



19 March 25 was detected. The TEC data calculated from a first principles model SD-WACCM-X were
20 also analyzed using the same method as that for the observational data. No TEC anomaly was found on
21 March 25 from the model outputs either. Thus the source of the TEC anomaly on March 25 is unlikely
22 from the lower atmosphere waves. In this study, we show the occurrence of TEC anomaly on March 25,
23 10 days before the Mw7.2 Mexico earthquake and this TEC anomaly may not be explained by lower
24 atmosphere or geomagnetic activity forcing.

25 **Key words:** GPS TEC, ionospheric TEC anomaly, Mw7.2 Mexico earthquake

26 **1. Introduction**

27 The abnormal ionospheric density variations before and/or after earthquakes have attracted much
28 attention from the geophysicists for many years (e.g., Pulinets & Boyarchuk, 2004a; Le et al., 2015).
29 However, identifying and determining ionospheric density disturbances that are associated with an
30 earthquake have been challenging so far. For instance, Heki (2011) reported the enhancement of
31 ionospheric total electron content (TEC) ~40 min before the 2011 Mw9.0 Tohoku-oki earthquake.
32 However, in subsequent studies, some scientists questioned the true cause of this pre-seismic abnormality
33 in their comments (Kamogawa and Kakinami, 2013; Utada and Shimizu, 2014; Masci et al., 2015).
34 Heki and Enomoto (2013, 2015) later applied several data analysis methods and more case studies to
35 demonstrate the correlation between pre-seismic TEC abnormality and the earthquake. Despite these
36 controversies, several data analysis methods have been employed to explore potential



37 seismo-ionospheric perturbations in previous studies. The running mean method is a common approach
38 in analyzing time series data. Liu et al. (2000) utilized the running median of the critical frequency of
39 the ionospheric F₂ layer (f_oF_2) and the inter-quartile range (IQR) of f_oF_2 as the upper and lower bounds
40 to extract seismo-ionospheric precursors that may be associated with $M \geq 6.0$ earthquakes around Taiwan
41 from 1994 to 1999. Pulinets et al. (2005) calculated the monthly mean (M) of ionospheric TEC, and used
42 $M \pm \sigma$ as the thresholds to find TEC disturbances before the Colima Mexico earthquake, where σ is the
43 standard deviation. In order to obtain the location characteristics of the ionospheric perturbations
44 associated with earthquakes, spatial analysis methods have been applied in some researches. Liu et al.
45 (2011) studied the locations of extreme TEC anomalies (enhancements or depletions) in the 12 2-hour
46 intervals for a day. They compared the data with previous 30-day data in each grid to determine if they
47 are maximum or minimum values in that 30-day period, and to see whether these TEC anomalies occur
48 only nearby the earthquake region or randomly worldwide. Liu et al. (2016) found, by calculating the
49 spatial relative changes, that ionospheric TEC, electron and ion densities were simultaneously enhanced
50 at different altitudes near the epicenter of the 2005 Sumatra Indonesia Ms 7.2 earthquake. A correlation
51 analysis between different stations was applied to demonstrate the local disturbance near the epicenter.
52 Pulinets et al. (2004b) calculated the cross-correlation coefficient between two measurement points
53 located inside or outside the earthquake preparation zone, and found that the coefficient sharply dropped
54 before strong seismic shocks. Iwata and Umeno (2016) detected the preseismic TEC anomalies before



55 the main shock, foreshock and aftershock of 2011 Tohoku-Oki M_w 9.0 earthquake by correlation
56 analysis.

57 Statistics analysis of seismo-ionospheric disturbances has also been attempted by scientists when
58 there are sufficient data. Liu et al. (2010) utilized the z -test for 150 $M \geq 5.0$ earthquakes during
59 2001-2007 to try to correlate the change of the ionospheric equatorial ionization anomaly (EIA) with
60 earthquakes. Parrot (2012) applied a software to automatically detect the abrupt enhancement of ion
61 density observed by the Detection of Electro-Magnetic Emissions Transmitted from Earthquake
62 Regions (DEMETER) satellite. Based on the statistical analysis of 17,366 $M > 4.8$ earthquakes, he found
63 that perturbations in ionospheric ion density before earthquakes are more obvious than those prior to
64 random selected pseudo-earthquake events. Therefore, there is not a unified and standard method to
65 extract ionospheric density anomalies that may be related to earthquakes.

66 The ionosphere shows strong variability of different temporal and spatial scales. This variability
67 can be of different sources, including the effects of large-scale lower atmospheric waves, geomagnetic
68 and solar activity, and possibly, the earthquakes. Some ionospheric oscillations have known periods and
69 zonal structures (e.g., Forbes et al., 2008; Pancheva & Mukhtarov, 2012; Luan et al., 2012). In this study,
70 we use a new approach to extract possible ionospheric anomalies related to earthquakes. We obtain
71 TEC residuals by removing the known and identified oscillations in the ionosphere TEC data. Since
72 earthquakes are mostly single occurrence events at particular locations and times, these TEC residuals



73 can manifest earthquake effects on the ionosphere better. We use the TEC data from the madrigal
74 database at the Massachusetts Institute of Technology (MIT) Haystack Observatory. The TEC data have
75 high temporal resolution from a significantly large number of Global Position System (GPS) stations
76 (about 1500 sites from 2000, now almost 6000 sites) all over the world. Therefore, the database can
77 provide high temporal and spatial resolution data for our analysis. In this paper, section 2 describes the
78 data and analysis method. In section 3, MIT TEC data before and after the Mw7.2 Mexico earthquake
79 occurred on April 4 2010 are analyzed to obtain ionospheric TEC perturbations. In section 4, we use
80 more observational data from other time periods and first principles numerical simulations by
81 SD-WACCM-X to show the uniqueness of the TEC disturbances that occurred before the Mw7.2
82 Mexico earthquake. Finally, the conclusions are drawn in section 5.

83 **2. Observations and the Method for Data Analysis**

84 Based on a network of worldwide GPS receivers, MIT TEC is calculated by using an automated
85 software suite (Rideout & Coster, 2006). It includes downloading data, determining satellite and
86 receiver biases, removing data outliers, mapping from slant TEC to vertical TEC, and so on. The data
87 are provided as estimates of vertical TEC in 1° by 1° grids distributed around locations where data are
88 available. The temporal resolution of the TEC maps is 5 minutes. The advantage of MIT TEC is that it
89 is strictly data driven with no underlying models that smooth out the real gradients in the TEC. In this



90 study, the TEC data are downloaded from the MIT Haystack Observatory madrigal database
91 (<http://madrigal.haystack.mit.edu/madrigal/>).

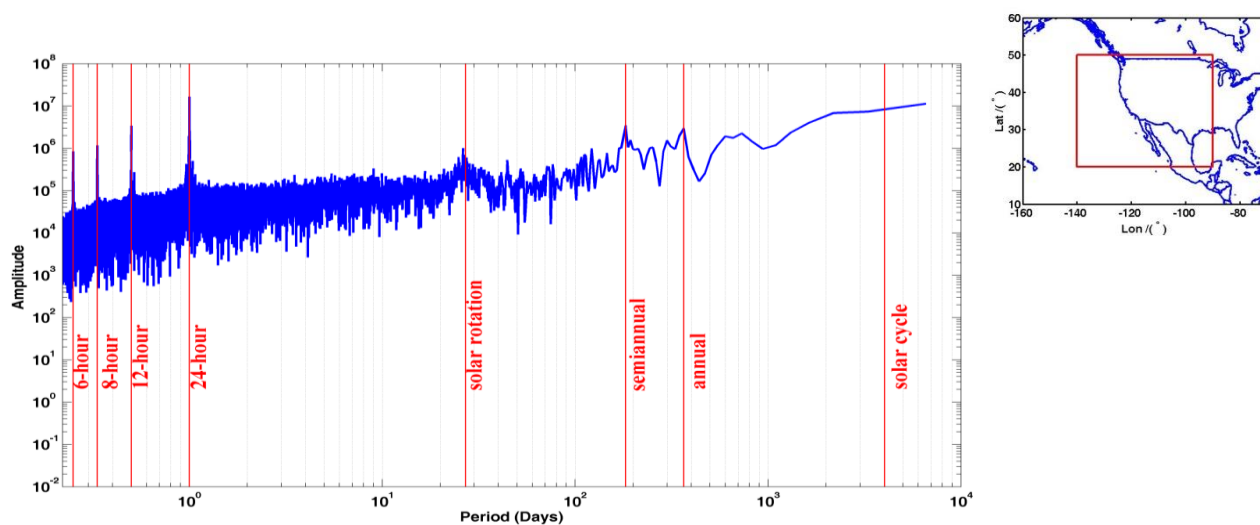
92 The Fast Fourier Transform (FFT) algorithm was applied to obtain the spectral distribution of the
93 TEC mean value in the northern American region (20°N-50°N in latitude, 90°W-140°W in longitude)
94 from 2000 to 2017 (Figure 1). Multi-day spectral peaks are seen in the figure, including 27-day solar
95 rotation, semiannual and annual oscillations. The high-frequency tidal spectral peaks at 24-hour
96 (diurnal), 12-hour (semidiurnal), 8-hour (terdiurnal) and 6-hour (quad diurnal) are also obvious in
97 Figure 1. The TEC data in each day can be expressed as a superposition of tide-like components (Forbes
98 et al., 2008; Luan et al., 2012). In this paper we used Eq. (1) to express TEC data, which includes 6
99 terms:

$$f(t)=A(0)+A(1)*\cos\left(\frac{2\pi}{24}t+B(1)\right)+A(2)*\cos\left(\frac{2\pi}{12}t+B(2)\right)+A(3)*\cos\left(\frac{2\pi}{8}t+B(3)\right) \\ +A(4)*\cos\left(\frac{2\pi}{6}t+B(4)\right)+A(5) \quad (1)$$

101 where $A(0)$ is the daily mean TEC, $A(1)$, $A(2)$, $A(3)$, $A(4)$ and $B(1)$, $B(2)$, $B(3)$, $B(4)$ are the amplitudes
102 and phases of diurnal, semidiurnal, terdiurnal, and 6-hour oscillations, respectively. $A(5)$ is the residual,
103 which includes higher frequency oscillations and/or some unknown variability, for example, the
104 perturbations caused by an earthquake. In this study, three steps were taken to obtain $A(5)$. First, if there
105 were extremely large values in the raw data, which may be caused by data error, the data were canceled
106 at its observation time. Second, a linear fitting between the solar 10.7 cm radio flux index ($F_{10.7}$), the



107 geomagnetic activity index (*AE*) and the TEC data was applied to remove solar and magnetic activity
108 effects. Similar to Pi et al. (2003) and Lei et al (2004), who used nonlinear least square minimization to
109 minimize the difference between the model results and observational data to investigate the mechanisms
110 of ionospheric variations, we also employed a nonlinear fitting method to obtain the coefficients in Eq.
111 (1) for each day in the third step. The data would not be fitted if the number of data in a day is less than
112 72 (the total number of data is 288 for each day at the 5-minute cadence) and if the data gaps are larger
113 than 6 hours. A running mean method was applied to do these fitting, with 1-day window and 1-hour
114 step, which means there will be 24 fitted data in each day. From these three steps, $A(5)$ was extracted
115 using Eq. (1) to analyze the TEC changes before and after the earthquake. Hereafter, we will show TEC
116 residual values of $A(5)$ obtained from the above described data analysis processes.

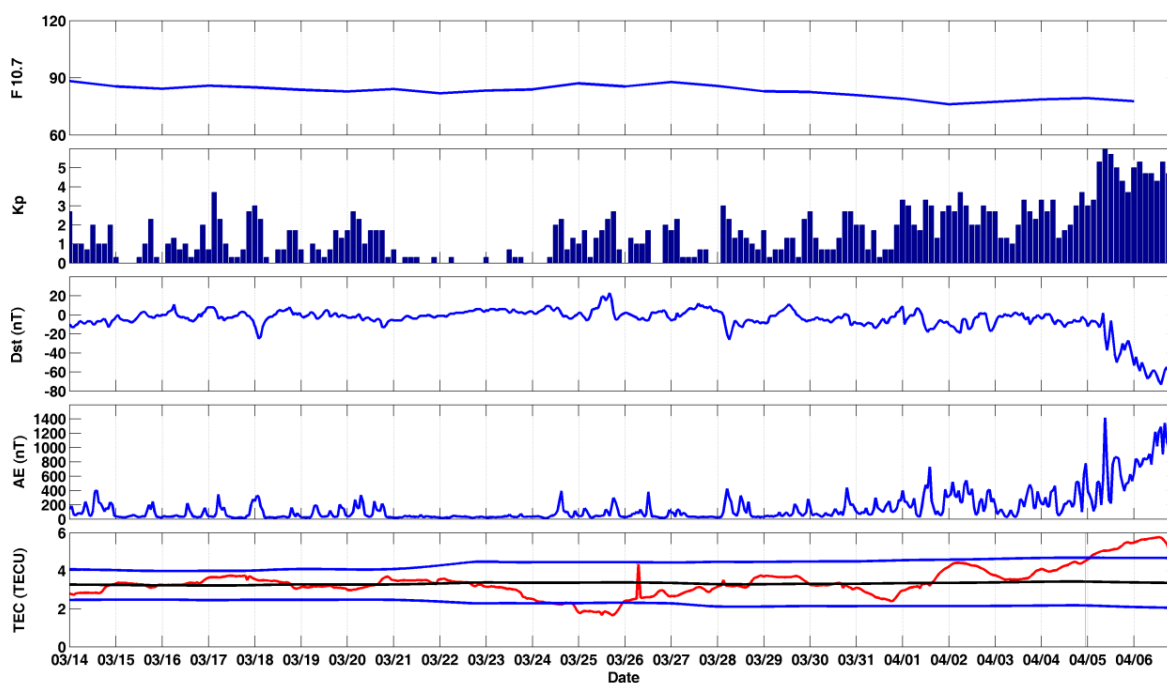


118
119 **Figure 1:** The FFT spectrum of the TEC mean values from 2000 to 2017 in the North American region (20°N-50°N in
120 latitude, 90°W-140°W in longitude), which is shown with the red rectangle in the subplot.



121 3. Results

122 The Mexico Mw7.2 earthquake with 10 km depth occurred at 22:40 UT (universal time) on April 4
123 2010. The epicenter was located at (32.286°N, 115.295°W). The TEC data around the epicenter in a
124 region with latitude from 30°N-34°N and longitude from 113°W-117°W were obtained and analyzed
125 from March 14 to April 6, 2010. The mean TEC residual in this region is shown in the bottom panel of
126 Figure 2. From the time series of $F_{10.7}$ and geomagnetic activity indices (K_p , Dst , and the AE) in Figure
127 2, it can be seen that there was no geomagnetic activity from March 14 to March 31. It became more
128 active since April 1, especially after April 5 when Dst dropped to below -40 nT.



129 **Figure 2:** Time series of TEC residual (A(5)) around the epicenter from March 14 to April 6, 2010. Panels from top to
130 bottom represent time series of $F_{10.7}$, K_p , Dst , the AE index and TEC residual, respectively. In the bottom panel, the red
131



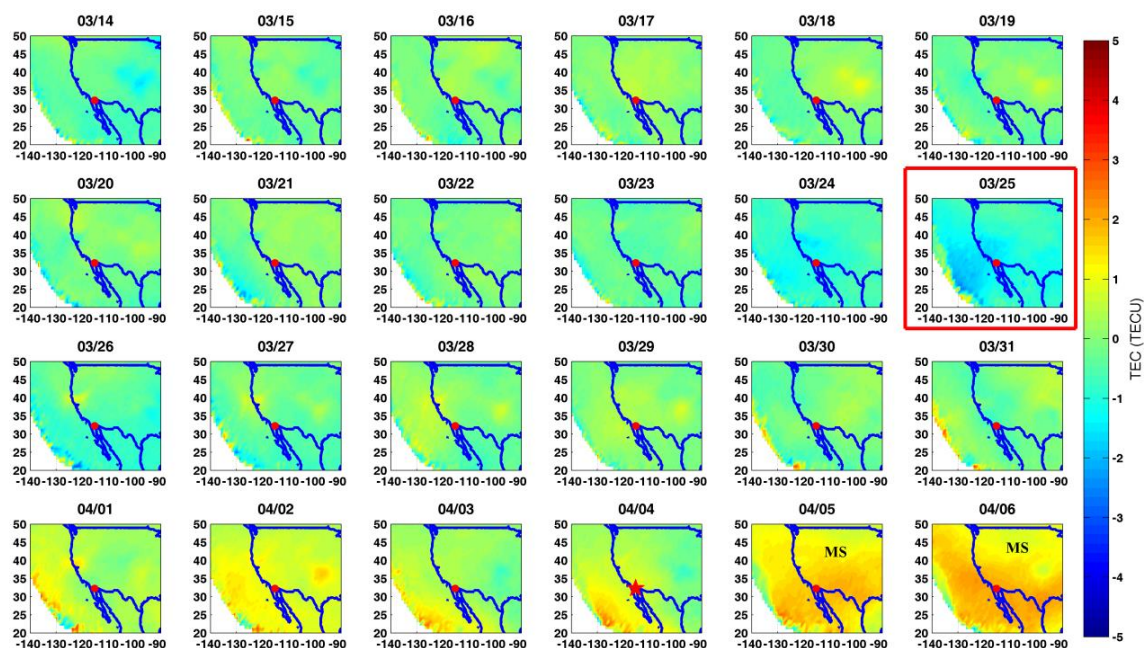
132 line is the mean value of the TEC residual in the region of latitude 30°N-34°N and longitude 113°W-117°W. The black
133 and blue lines represent the mean values and $M \pm 1.5 \cdot \sigma$ of ± 15 days of data centered around a particular day. The
134 vertical red line on April 4 indicates the occurrence time of the Mexico Mw7.2 earthquake.

135 Under the assumption of a normal distribution, the probability of data in the range of $\pm \sigma$ and $\pm 2\sigma$ is
136 68.26% and 95.44%, respectively. In order to avoid the probability being too low or high, we used
137 $M \pm 1.5 \cdot \sigma$ (the probability is 86.64%) as the threshold to extract the disturbances that may be related to
138 earthquakes, where M and σ stands for the mean value and standard deviation of TEC residuals of ± 15
139 days centered around a particular day, respectively. Before the Mexico earthquake, the TEC value was
140 lower than the threshold on March 25. In other days from March 14 to April 5, TEC residual value for
141 each day was within the thresholds. When magnetic activity became stronger, the TEC was also
142 increasing with the values over the upper bound from April 5 to 6. Therefore, except the depletion of
143 TEC residual on March 25 and the increase of TEC residual associated with the magnetic activity on
144 April 5 and 6, no TEC anomalies exceeding the thresholds were detected in other days during the
145 72-day period.

146 In order to further analyze the TEC changes potentially related to the earthquake, we expanded the
147 region of interest to include the area of latitude 20°N-50°N and longitude 90°W-140°W. In this analysis
148 of TEC spatial structure, the differences between the TEC residuals and the mean values of ± 15 -day
149 data for a particular day were obtained before and after the earthquake. For each day, the mean value of



150 the 24-hour data was used to represent the TEC residual, as shown in Figure 3. The TEC depletion on
151 March 25 is evident in the region surrounding the epicenter, similar to the analysis result of the time
152 series shown in Figure 2. The TEC depletion can also be seen south and west to the epicenter. In all the
153 days shown in Figure 3, only on March 25 did the TEC residual data show depletion in the region
154 around the epicenter. The TEC residuals began to increase from April 1 in the western and southern part
155 of the region. Large TEC values occurred in a large region when magnetic activity became strong on
156 April 5 and 6.

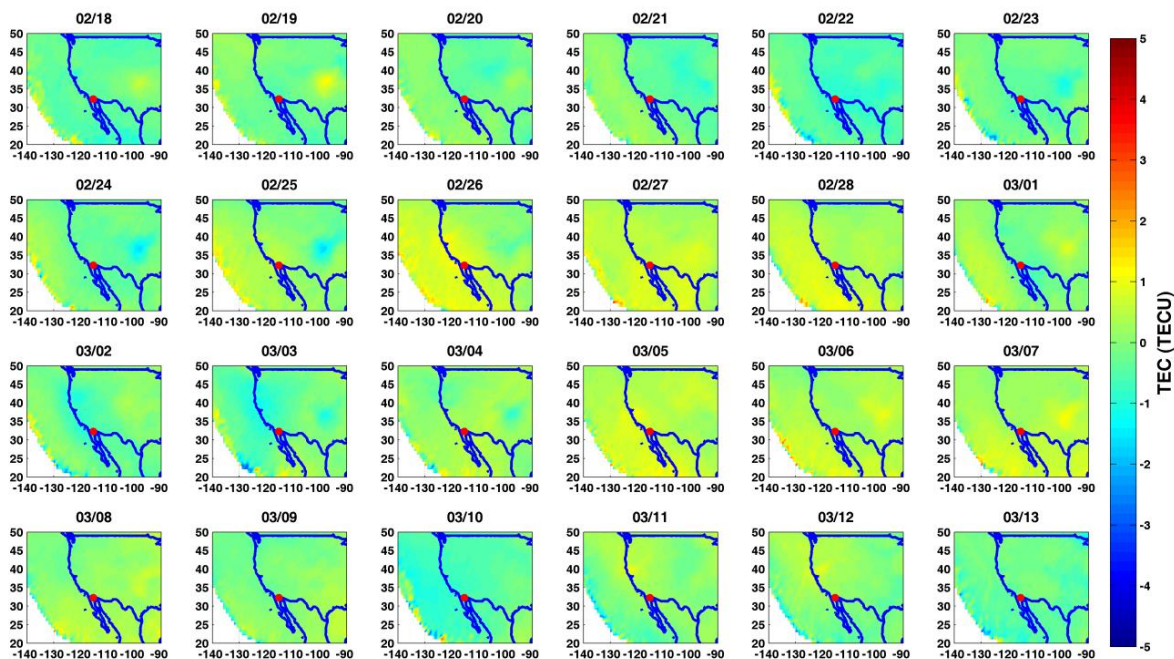


157
158 **Figure 3:** TEC changes in the region of latitude 20°N-50°N and longitude 90°W-140°W from March 14 to April 6,
159 2010. The red dot indicates the epicenter. The red star shows the day of the earthquake. The blue lines represent the
160 continent. ‘MS’ on April 5 and 6 mean ‘magnetic storm’. The date of the map is marked on the top of each subpanel.
161 The subpanel on March 25 is highlighted with the red rectangle.



162 **4. Discussion**

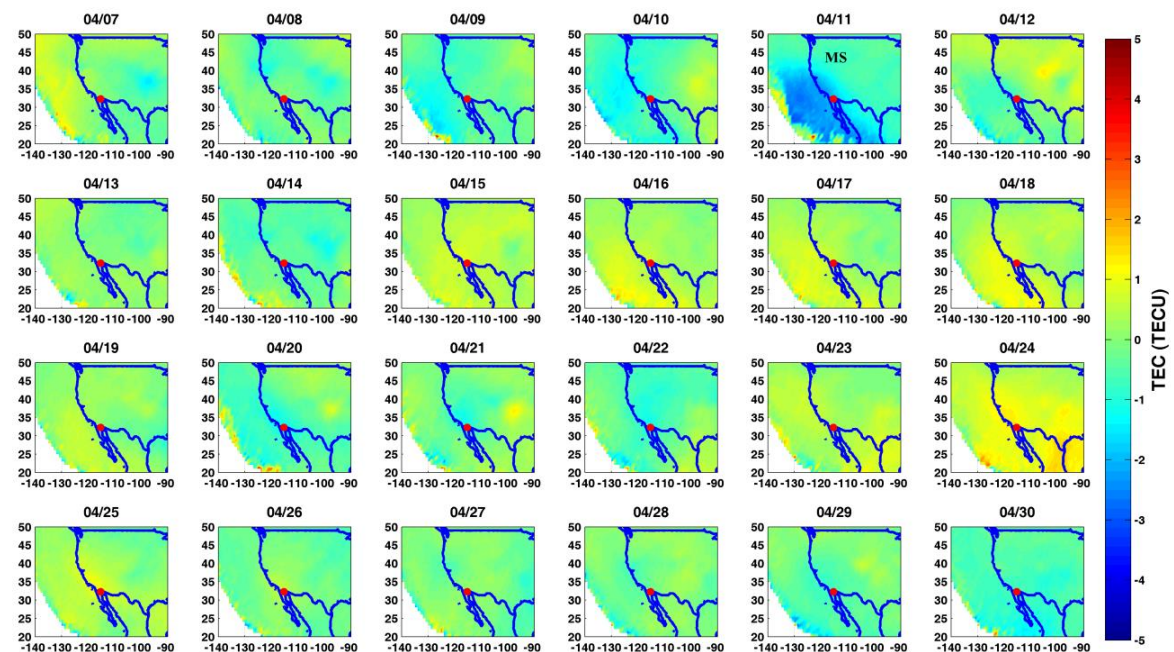
163 In order to further establish the possible correlation between the TEC depletion on March 25 and
164 the Mexico Mw7.2 earthquake on April 4, we carried out a number of detailed analysis of the TEC
165 variations in the region. Firstly, the analysis time was extended to more days before March 14 and after
166 April 6 to determine the TEC changes over a longer period of time. The results are given in Figures 4
167 and 5. We can see that, except the TEC decrease on April 11 when a magnetic storm occurred with a
168 minimum *Dst* value of -55 nT, there were no extremely high or low TEC values in the region for all the
169 days. Therefore, we can see from Figures 3-5 that, in 72 days from February 18 to April 30, there were
170 four days of TEC anomalies: TEC depletion on March 25 under geomagnetically quiet conditions, TEC
171 enhancements on April 5 and 6 under geomagnetically active conditions, and TEC depletion on April
172 11 under geomagnetic storm conditions. The unique occurrence of TEC depletion on the
173 geomagnetically quiet day of March 25 is thus potentially connected to the occurrence of the earthquake
174 on April 4 in the same region.



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176

Figure 4: Same as Figure 3, but for February 18 to March 13, 2010.



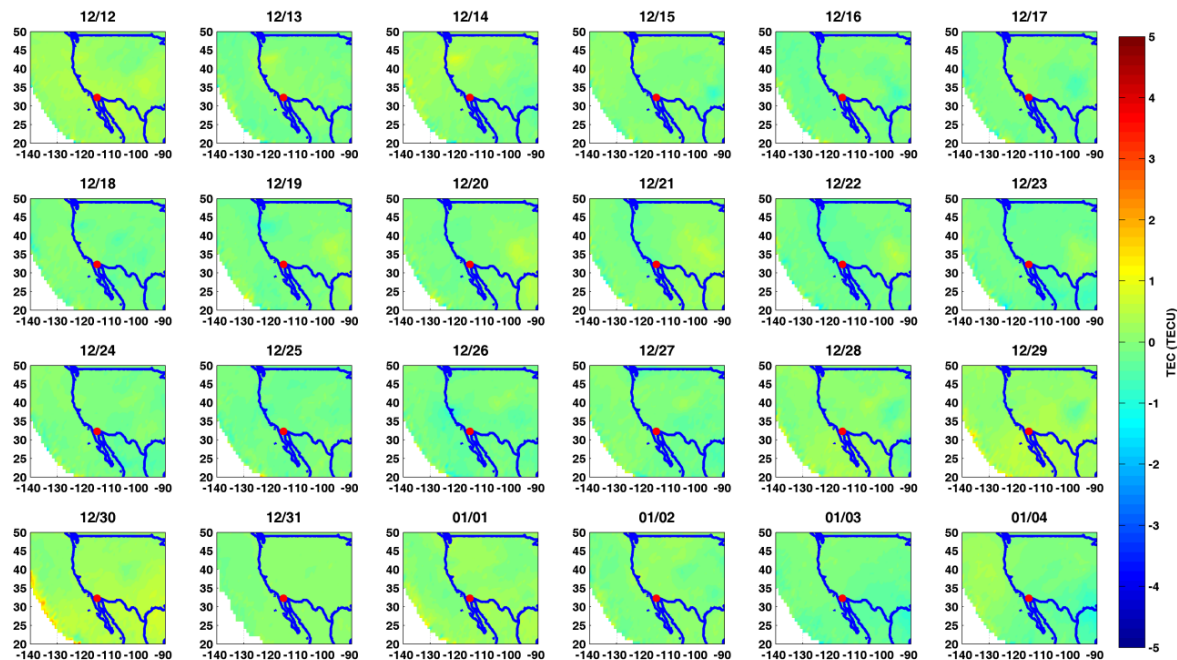
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Figure 5: Same as Figure 3, but for April 7 to April 30, 2010.



179 Secondly, we analyzed the TEC changes in the geomagnetically quiet days of other years ($-30 \text{ nT} <$
180 $Dst < 20 \text{ nT}$, $Kp < 3$, $AE < 500 \text{ nT}$) using the same analysis method described above. We found that the
181 time period from November 27 2009 to January 19 2010 was geomagnetically quiet and there were
182 continuous TEC data to allow a meaningful data analysis. In this study, since we use ± 15 days data as
183 the background, the first day with analysis results should be 15 days later than November 27 2009. The
184 distribution of the differences of TEC residuals from December 12 2009 to January 4 2010 is given in
185 Figure 6. There was no TEC anomaly in all those days, which means that when geomagnetic activity is
186 quiet, the TEC anomaly seen on March 25 before the Mexico Mw7.2 earthquake may not appear.



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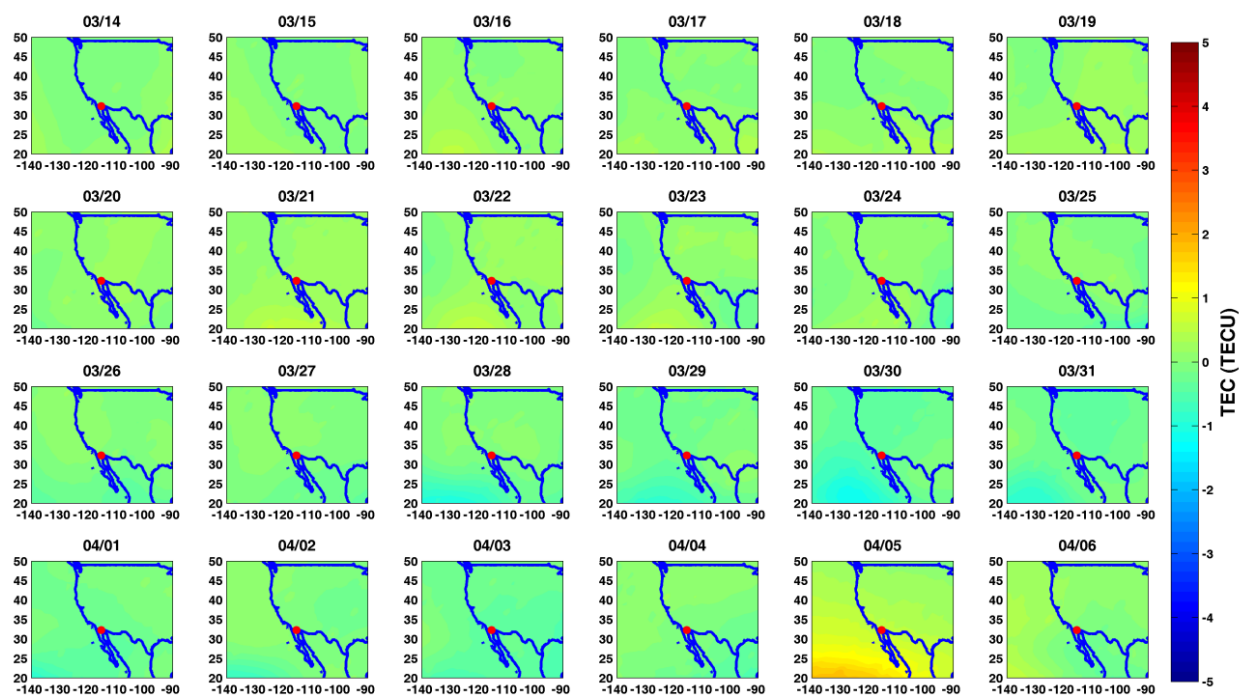
Figure 6: Same as Figure 3, but for December 12 2009 to January 4 2010 in geomagnetically quite days.



189 Thirdly, it is important to distinguish the seismic disturbance and the medium-scale traveling
190 ionospheric disturbances (MSTIDs). Iwata and Umeno (2017) calculated the rate of anomalous area and
191 the propagation velocity to detect the preseismic ionospheric disturbances. Otsuka et al. (2011) reported
192 that the propagation of atmospheric gravity waves and auroral activity are the main sources of the
193 MSTIDs. In this study, first principles simulations were employed to further examine the potential
194 source of the TEC anomaly seen on March 25, 2010, especially the lower atmospheric waves. Here we
195 used the thermosphere and ionosphere extension of the Whole Atmosphere Community Climate Model
196 (WACCM-X) (Liu et al., 2018). The top boundary of WACCM-X is set at 4.0×10^{-10} hPa (~500 to ~700
197 km altitude, depending on solar activity). WACCM-X can well simulate the chemical and physical
198 processes in the atmosphere (Marsh et al., 2013; Neale et al., 2013). WACCM-X can be configured
199 either for free climate simulations (lower atmosphere unconstrained), or to have the tropospheric and
200 stratospheric dynamics constrained to meteorological reanalysis fields for specifically targeted time
201 periods. The later one is called specified dynamics WACCM-X (SD-WACCM-X, Sassi et al., 2013). A
202 detailed description of the SD-WACCM-X can be found in Marsh (2011). With the lower atmospheric
203 dynamics constrained by the observational data, SD-WACCM-X can accurately represent the
204 large-scale and medium-scale lower atmospheric waves that can propagate upward and affect the
205 thermosphere and ionosphere. In this study, we used the SD-WACCM-X to determine whether the TEC
206 depletion seen on March 25 is related to lower atmospheric forcing. The SD-WACCM-X has a



207 horizontal resolution of 1.9° in latitude and 2.5° in longitude. Using the same data analysis method
208 described in section 2, the distributions of the differences of the simulated TEC residual from March 14
209 to April 6, 2010 are shown in Figure 7 within the region of latitude 20°N - 50°N and longitude
210 90°W - 140°W to be consistent with the data. Except the TEC enhancements around the northern crest of
211 EIA on April 5, no TEC anomaly is identified around the epicenter. The TEC depletion on March 25 is
212 not detected in SD-WACCM-X outputs, which means that the TEC anomaly source may not result from
213 the lower atmospheric forcing.



214

215

Figure 7: Same as Figure 3, but for SD-WACCM-X outputs from March 14 to April 6, 2010.



216 Takeuchi et al. (2006) pointed out that the stress of the p-holes in the crust can reach the surface
217 and create an upward electric field. Pulinets and Ouzounov (2011) supported the hypothesis of
218 atmospheric electricity changes caused by radon emanation and air ionization near the fault zone.
219 Sorokin et al. (2005) suggested that the DC electric field that forms above seismically active regions
220 could penetrate into the ionosphere. Kuo et al. (2011, 2014) proposed the electrical coupling between
221 the ionosphere and the surface charges in the earthquake fault zone. They suggested that the vertical
222 surface electric field drives currents in the atmosphere and electric fields at the bottom of the ionosphere.
223 If there is an upward electric field in the ionosphere, with the $\mathbf{E} \times \mathbf{B}$ drifts, the plasma moves westward.
224 Furthermore, Pulinets and Boyarchuk (2004a) suggested that the seismo-ionospheric phenomena did not
225 coincide with the vertical projection of the epicenter, but shifted equatorward. In our study, we found
226 that the depletion of TEC residuals occurred not only over the epicenter but also south and west to the
227 epicenter, which may be related to the above mentioned lithosphere-atmosphere-ionosphere coupling.

228 **5 Conclusions**

229 In this study, we applied a decomposition and nonlinear fitting method on the MIT TEC data to
230 obtain ionospheric TEC anomaly that is possibly associated with earthquakes. We analyzed the MIT
231 TEC data near the Mexico Mw7.2 earthquake that occurred on April 4 2010. We also carried out
232 numerical simulations using first principles model SD-WACCM-X. The main conclusions of this work
233 are as follows:



234 The TEC decreased on March 25 around the epicenter, 10 days before the earthquake. Except for
235 the TEC perturbations that were clearly related to geomagnetic activity, no TEC anomaly similar to that
236 on March 25 was seen in other 68 days around the day of the earthquake. Furthermore, the TEC
237 anomaly seen on March 25 cannot be found in geomagnetically quiet days from December 12 2009 to
238 January 4 2010 in the same region, either.

239 The TEC simulated by the SD-WACCM-X runs did not show TEC decrease around the epicenter
240 on March 25. SD-WACCM-X includes lower atmospheric large-scale waves and their coupling effects
241 on the ionosphere. Therefore, the model results suggest that the ionosphere TEC anomaly on March 25
242 might not be the result of lower atmospheric forcing. Our data analysis and model simulations thus
243 indicate that the TEC anomaly seen on March 25 may be potentially related to the Mexico Mw7.2
244 earthquake.

245 Although we identify ionospheric anomalies (TEC depletion) that are possibly associated with the
246 2010 Mexico Mw7.2 earthquake in this work, more case studies and physics-based simulations are
247 needed in the future to fully understand the physical mechanism of the seismo-ionospheric coupling.

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252 Observatory for providing the GPS TEC data, and the data can be downloaded from
253 <http://madrigal.haystack.mit.edu/madrigal/>. The authors acknowledge the WACCM-X development
254 teams at NCAR. The WACCM-X is an open-source community model. The outputs from model runs
255 used in this study are calculated using the NCAR supercomputer. NCAR is sponsored by the National
256 Science Foundation.

257 **Data availability**

258 The MIT TEC data can be downloaded from <http://madrigal.haystack.mit.edu/madrigal/>.

259 **Author contribution**

260 Jing Liu analyzed the data and wrote the manuscript. Wenbin Wang proposed the topic, conceived
261 and designed the study. Xuemin Zhang helped in the interpretation. All authors read and approved the
262 final manuscript.

263 **Competing interests**

264 The authors declare that they have no competing interest.

265 **Statement**

266 The journal of Annales Geophysicae has the right to reproduce materials in this manuscript,
267 including figures, tables, and maps.

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