that the ionosphere may generate these waves. Farges et al. (2003) identified two origins for the waves: the 40 thermosphere and the lower atmosphere. These observations are consistent with the work of Šauli et al. (2007), who documented the emission of two waves: one ascending from an altitude of 200 km and the other descending to below 200 km. Gomez (2021) conducted a recent study that demonstrated the vertical travel of these waves through all layers, from the troposphere to the thermosphere, and their potential to reach altitudes exceeding 300 km within the umbra. These waves are known to have several properties, including an increase in amplitude with height (Fritts and Luo, 1993). Jones et al. (2004) stated that the waves synchronize with the moon's shadow, exhibiting increased strength and greater frequency during the phase of maximum obscuration (Lin et al., 2018). This indicates that the eclipse is the origin of these waves and that it synchronizes their movement with the moon's umbra (Nayak and Yiğit, 2018). The evidence suggests that gravity waves are stronger and more intense during the maximum obscuration phase, so it is important to closely examine them to help us understand the mysterious phenomena and
- changes that occur during a solar eclipse. Initially, we disclose the impact of a solar eclipse on the atmosphere and

- the earth's surface, as well as the subsequent events that predominantly occur within the central shadow. Secondly,
- after examining this literature, we distill the most significant findings and insights regarding the lunar umbra.

2 Experiments and results during a solar eclipse

2.1 Gravity waves and their properties in maximum obscuration

- In addition to the property of increasing the wave amplitude with height, we add in this section the most important
- known features of gravity waves that appear during solar eclipses. According to Šauli et al. (2007), gravity waves
- manifest vertically, and their propagation direction is oblique, they also discovered that these waves exhibit both
- upward and downward motion. In his work, Venkat Ratnam et al. (2012) established that the hodograph analysis
- demonstrated the wave's upward motion and clockwise rotation, along with its downward and counterclockwise
- rotation. This aligns with Sheng-Yang Gu's (2023) observation that gravity waves show more intensity on the north
- side of the central eclipse than on the south side. Several studies also address a related topic, observing simultaneous
- upward and downward energy-pulling processes (Barad and Sripathi, 2023; Chen et al., 2021). Chen et al. (2014)
- and Manju et al. (2014) uncovered another characteristic of gravity waves, demonstrating its elliptical polarization
- through their research.

2.2 During the maximum obscuration, the atmosphere changes

 Here, we summarize the key details of the variations that occur at different atmospheric layers during a solar eclipse. The data indicate that there has been a discernible drop in ozone levels, particularly during the period of maximum obscuration (Girach et al., 2012; Manchanda et al., 2012; Zerefos et al., 2007). Pradipta et al. (2018) and Chen et al. (2019) indicated a progressive reduction in total electron content (TEC) during the solar eclipse phase. The TEC drop value can reach a maximum of 60% within the umbra (Aa et al., 2020; Coster et al., 2017) this percentage is deemed elevated compared to other locations. Numerous studies have documented clear disturbances in the ionosphere layers during a solar eclipse. For instance, during the 2009 eclipse in China, Gang Chen et al. (2011) observed that gravity waves dragged the ES layer upward. In the same vein, J. Wang et al. (2021) verified that the highest upward lift rate of the ES layer occurs at maximum obscuration in Xiamen, China (24.2°N, 118.07°E, 97.8% obscuration), and it is lower in places where the obscuration ratio is low. The upward pulling process may approach 400 km in the F and E layers, particularly at the core of the eclipse shadow compared to the surrounding areas (Pradipta et al., 2018). According to Martínez et al. (2020), at Temuco station (38.7°S, 72.6°W), the F1 and E ionospheric layers experience the most significant reduction, while the TEC drift and critical frequencies exhibit the highest changes when the obscuration ratio is 100%, in contrast to other stations with lower obscuration ratios. Tian

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- (2022) pointed out the upward movement in the F2 layer during the first stage of entering the lunar shadow and a
- subsequent decline after the shadow leaves.
- 82 During an eclipse, intense gravity waves propagate from the lower atmosphere to the ionosphere, facilitating the
- occurrence of traveling ionospheric disturbances (TIDS). These TIDS follow the moon's shadow path and appear as
- curved wavefronts (Jones, 1999; Yan et al., 2021). Sun et al. (2018) examined the emergence of the ionosphere bow

85 wave as a feature associated with the onset of maximum obscuration. Zhang et al. (2017) state that its structure is

elliptical.

2.3 The eclipse cyclone is located within the path of the moon's shadow

- The progress made in elucidating the relationship between gravity waves and wind system changes during a solar
- eclipse, along with their motion characteristics, is summarized in this section. Historically, the solar eclipse of May
- 28, 1900, that crossed the United States of America marked the beginning of the observation of the eclipse cyclone
- 91 phenomenon by Clayton (1901), which concluded by noting the appearance of rings rotating in opposite directions
- from each other around the totality region and differing in radial distance. However, the scientific community at the
- time did not agree with Clayton's theory (Aplin et al., 2016). Almost a century later, Müller-Wodarg (1998)
- confirmed that the winds show counterclockwise rotational motion around the low-pressure zone. In the same
- 95 context, a simulation conducted by Prenosil (2000) of a solar eclipse traversing Europe exhibited similar features to
- 96 those of an eclipse cyclone. Similarly, Aplin & Harrison's (2003) model closely agrees with the Clayton theory. This
- model features a chilly core at its center, encircled by a clockwise-rotating ring. It possesses an autonomous internal
- rotation in a counterclockwise direction, with a diameter of approximately 170 km.
- During the maximum obscuration of a solar eclipse in India, researchers detected wind disturbances in two different
- levels of the atmosphere, accompanied by the elliptical motion of winds. In addition, the investigation revealed the
- presence of vertical wavelengths in the troposphere that might approach 6 km in wind structure (Ramkumar et al.,
- 2013), as is widely known, gravity waves can create winds (Hines, 1974). In recent studies, the eclipse winds have
- been better explained based on multiple observations, where they are perpendicular and appear at the first contact by
- entering the region of totality, persisting alongside the lunar shadow it shows downward winds in the anterior arch
- and upward winds in the posterior arch at an altitude of 400 km (Cnossen et al., 2019). Piscitelli & Saurral (2021)

also observe that when the local area surpasses its maximum obscuration, the speed of these winds increases.

2.4 The elliptical motion of a pendulum occurs at the maximum obscuration

- To provide comprehensive insights into the analysis and linking of ideas and observations associated with the solar
- eclipse, we present in this section the most important experiments of the Foucault pendulum, as some reports must
- be considered and these individual results cannot be ignored. Researchers have performed independent experiments
- to observe this impact inside and outside the lunar umbra. During the solar eclipse in Mexico, researchers
- experimented with a paraconic pendulum, which could rotate about its axis and oscillate in all directions. The
- experiment occurred in Pato-Branco, located at the lunar umbra, Savrov (1991) noted a pendulum transition from a
- linear path to an elliptical trajectory. Additionally, the rotation rate increased significantly by three times for the
- local Foucault effect around the azimuth of the major axis. The observation of elliptical motion is consistent with the
- findings of Kuusela's (1992) experiment, which was conducted on the same day as Savrov's experiment, specifically
- in the lunar umbra, during the experiment, the sun's height approached 80°. On November 3, 1994, Savrov repeated
- the experiment utilizing two pendulums during the maximum obscuration phase of a solar eclipse in Brazil and saw
- an identical effect. The elliptical motion emerged when the pendulum's rotation speed increased, surpassing the

- local Foucault effect. Despite the clarity of these results, the existence of this effect remained subject to
- disagreement, as it has been considered that such effects during eclipses may be due to errors in the instruments,
- measurements, or calculations. However, the observers continued their experiments and monitored the pendulum's
- movement, as Olenici & Pugach (2012) used a pendulum and a torsion balance during the partial eclipse in their
- experiment to confirm Allais' observation, their observations revealed convex curves, suggesting this force could
- rotate the torsion balance. Nonetheless, Olenici et al. (2014) did not settle for that limit, but instead repeated the
- experiment, reaffirming their results with the presence of the effect.

3 Key observations

- This section will emphasize the most intriguing observations noted in the introduction and sections about the impact of solar eclipses on planet Earth. Results show that the largest drift in TEC and critical frequencies occurs after the maximum obscuration, i.e., in the lunar umbra, and this is the region where the highest upward ionospheric drag occurs, as the origin of gravity waves is linked to the lunar umbra, which moves in sync with it. Within this region, gravity waves are characterized by vertical orientation and oblique propagation, and the wave amplitude increases
- with height. We also observe a high-frequency flow, which gradually decreases in the surrounding regions as we
- move away from the maximum obscuration. The results also show that the waves rotate clockwise upward and
- counterclockwise downward. Where the lunar umbra cone contains these waves, Studies have shown that gravity
- waves within the maximum obscuration have a vertical wavelength of up to 90 km and travel to heights of up to 300
- km, allowing them to pass through all layers of the atmosphere. Therefore, the wind structure exhibits vertical
- structure and elliptical rotation due to the influence of gravity waves. In the same context, the pendulum undergoes
- an elliptical movement whenever the experiment falls within the maximum obscuration (The lunar umbra).
- These observations lead us to consider the possibility that the source of the influence on these waves is outside the
- atmosphere, specifically in the lunar umbra. The question is whether these waves will ascend further. What types of
- physical phenomena may manifest in this area? Consequently, we must examine the physical processes occurring
- within this conical component in general.

4 The modifications in the IMF lines within the lunar umbra

- In this section, we will attempt to shed light on the most notable results in the structure of the lunar umbra near the
- moon, as many studies have demonstrated the formation of a structure directly behind it, known as the lunar wake,
- upon absorption of the solar wind by the lunar surface, the IMF lines traverse unimpeded to the lunar nightside,
- 148 thereby forming a complicated structure in that area (Lyon et al., 1967).
- The substantial augmentation of the magnetic field, especially within the central lunar wake, is contingent upon the
- angle of interaction between the solar wind and the IMF lines. In the perpendicular case, the expansion of the IMF is
- restricted to the nearby moon. In the case of parallel solar winds, there is an increase in the magnetic field, and it
- extends to long distances downstream (Holmström et al., 2012; Wang et al., 2011; Xie et al., 2013; Zhang et al.,
- 2012). The observed phenomenon results from the electric field surrounding the shadow cone created by the

- 154 accelerated motion of electrons behind the moon (Birch and Chapman, 2001, 2002). The electric current creates
- closed cylinder loops perpendicular to the IMF lines, putting pressure on them as illustrated in Fig 4 of the study by
- Fatemi et al. (2013). The application of pressure causes the IMF lines to bend, resulting in an axially symmetric
- configuration centered on the central lunar wake axis (see Fig 1). As a result, it increases the intensity of the
- magnetic field on the far side of the lunar wake, we also note that there is a rotation of the field attributed to the
- Mach cone (Michel, 1968; Owen et al., 1996).
- Observations suggest that the IMF within the lunar wake (The lunar umbra) exhibits a clockwise rotation at Z>0 and
- a counterclockwise rotation at Z<0, It is also depicted in Fig 14 of Zhang et al. (2014). On the other hand, the lunar
- satellite Kaguya observed a reflection of magnetic field rotation between the Northern and Southern Hemispheres at
- a 90° angle in the lunar wake, i.e., in the lunar umbra. Simultaneously, solar winds were found to align parallel to
- the magnetic field, and the polarization exhibited an extremely elliptical nature (Nakagawa et al., 2021).

Figure 1: The white lines represent IMF, and as they pass the moon, the electrical current indicated by the red rings

forces these lines to bend at the edges of the Umbra lunar cone, as illustrated by the two blue lines.

5 Similarities

This section will compare the results and observations gathered regarding gravity waves and the IMF. On the lunar

nightside, the IMF lines are bent inward, which leads to a significant increase in the expansion of the magnetic field

- downstream. First, this bending is consistent with gravity waves' oblique propagation and vertical direction during a
- solar eclipse. Second, the stretched IMF and the high-frequency gravity waves share in the umbra cone. Another
- comparison reveals that the IMF lines behind the moon which represents the lunar umbra cone, rotate in two
- directions: clockwise and counterclockwise, signifying a push and pull force, that aligns with the upward and

- downward motion of gravity waves and the wind movement during the maximum obscuration phase. On the other
- 176 hand, the satellite Kaguya results indicate the presence of an elliptical motion behind the moon at an angle of 90°,
- which is consistent with the elliptical polarization of gravity waves, which in turn stimulates the elliptical motion of
- the winds. In addition, this result is consistent with the elliptical motion of the Foucault pendulum when the
- obscuration ratio is 100% (the lunar umbra). The temporal dependence in comparing the duration of disturbances
- and oscillations during the eclipse with the vibration period in the IMF structure behind the moon is also significant.
- This comparison relies on data obtained by the ARTEMIS mission, which indicated that the duration ranged from 5
- to 60 minutes on the edge of the lunar wake (Poppe et al., 2014). Researchers typically estimate this period to be
- between 40 and 60 minutes after the eclipse maximum (Altadill et al., 2001; Chen et al., 2015; Farges et al., 2003;
- Wang et al., 2021), highlighting the importance of this time in analyzing atmospheric fluctuations during eclipses.
- **6 Discussion and conclusions**
- The IMF lines are similar to magnetohydrodynamics (MHD) in that they interact with the solar wind and other
- matter in space. As a result of this interaction, the magnetic field lines can become curved, becoming parallel,
- creating what is known as magnetic vortices (Durand-Manterola and Flandes, 2022). It is known in scientific circles
- that the pressure of electric current on vertical magnetic lines, when they are in the form of closed loops leads to the
- bending of these lines, which become parallel around a fixed axis and end in the form of a vortex. Thus, this
- description is consistent with the pressure exerted by the electric current in space on the IMF lines, causing them to
- bend inward (refer to sectional figure 2). This increases the expansion of the magnetic field in the central lunar wake
- axis, which means a magnetic vortex. Observing the Kaguya satellite, which showed an elliptical rotation at a 90°
- angle within the center of the wake, confirms the formation of a magnetic vortex behind the moon.

 Figure 2: Depicts a cross-section of the lunar umbra cone, with the white color indicating the symmetrical and curved IMF lines, the large blue circle representing the cone's boundaries, and the small circle representing the shadow's center (The maximum obscuration).

- It is worth noting that the IMF lines at the edges of the lunar umbra cone rotate in two directions: clockwise and counterclockwise, indicating a push and pull force. This interaction is also an indicator of the presence of a magnetic vortex. From this perspective, we can examine the lowest section of this cone as a shared component of the lunar umbra. By identifying common characteristics, we intend to demonstrate that vortex structure may extend to the atmosphere and the Earth's surface. Models of vortex formation can elucidate this pattern, especially when considering the effects of magnetic fields.
- Our literature has noted the elliptical motion of gravity waves and winds in the atmosphere during eclipses.
- Additionally, the pendulum experiences elliptical motion as its rotation rate increases. This rotation property is
- regarded as one of the fundamentals of the vortex, providing an essential explanation for the core of any vortex,
- which represents the umbra region. A dominant feature of vortices is their coordinated upward and downward
- motion, indicating a clockwise direction during ascent and a counterclockwise direction during descent. This
- property is consistent during eclipses with the motion of IMF lines, gravity waves, and winds even at altitudes of up
- to 400 km around the region of totality. Furthermore, in vortices or tornadoes, an intersection exists between phase
- motion and energy transfer, as energy accumulates and delineates a certain trajectory (such as the descent of cold
- air). At the same time, other structural components revolve circularly around the center. This leads us to the research
- indicating that cold air descends on the umbra area during a solar eclipse, replacing warm air (Elmhamdi et al.,
- 2024). This serves as additional evidence of the presence of a vortex structure in the atmosphere.
- Given our reliance on models of cyclones and magnetic vortices for comparison, we can highlight the effect of the vortex and its magnetic field on charged particles. This field forces the particles to move in helical paths within this conical structure at an inclined angle, forming a structure similar to the motion of these particles within it. Additionally, this field can elevate plasma to higher layers, increasing its flow toward upper regions. Some studies have indicated the occurrence of this phenomenon in the atmosphere, particularly around the shadow center. For 221 instance, Wang et al. (2019) noted that winds near the shadow center contribute to lifting plasma upward, altering its density. Moreover, Tian et al. (2022) demonstrated that the F2 layer exhibited a tilted structure, with plasma rising in the west while decreasing in the east. The results also showed that plasma ascended during the initial contact of 224 the obscuration phase. In this context, a study by Manju et al. (2014) revealed that horizontal winds exhibit an inclined elliptical shape, based on hodograph analysis of the ionospheric E layer. These findings are supported by studies by Jie Wang et al. (2023) and Chi-Yen Lin et al. (2021) The former reported ascending and descending echo structures with inclined geometry during the solar eclipse on June 21, 2020, while the latter suggested that the moon's shadow could cause horizontal tilting in the ionosphere. These studies reflect ascending and descending fields at inclined angles, which aligns with the concept of a magnetic vortex structure. Thus, we believe that the changes in gravity waves, winds, and the Allais effect during a solar eclipse are mostly attributable to the IMF lines inside the lunar umbra cone. We consider IMF lines can affect charged particles by moving them toward the upper

- layers, creating atmospheric pressure, and thus causing variations in wind patterns and gravity waves. Since the IMF
- lines are vortex-shaped, they force the wind to direct its movement as a cyclone.
- The lower temperature and the velocity of the moon's shadow are frequently posited as the primary variables
- contributing to the oscillations observed during a solar eclipse. A series of accurate measurements in China utilizing
- the Lacoste-Romberg gravimeter during the solar eclipse on March 9, 1997, indicated a reduction in the gravity
- 237 value of approximately 0.7 millionths of a percent of the standard gravity value (Yang and Wang, 2002). This
- reduction was attributed to the supersonic transit of the moon's shadow, temperature, and barometric pressure
- fluctuations. Despite barometric pressure, the influence of the moving shadow in the atmosphere remains a hundred
- thousand times insufficient to account for the retrograde motion of a pendulum (Bagdoo, 2023).
- From the perspective of theoretical analysis, It is believed that the start of darkness induces a decrease in
- temperature, resulting in the cooling of the ozone layer and subsequently the emergence of gravity waves. We wish
- to cite a comparable instance, the Terminator, which traverses at a velocity exceeding the speed of sound before
- dusk. The absence of light lowers the environment's temperature and facilitates the generation of standard gravity
- waves. Nevertheless, these waves are often less intense and less pronounced, exhibiting a maximum vertical
- wavelength of 30 km (Preusse et al., 2002). They can ascend to the mesosphere under optimal atmospheric
- conditions, although the cooling duration exceeds the maximum obscuration phase of a solar eclipse. Conversely,
- gravity waves are more intense during a solar eclipse, propagating vertically and horizontally and attaining greater

- heights. As stated in Section 2.3, these waves also contribute to the associated consequences of winds. The
- manifestation of elliptical motion in the umbra should also be noted.

 Figure 3: The white lines represent the IMF. After crossing the moon, they bend inward to form a magnetic vortex structure at the edges of the lunar umbra cone, which extends to the Earth.

 Our comparison reveals a distinct resemblance and motion consistency between the IMF lines, the gravity, and the changes in the winds inside the lunar umbra cone. Figure 3 shows our perspective based on the analysis of observations, where we suggest that during the solar eclipse, there is a magnetic vortex stretching from the region behind the moon to the Earth's surface. Conical vortices are common in nature. Examples include solar magnetic tornadoes and plasma vortices flowing in strong magnetic fields, such as those found in black holes or electrical discharges in nuclear fusion experiments.

- According to our proposal, this conical structure moves in tandem with the lunar umbra cone, with the leading edge
- functioning as an arched barrier of rising IMF lines, thereby influencing the atmospheric layers, the idea of a
- magnetic vortex fits with the Clayton cyclone theory that was seen during a solar eclipse. This conclusion also fits
- with the concept proposed by Olenici et al. (2014), who say solar influences cause a vortex-like effect during the
- eclipse phase. However, they did not explain the reasons behind this phenomenon and its formation.
- Through this combination of observations in outer space, the atmosphere, and on the Earth's surface, this study
- presents evidence that the IMF lines link to the Earth during a solar eclipse, forming a vortex at the edges of the
- lunar umbra cone. This paper is the first work to indicate the existence of a magnetic vortex during a solar eclipse,
- so we consider it suitable for future experimental applications to verify this concept. Therefore, we suggest
- examining the elliptical motion at two or three altitudes within the lunar umbra cone. The first method involves

- conducting a pendulum experiment on the earth's surface, the second involves using balloons or rockets inside the
- atmosphere, and the third method consists of having a satellite outside the atmosphere, and making sure that the
- three experiments happen at the same time and on the same vertical path. We recommend selecting the height of the
- Sun during the eclipse closer to 90°. We should also consider other factors that may affect the structure of the
- vortex, such as the earth's position in the year, the speed of the solar wind, the time of day, and geography.
- Therefore, the recurrence of the elliptical motion at these three points can prove the validity of this theory. This
- theory is anticipated to establish a new avenue for researchers to explore the interaction of IMF lines within the
- atmosphere and utilize the magnetic vortex structure to elucidate numerous poorly known phenomena that transpire
- during solar eclipses.

Data Availability Statement

- This study does not require data sharing because it relies solely on theoretical analysis and does not create or
- analyze any new data.
- **Competing interests**
- The contact author has declared that none of the authors has any competing interests.
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