

1 **Model of Propagation of VLF Beams in the Waveguide Earth-Ionosphere. Principles of Tensor Impedance Method in**
2 **Multilayered Gyrotropic Waveguides.**

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17 **Abstract.** The modeling of very low frequency electromagnetic beam propagation in the earth-ionosphere waveguide is
18 considered. A new tensor impedance method for modeling propagation of electromagnetic beams in a multi-
19 layered/inhomogeneous waveguide is presented. The waveguide is assumed to possess the gyrotropy and inhomogeneity
20 with a thick cover layer placed above the waveguide. The influence of geomagnetic field inclination and carrier beam
21 frequency on the characteristics of the polarization transformation in the Earth-ionosphere waveguide is determined. The
22 new method for modeling of propagation of electromagnetic beams allows us to study: (i) propagation of the very low
23 frequency modes in the earth-ionosphere waveguide and, in perspective, their excitation by the typical Earth-ionosphere
24 waveguide sources, such as radio wave transmitters and lightning discharges and (ii) leakage of Earth-ionosphere waveguide
25 waves into the upper ionosphere/magnetosphere. The proposed approach can be applied to the variety of problems related to
26 the analysis of propagation of electromagnetic waves in layered gyrotropic/anisotropic active media in a wide frequency
27 range, e.g. from the Earth-ionosphere waveguide to optics waveband, an artificial signal propagation such as metamaterial
28 microwave or optical waveguides.

29 **Keywords** — ionosphere, atmosphere, VLF, tensor impedance, gyrotropy, layered waveguide, beam, electromagnetic
30 waves, boundary conditions, ionospheric disturbances, vertical coupling processes

31

1 Introduction

32 The results of analytical and numerical study of very low frequency (VLF) electromagnetic (EM) wave/beam propagation in
33 the **Lithosphere–Atmosphere–Ionosphere–Magnetosphere system (LAIM)**, in particular in the waveguide “Earth–
34 Ionosphere” (WGEI) are presented. The amplitude and phase of the VLF wave propagates in the earth-ionosphere waveguide
35 (WGEI) can change, and these changes may be observable using ground-based and/or satellite detectors. This reflects the

36 variations in the ionospheric electrodynamic characteristics (complex dielectric permittivity) and the influences on the
37 ionosphere, for example, “from above” by chain Sun – Solar Wind – Magnetosphere – Ionosphere (Patra et al., 2011;
38 Koskinen, 2011; Boudjada et al., 2012; Wu et al., 2016; [Yiğit et al., 2016](#)). Then the influence on the ionosphere “from
39 below” comes from the most powerful meteorological, seismogenic and other sources in the lower atmosphere and
40 lithosphere/Earth, such as cyclones and hurricanes (Nina et al., 2017; Rozhnoi et al., 2014; Chou et al., 2015) as well as from
41 earthquakes (Hayakawa, 2015; Surkov and Hayakawa, 2014; Sanchez-Dulcet et al., 2015) and tsunamis. From inside the
42 ionosphere the strong thunderstorms, lightning discharges, and terrestrial gamma-ray flashes or sprite streamers (Cummer et
43 al., 1998; Qin et al., 2012; Dwyer, 2012; Dwyer and Uman, 2014; Cummer et al., 2014; Mezentsev et al., 2018) influence the
44 ionospheric electrodynamic characteristics as well. Note that the VLF signals are very important for the merging of the
45 atmospheric physics and space plasma physics with astrophysics and high-energy physics. The corresponding “intersection
46 area” for these four disciplines includes cosmic rays and very popular now objects of investigation – high-altitude discharges
47 (sprites), anomalous X-ray bursts, and powerful gamma-ray bursts. The key phenomena for the occurrence of all of these
48 objects is the appearance of runaway avalanche in the presence of high energy seed electrons. In the atmosphere, there are
49 cosmic ray secondary electrons (Gurevich and Zybin, 2001). Consequently, these phenomena are intensified during the air
50 shower generating by cosmic particles (Gurevich and Zybin, 2001; Gurevich et al., 2009). The runaway breakdown and
51 lightning discharges including **high-altitude** ones can **cause** radio emission both in HF range, which could be observed
52 using the Low-Frequency Array (LOFAR) radio telescope network facility and other radio telescopes (Buitink et al., 2014;
53 Scholten et al., 2017; Hare et al., 2018), and in the VLF range (Gurevich and Zybin, 2001). The corresponding experimental
54 research includes the measurements of the VLF characteristics by the international measurement system of the pairs
55 “transmitted-receiver” separated by a distance of a couple of thousand km (Biagi et al., 2011; 2015). The World Wide
56 Lightning Location Network is the one more international facility for VLF measurements during thunderstorms with
57 lightning discharges (Lu et al. 2019). Intensification of magnetospheric research, wave processes, particle distribution and
58 wave-particle interaction in the magnetosphere including radiation belts leads to the great interest to the VLF plasma waves,
59 in particular whistlers (Artemyev et al., 2013, [2015](#); Agapitov et al., 2014; Agapitov et al., 2018).

60 The differences of the proposed model for the simulation of VLF waves in the WGEI from others can be
61 summarized in three main points. (i) In distinction to the impedance invariant imbedding model (Shalashov and
62 Gospodchikov, 2010; Kim and Kim, 2016), our model provides an optimal balance between the analytical and numerical
63 approaches. It combines analytical-numerical approaches basing on matrix sweep method (Samarskii, 2001). As a result, this
64 model allows to obtain analytically the tensor impedance and, at the same time, provides high effectiveness and stability of
65 the modeling. (ii) In distinction to the full-wave finite difference time domain models (Chevalier and Inan, 2006; Marshall et
66 al., 2017; [Yu et al., 2012](#); Azadifar et al., 2017), our method provides the physically clear lower and upper boundary
67 conditions, in particular physically justified upper boundary conditions corresponding to the radiation of the waves
68 propagation in the WGEI to the upper ionosphere/magnetosphere. [This allows to determine the leakage modes and to](#)
69 [interpret not only ground-based, but also satellite measurements of the VLF beam characteristics.](#) (iii) In distinction to the

70 models considered in (Kuzichev and Shklyar, 2010; Kuzichev et al., 2018; Lehtinen and Inan, 2009; Lehtinen and Inan, 71 2008) based on the mode presentations and made in the frequency domain, we use the combined approach. This approach 72 includes the radiation condition at the altitudes of the F-region, equivalent impedance conditions in the lower E-region and at 73 the lower boundary of the WGEI, mode approach, and finally, the beam method. **This combined approach, finally, creates** 74 **the possibility to interpret adequately data of both ground-based and satellite detection on the VLF EM wave/beam** 75 **propagating in the WGEI and those, which experienced a leakage from the WGEI into the upper ionosphere and** 76 **magnetosphere. Some other details on the distinctions from the previously published models are given below in Sect. 3.**

77 The methods of effective boundary conditions such as effective impedance conditions (Tretyakov, 2003; Senior and 78 Volakis, 1995; Kurushin and Nefedov, 1983) are well-known and can be used, in particular, for the layered metal-dielectric, 79 metamaterial and gyrotropic active layered and waveguiding media of different types (Tretyakov, 2003; Senior and Volakis, 80 1995; Kurushin and Nefedov, 1983; Collin, 2001; Wait, 1996) including plasma-like solid state (Ruibys, and Tolutis, 1983) 81 and space plasma (Wait, 1996) media. The plasma wave processes in the waveguide structures metal-semiconductor- 82 dielectric, placed into the external magnetic field, were widely investigated (Ruibys and Tolutis, 1983; Maier, 2007; 83 Tarkhanyan and Uzunoglu, 2006) from radio to optical frequency ranges. Corresponding waves are applied in modern 84 plasmonics and in non-destructive testing of semiconductor interfaces. It is interesting to realize the resonant interactions of 85 volume and surface electromagnetic waves in these structures, so the simulations of the wave spectrum are important. To 86 describe such complex layered structures, it is very convenient and effective to use impedance approach (Tretyakov, 2003; 87 Senior and Volakis, 1995; Kurushin and Nefedov, 1983). As a rule, impedance boundary conditions are used, when the layer 88 covering waveguide is thin (Senior and Volakis, 1995; Kurushin and Nefedov, 1983). One of the known exclusions is the 89 impedance invariant imbedding model. The difference between our new method and that model is already mentioned above 90 and is explained in more details in the Subsection 3.3. Our new approach, i.e a new tensor impedance method for modeling 91 propagation of electromagnetic beams (TIMEB), includes a set of very attractive features for practical purposes. These 92 features are: (i) the surface impedance characterizes cover layer of finite thickness, and this impedance is expressed 93 analytically; (ii) the method allows an effective modelling of 3D beam propagating in the gyrotropic waveguiding structure; 94 (iii) finally, if the considered waveguide can be modified by any external influence such as bias magnetic or electric fields, 95 or by any extra wave or energy beams (such as acoustic or quasistatic fields etc.), the corresponding modification of the 96 characteristics (phase and amplitude) of the VLF beam propagating in the waveguide structure can be modelled.

97 Our approach was targeting properly and is suitable for the further development which will allow to solve also the 98 following problems: (i) the problem of the excitation of the waveguide by the waves incident on the considered structure 99 from above could be solved as well with the slight modification of the presented model, with inclusion also ingoing waves; 100 (ii) to consider a plasma-like system placed into the external magnetic field, such as the LAIM system (Grimalsky et al., 101 1999 a, b) or dielectric-magnetized semiconductor structure. The **electromagnetic waves** radiated outside the waveguiding 102 structure, such as helicons (Ruibys and Tolutis, 1983) or whistlers (Wait, 1996), and the waveguide modes could be 103 considered altogether. An adequate boundary radiation conditions on the upper boundary of the covering layer are derived.

104 Based on this and absence of ingoing waves, the leakage modes above the upper boundary of the structure (in other words,
105 upper boundary of covering layer), will be searched with the further development of the model delivered in the present
106 paper. Namely, it will be possible to investigate the process of the leakage of electromagnetic waves from the open
107 waveguide. Then their transformation into magnetized plasma waves, propagating along magnetic field lines, and the
108 following excitation of the waveguiding modes by the waves incident on the system from external space (Walker, 1976), can
109 be modeled as a whole. Combining with the proper measurements of the phases and amplitudes of the electromagnetic
110 waves, propagating in the waveguiding structures and leakage waves, the model can be used for searching, and even
111 monitoring the external influences on the layered gyrotropic active artificial or natural media, for example, microwave or
112 optical waveguides or the system LAIM and the earth-ionosphere waveguide, respectively.

113 An important effect of the gyrotropy and anisotropy is the corresponding transformation of the field polarization
114 during the propagation in the WGEI, absent in the ideal metal planar waveguide without gyrotropy and anisotropy. We will
115 search, how such an effect depends on the carrier frequency of the beam, propagating in the WGEI, inclination of the
116 geomagnetic field and perturbations in the electron concentration, which could vary under the influences of the powerful
117 enough sources placed “below”, “above” and/or “inside” the ionosphere.

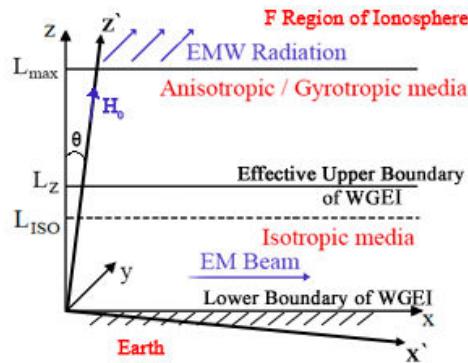
118 In Sect. 2 formulation of the problem is presented. In Sect. 3 the algorithm is discussed including the determination
119 of the VLF waves/beams radiation conditions into the upper ionosphere/magnetosphere at the upper boundary, placed in the
120 F-region at the altitude (250-400) km. The effective tensor impedance boundary conditions at the upper boundary (~ 85 km)
121 of the effective earth-ionosphere waveguide and the 3D model TIMEB of the propagation of the VLF beam in the WGEI are
122 discussed as well. The issues regarding the VLF beam leakage regimes are considered only very briefly, since the relevant
123 details will be presented in the following articles. In Sect. 4 the results of the numerical modeling are presented. In Sect. 5
124 the discussion is presented, including an example of the qualitative comparison between the results of our theory and an
125 experiment including the future rocket experiment on the measurements of the characteristics of VLF signal radiated from
126 the artificial VLF transmitter, which is propagating in the WGEI and penetrating into the upper ionosphere.

127 2 Formulation of the problem

128 The VLF electromagnetic (EM) waves with frequencies $f = (10 - 100)$ kHz can propagate along the Earth’s surface for long
129 distances > 1000 km. The Earth’s surface of a high conductivity $z = 0$ (z is vertical coordinate) and the ionosphere F -layer z
130 = 300 km form the VLF waveguide, see Fig. 1. The propagation of the VLF electromagnetic radiation excited by a near-
131 Earth antenna within the WGEI should be described by the full set of Maxwell’s equations in the isotropic atmosphere $0 < z$
132 < 60 km, the approximately isotropic ionosphere D -layer $60 \text{ km} < z < 75$ km, and the anisotropic E - and F - layers of the
133 ionosphere, due to the geomagnetic field \vec{H}_0 , added by the boundary conditions at the Earth’s surface and at the F -layer. In
134 Fig. 1, θ is the angle between the directions of the vertical axis z and geomagnetic field \vec{H}_0 . Note that theta θ angle is
135

137 complementary to the angle of inclination of the geomagnetic field. Geomagnetic field \vec{H}_0 is directed along z' axis, lies in
 138 the plane xz , while the planes $x'z'$ and xz coincide with each other.

139



140

141 **Figure 1. The geometry of the anisotropic/gyroscopic waveguide. EM waves propagate in OX direction.** \vec{H}_0 is the external
 142 magnetic field. The effective WGEI for EM waves occupies the region $0 < z < L_z$. The isotropic medium occupies the region $0 < z <$
 143 L_{ISO} , $L_{ISO} < L_z$. The anisotropic/gyroscopic medium occupies the region $L_{ISO} < z < L_{max}$. Covering layer occupies the region $L_z < z <$
 144 L_{max} . WG includes isotropic region $0 < z < L_{ISO}$ and a part of anisotropic region $L_z < z < L_{max}$. It is supposed that the anisotropic
 145 region is relatively small part of the WG, $(L_z - L_{ISO})/L_z \sim (0.1-0.2)$. At the upper boundary of covering layer ($z = L_{max}$) the radiation
 146 of EM to the external region ($z > L_{max}$) is accounted for with the proper boundary conditions. Integration of the equations
 147 describing the EM field propagation allows to obtain effective impedance boundary conditions at the upper boundary of effective
 148 WG ($z = L_z$). These boundary conditions effectively includes all the effects on the wave propagation of the covering layer and the
 149 radiation (at $z = L_{max}$) to the external region ($z > L_{max}$).

150

151 3. Algorithm

152 The boundary conditions and calculations of impedance and beam propagation in the WGEI is considered in this Section. The
 153 other parts of the algorithm, e.g. the reflection of the EM waves from the WGEI effective upper boundary and the leakage of
 154 EM waves from the WGEI to the upper ionosphere/magnetosphere, will be presented very briefly as they are the subjects of
 155 the next papers.

156 3.1 Direct and inverse tensors characterizing the ionosphere

157 In the next subsections we derived the formulas describing the transfer of the boundary conditions at the upper boundary ($z =$
 158 L_{max}), Fig. 1, resulting in the tensor impedance conditions at the upper boundary of the effective WGEI ($z = L_i$). Firstly let's
 159 describe the tensors characterizing the ionosphere.

160 The algorithm's main goal is to transfer the EM boundary conditions from the upper ionosphere at the height $L_z \sim 250 - 400$
 161 km to the lower ionosphere $L_z \sim 70 - 90$ km. All components of the monochromatic EM field is considered to be
 162 proportional to $\exp(i\omega t)$. The anisotropic medium is inhomogeneous along OZ axis only and characterized by the
 163 permittivity tensor $\hat{\epsilon}(\omega, z)$ or by the inverse tensor $\hat{\beta}(\omega, z) = \hat{\epsilon}^{-1}(\omega, z)$: $\vec{E} = \hat{\beta}(\omega, z) \cdot \vec{D}$, where \vec{D} is the electric induction.

164 Below the absolute units are used. The expressions for the components of the effective permittivity of the ionosphere are in
 165 the coordinate frame $X'YZ'$ where OZ' axis is aligned along the geomagnetic field \vec{H}_0 :

166

$$\begin{aligned}
 \hat{\varepsilon}' &= \begin{pmatrix} \varepsilon_1 & \varepsilon_h & 0 \\ -\varepsilon_h & \varepsilon_1 & 0 \\ 0 & 0 & \varepsilon_3 \end{pmatrix}, \quad \varepsilon_1 = 1 - \frac{\omega_{pe}^2 \cdot (\omega - i\nu_e)}{((\omega - i\nu_e)^2 - \omega_{He}^2) \cdot \omega} - \frac{\omega_{pi}^2 \cdot (\omega - i\nu_i)}{((\omega - i\nu_i)^2 - \omega_{Hi}^2) \cdot \omega}, \quad \varepsilon_h \equiv ig; \\
 167 \quad g &= -\frac{\omega_{pe}^2 \cdot \omega_{He}}{((\omega - i\nu_e)^2 - \omega_{He}^2) \cdot \omega} + \frac{\omega_{pi}^2 \cdot \omega_{Hi}}{((\omega - i\nu_i)^2 - \omega_{Hi}^2) \cdot \omega}, \quad \varepsilon_3 = 1 - \frac{\omega_{pe}^2}{(\omega - i\nu_e) \cdot \omega} - \frac{\omega_{pi}^2}{(\omega - i\nu_i) \cdot \omega}; \\
 \omega_{pe}^2 &= \frac{4\pi e^2 n}{m_e}, \quad \omega_{pi}^2 = \frac{4\pi e^2 n}{m_i}, \quad \omega_{He} = \frac{eH_0}{m_e c}, \quad \omega_{Hi} = \frac{eH_0}{m_i c}
 \end{aligned} \tag{1}$$

168

169 Here $\omega_{pe}, \omega_{pi}, \omega_{He}, \omega_{Hi}$ are the plasma and cyclotron frequencies for electrons and ions, respectively; m_e, m_i, ν_e, ν_i are the
 170 masses and collision frequencies for electrons and ions, respectively, n is electron concentration. **The approximations of**
 171 **the three-component plasma-like ionosphere (including electron, one effective ion and one effective neutral**
 172 **components) and quasi-neutrality are accepted.** The expressions for the components of permittivity tensor $\hat{\varepsilon}(\omega, z)$ are
 173 obtained from (1) by means of multiplication with the standard rotation matrices (Spiegel, 1959) dependent on angle θ . For
 174 the medium with a scalar conductivity σ , e.g. lower ionosphere or atmosphere, the effective permittivity in (1) reduces to $\varepsilon =$
 175 $1 - 4\pi i \sigma / \omega$.

176 **3.2 The equations for the EM field and upper boundary conditions**

177 The case of the VLF waveguide modes k_x is slightly complex and should be calculated accounting for boundary conditions at
 178 the Earth's surface and upper surface of the effective WGEI. The EM field depends on the horizontal coordinate x as
 179 $\exp(-k_x x)$. Taking into account $k_x \leq k_0$ ($k_0 = \omega/c$), in simulations of the VLF beam propagation, we assume $k_x = k_0$.

180 Therefore, Maxwell's equations are:

$$\begin{aligned}
 -\frac{\partial H_y}{\partial z} &= ik_0 D_x, \quad \frac{\partial H_x}{\partial z} + ik_x H_z = ik_0 D_y, \quad -ik_x H_y = ik_0 D_z \\
 181 \quad -\frac{\partial E_y}{\partial z} &= -ik_0 H_x, \quad \frac{\partial E_x}{\partial z} + ik_x E_z = -ik_0 H_y, \quad -ik_x E_y = -ik_0 H_z
 \end{aligned} \tag{2}$$

182 Here, $E_x = \beta_{11} D_x + \beta_{12} D_y + \beta_{13} D_z$ etc. All components of the EM field can be represented through the horizontal components
 183 of the magnetic field H_x and H_y . The equations for these components are:

184

$$\frac{\partial}{\partial z} \left(\frac{\beta_{22}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \frac{\partial H_x}{\partial z} \right) - \frac{\partial}{\partial z} \left(\frac{\beta_{21}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \frac{\partial H_y}{\partial z} \right) - ik_x \frac{\partial}{\partial z} \left(\frac{\beta_{23}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} H_y \right) + k_0^2 H_x = 0 \quad (3a)$$

185

$$\begin{aligned} & \frac{\partial}{\partial z} \left(\left(\beta_{11} + \frac{k_x^2}{k_0^2} \frac{\beta_{12} \cdot \beta_{21}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \right) \frac{\partial H_y}{\partial z} \right) - \frac{\partial}{\partial z} \left(\frac{\beta_{12}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \frac{\partial H_x}{\partial z} \right) + \\ & + ik_x \frac{\partial}{\partial z} \left(\left(\beta_{13} + \frac{k_x^2}{k_0^2} \frac{\beta_{12} \cdot \beta_{23}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \right) H_y \right) + ik_x \left(\beta_{31} + \frac{k_x^2}{k_0^2} \frac{\beta_{32} \cdot \beta_{21}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \right) \frac{\partial H_y}{\partial z} - \\ & - ik_x \frac{\beta_{32}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \frac{\partial H_x}{\partial z} + k_0^2 \left(1 - \beta_{33} \frac{k_x^2}{k_0^2} - \frac{k_x^4}{k_0^4} \frac{\beta_{23} \cdot \beta_{32}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \right) H_y = 0 \end{aligned} \quad (3b)$$

186 The expressions for the horizontal components of the electric field E_x, E_y are:

187

$$E_x = \frac{i}{k_0} \left(\left(\beta_{11} + \frac{k_x^2}{k_0^2} \frac{\beta_{12} \cdot \beta_{21}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \right) \frac{\partial H_y}{\partial z} - \frac{\beta_{12}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \frac{\partial H_x}{\partial z} \right) - \frac{k_x}{k_0} \left(\beta_{13} + \frac{k_x^2}{k_0^2} \frac{\beta_{12} \cdot \beta_{23}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \right) H_y$$

188

$$E_y = \frac{i}{k_0} \left(- \frac{\beta_{22}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \frac{\partial H_x}{\partial z} + \frac{\beta_{21}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \frac{\partial H_y}{\partial z} \right) - \frac{k_x}{k_0} \frac{\beta_{23}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} H_y \quad (4)$$

189

In the region $z \geq L_{max}$ the upper ionosphere is assumed to be weakly inhomogeneous, and the geometric optics approximation is valid in the VLF range there. We should note that due to high inhomogeneity of the ionosphere in the vertical direction within E -layer (i.e. at the upper boundary of the effective VLF WGEI) such an approximation is not applicable. These conditions determine the choice of the upper boundary $z = L_{max} \sim (250 - 400)$ km, where the conditions of the radiation are formulated. The dispersion equation connected the wave numbers and the frequency of the outgoing waves is obtained from Eqs. (3), where $H_{x,y} \sim \exp(-ik_z z)$, while the derivatives like $\partial\beta_{11}/\partial z$ and the inhomogeneity of the media are neglected:

$$\begin{aligned}
& \left(\beta_{22} k_z^2 - k_0^2 (1 - \beta_{22} \frac{k_x^2}{k_0^2}) \right) \cdot \left((\beta_{11} (1 - \beta_{22} \frac{k_x^2}{k_0^2}) + \frac{k_x^2}{k_0^2} \beta_{12} \cdot \beta_{21}) k_z^2 + ((\beta_{13} + \beta_{31}) (1 - \beta_{22} \frac{k_x^2}{k_0^2}) + \right. \\
& \left. + \frac{k_x^2}{k_0^2} (\beta_{12} \cdot \beta_{23} + \beta_{32} \cdot \beta_{21}) k_x k_z - k_0^2 ((1 - \beta_{33} \frac{k_x^2}{k_0^2}) (1 - \beta_{22} \frac{k_x^2}{k_0^2}) - \frac{k_x^4}{k_0^4} \beta_{23} \cdot \beta_{32}) \right) - \\
& - (\beta_{21} k_z^2 + \beta_{23} k_x k_z) \cdot (\beta_{12} k_z^2 - \beta_{32} k_x k_z) = 0
\end{aligned} \tag{5}$$

196 Thus, generally Eq. (5) determined the wave numbers for the outgoing waves is of the fourth order (Wait 1996). The
197 boundary conditions at the upper boundary $z = L_{max}$ within the ionosphere F -layer are the absence of the ingoing waves, i.e.
198 the outgoing radiated (leakage) waves are present only. Two roots should be selected that possess the negative imaginary
199 parts $Im(k_{z1, z2}) < 0$, i.e. the outgoing waves dissipate upwards. However, in the case of VLF waves, some simplification can
200 be used. Namely, the expressions for the wave numbers $k_{1,2}$ are obtained from Eqs. (3), where the dependence on x is
201 neglected: $|k_{1,2}| \gg k_0$. This approximation is valid within F -layer where the first outgoing wave corresponds to the whistlers
202 with small dissipation, the second one is the highly dissipating slow wave. To formulate the boundary conditions for Eqs.
203 (3a, b) at $z \geq L_{max}$, the EM field components can be presented as:

$$H_x = A_1 e^{-ik_{z1}\tilde{z}} + \alpha_2 A_2 e^{-ik_{z2}\tilde{z}}, \quad H_y = \alpha_1 A_1 e^{-ik_{z1}\tilde{z}} + A_2 e^{-ik_{z2}\tilde{z}} \tag{6}$$

204 In the relations (6) $\tilde{z} = z - L_z$. Eqs. (3) are simplified in the approximation described above:

$$\beta_{22} \frac{\partial^2 H_x}{\partial z^2} - \beta_{21} \frac{\partial^2 H_y}{\partial z^2} + k_0^2 H_x = 0, \quad \beta_{11} \frac{\partial^2 H_y}{\partial z^2} - \beta_{12} \frac{\partial^2 H_x}{\partial z^2} + k_0^2 H_y = 0 \tag{7}$$

205 The solution of Eqs. (7) is searched for as: $H_{x,y} \sim \exp(-ik_z \tilde{z})$. The following equation has been obtained to get the wave
206 numbers $k_{z1, z2}$ from Eqs. (7):

$$\kappa^4 - (\beta_{22} + \beta_{11})\kappa^2 + \beta_{11}\beta_{22} - \beta_{12}\beta_{21} = 0, \quad \kappa^2 = \frac{k_0^2}{k_z^2} \tag{8}$$

212 Therefore, from Eq. (8) follows,

$$\kappa_{1,2}^2 = \frac{\beta_{11} + \beta_{22}}{2} \pm \left(\left(\frac{\beta_{11} - \beta_{22}}{2} \right)^2 + \beta_{12}\beta_{21} \right)^{1/2}; \quad \alpha_1 = \frac{\beta_{22} - \kappa_1^2}{\beta_{21}} = \frac{\beta_{12}}{\beta_{11} - \kappa_1^2}; \quad \alpha_2 = \frac{\beta_{11} - \kappa_2^2}{\beta_{12}} = \frac{\beta_{21}}{\beta_{22} - \kappa_2^2}; \quad k_{z1, z2}^2 = \frac{k_0^2}{\kappa_{1,2}^2} \tag{9}$$

214 The signs of $k_{z1, z2}$ have been chosen from the condition $Im(k_{z1, z2}) < 0$. From Eq. (5) at the upper boundary $z = L_{max}$, the
215 following relations are valid:

$$H_x = A_1 + \alpha_2 A_2, \quad H_y = \alpha_1 A_1 + A_2 \tag{10}$$

218

219 From Eq. (10) one can get

220 $A_1 = \Delta^{-1}(H_x - \alpha_2 H_y); A_2 = \Delta^{-1}(H_y - \alpha_1 H_x); \Delta = 1 - \alpha_1 \alpha_2 \quad (11)$

221 Thus, it is possible to exclude the amplitudes of the outgoing waves $A_{1,2}$ from Eqs. (9). As a result, at $z = L_{max}$ the boundary
222 conditions are rewritten in terms of H_x, H_y only:

223
$$\begin{aligned} \frac{\partial H_x}{\partial z} &= -i(k_{z1}A_1 + k_{z2}\alpha_2 A_2) = -\frac{i}{\Delta} \left((k_{1z} - \alpha_1 \alpha_2 k_{z2}) H_x + \alpha_2 (k_{z2} - k_{z1}) H_y \right) \\ \frac{\partial H_y}{\partial z} &= -i(k_{z1}\alpha_1 A_1 + k_{z2}A_2) = -\frac{i}{\Delta} \left((k_{z2} - \alpha_1 \alpha_2 k_{z1}) H_y + \alpha_1 (k_{z1} - k_{z2}) H_x \right) \end{aligned} \quad (12)$$

224 The relations (12) are the upper boundary conditions of the radiation for the boundary $z=L_{max} \sim 250 - 400$ km. These
225 conditions will be transformed/recalculated using the analytical numerical recurrent procedure into equivalent impedance
226 boundary conditions at $z = L_z \sim 70 - 90$ km.

227 Note that in the “whistler/VLF approximation” is valid at frequencies ~ 10 kHz for the F -region of the ionosphere.
228 In this approximation and $k_x \approx 0$, we receive the dispersion equation using Eqs. (5), (8), (9), in the form:

229 $k_z'^2 k^2 = k_0^2 g^2 \quad (13)$

230 where $k^2 = k_x^2 + k_z^2 = k_x'^2 + k_z'^2$; k_x' and k_z' are the transverse and longitudinal components of wave number relative to
231 **geomagnetic field**. For F -region of the ionosphere, where $v_e \ll \omega \ll \omega_{He}$, Eq. (13) reduces to the standard form of the
232 whistler dispersion equation $|k_z'| k = k_0 |g|$; $g \approx -\omega_{pe}^2 / (\omega \omega_{He})$; $\omega = c^2 k |k_z'| (\omega_{He} / \omega_{pe}^2)$. In a special case of the waves
233 propagating along geomagnetic field, $k_x' = 0$, for the propagating whistler waves, well-known dispersion dependence is
234 $\omega = c^2 k_z'^2 (\omega_{He} / \omega_{pe}^2)$ (Artcimovich and Sagdeev, 1979). For the formulated problem we can reasonably assume $k_x \approx 0$.
235 Therefore eq. (13) is reduced to $k_z^4 \cos^2 \theta = k_0^4 g^2$. As a result, we get $k_{z1} = \sqrt{g / \cos \theta} k_0$, $k_{z2} = -i \sqrt{g / \cos \theta} k_0$, and then,
236 similarly to the relations (12), the boundary conditions can be presented, in terms of the tangential components of electric
237 field as:

238
$$\frac{\partial \vec{U}}{\partial z} + \hat{B} \vec{U} = 0; \vec{U} = \begin{bmatrix} E_x \\ E_y \end{bmatrix}; \hat{B} = \frac{1}{2} \sqrt{\frac{g}{\cos \theta}} k_0 \begin{bmatrix} 1+i & -1-i \\ 1+i & 1+i \end{bmatrix} \quad (14)$$

239 Conditions (12) or (14) are the conditions of radiation (absence of ingoing waves) formulated at the upper boundary $z=L_{max}$
240 and suitable for the determination of the energy of the wave leaking from the WGEI into the upper
241 ionosphere/magnetosphere. Note, the equations (12), (14) expressed the boundary conditions of the radiation (more
242 accurately speaking, an absence of incoming waves, what is the consequence to the causality principle), are obtained as a
243 result of limitation by the small parameter k_x/k_0 $|k_x|/k_z| \rightarrow 0$ in Eq. (5). In spite of the disappearance of the dependence of
244 these boundary conditions explicitly on k_x , the dependence of the characteristics of the wave propagation process on k_x , as a
245 whole, is accounted for, and all results are still valid for the description of the wave beam propagation in the WGEI along the
246 horizontal axis x with finite $k_x \sim k_0$.

248 **3.3 Equivalent Tensor Impedance Boundary Conditions**

249 The tensor impedance at the upper boundary of the effective WGEI $z=L_z$ (see Fig. 1), is obtained by the conditions
 250 of radiation (12) or (14), recalculated from the level $z = L_z \sim 80 - 90$ km, placed in the F region of the ionosphere, at $z=L_{max} \sim$
 251 $250 - 400$ km.

252 The main idea of the effective tensor impedance method is the unification of analytical and numerical approaches
 253 and derivation of the proper impedance boundary conditions without “thin cover layer” approximation. This approximation
 254 is usually used in the effective impedance approaches, applied either for artificial or natural layered gyrotropic structures,
 255 (see e.g. Tretyakov, 2003; Senior and Volakis, 1995; Kurushin and Nefedov, 1983; Alperovich and Fedorov, 2007). There is
 256 one known exception, namely invariant imbedding impedance method (Shalashov and Gospodchikov, 2011; Kim and Kim,
 257 2016). The comparison of our method with the invariant imbedding impedance method will be presented at the end of this
 258 subsection. Eqs. (3) jointly with the boundary conditions (12), have been solved by finite differences. The derivatives in Eqs.
 259 (3) are approximated as

260

$$\frac{\partial}{\partial z} \left(C(z) \frac{\partial H_x}{\partial z} \right) \approx \frac{1}{h} \left(C(z_{j+1/2}) \frac{(H_x)_{j+1} - (H_x)_j}{h} - C(z_{j-1/2}) \frac{(H_x)_j - (H_x)_{j-1}}{h} \right),$$

$$\frac{\partial}{\partial z} (F(z) H_x) \approx \frac{1}{2h} (F(z_{j+1}) (H_x)_{j+1} - F(z_{j-1}) (H_x)_{j-1}) \quad \text{etc.} \quad (15)$$

261 In Eq. (15) $z_{j+1/2} = h \cdot (j + 0.5)$. In Eqs. (10) the approximation is $\partial H_x / \partial z \approx [(H_x)_N - (H_x)_{N-1}] / h$. Here h is the
 262 discretization step along OZ axis; N is the total number of nodes. At each step j the difference approximations of Eqs. (3)
 263 take the form:

265
$$\hat{\alpha}_j^{(-)} \cdot \vec{H}_{j-1} + \hat{\alpha}_j^{(0)} \cdot \vec{H}_j + \hat{\alpha}_j^{(+)} \cdot \vec{H}_{j+1} = 0 \quad (16)$$

266 where $\vec{H}_j = \begin{pmatrix} H_x \\ H_y \end{pmatrix}$, $j = N-1, N-2, \dots, 1$, $z_j = h \cdot j$, $L_z = h \cdot N$. Due to the complexity of expressions for the matrix
 267 coefficients in Eq. (16) we have shown them in Appendix. The set of the matrix Eq. (16) has been solved by factorization
 268 method also known as an elimination/matrix sweep method (see Samarskii, 2001). It can be written as:

269
$$\vec{H}_j = \hat{b}_j \cdot \vec{H}_{j-1}, \quad j = N, \dots, 1 \quad (17a)$$

270
$$H_{xj+1} = b_{11j+1} H_1 + b_{12j+1} H_2; \quad H_{yj+1} = b_{21j+1} H_1 + b_{22j+1} H_2; \quad H_1 \equiv H_{xj}; \quad H_2 \equiv H_{yj} \quad (17b)$$

271 This method is a variant of the Gauss elimination method for the matrix 3-diagonal set of the Eq. (16). The value of \hat{b}_N is
 272 obtained from the boundary conditions (12) as:

273
$$\hat{\alpha}_N^{(-)} \cdot \vec{H}_{N-1} + \hat{\alpha}_N^{(0)} \cdot \vec{H}_N = 0 \quad (18)$$

274 Therefore $\hat{b}_N = -(\hat{\alpha}_N^{(0)})^{-1} \cdot \hat{\alpha}_N^{(-)}$. Then the matrices \hat{b}_j have been computed sequentially down to the desired value of $z =$
 275 $L_z = h \cdot N_z$, where the impedance boundary conditions are assumed to be applied. At each step the expression for \hat{b}_j follow
 276 from (16), (17) as:

$$277 \quad (\hat{\alpha}_j^{(0)} + \hat{\alpha}_j^{(+)} \cdot \hat{b}_{j+1}) \cdot \vec{H}_j = -\hat{\alpha}_j^{(-)} \cdot \vec{H}_{j-1} \quad (19)$$

278 Therefore, for (17), we obtain $\hat{b}_j = -(\hat{\alpha}_j^{(0)} + \hat{\alpha}_j^{(+)} \cdot \hat{b}_{j+1})^{-1} \cdot \hat{\alpha}_j^{(-)}$. The derivatives in Eqs. (4) have been approximated as:

$$279 \quad \left(\frac{\partial H_x}{\partial z} \right)_{N_z} \approx \frac{(H_x)_{N_z+1} - (H_x)_{N_z}}{h} = \frac{(b_{N_z+1-11} - 1) \cdot (H_x)_{N_z} + b_{N_z+1-12} \cdot (H_y)_{N_z}}{h} \quad (20)$$

280 and similar equation can be obtained for $\left(\frac{\partial H_y}{\partial z} \right)_{N_z}$. Note, that as a result of this discretization, only the values at the grid
 281 level N_z are included into the numerical approximation of the derivatives $\partial H_{x,y} / \partial z$ at $z = L_z$. We determine tensor
 282 impedance \hat{Z} at $z = L_z \sim 85$ km level. The tensor values are included into the following relations, all of which are
 283 corresponded to altitude (in other words, to the grid with number N_z , corresponded to this altitude):

$$284 \quad \vec{n} \times \vec{E} = \hat{Z}_0 \cdot \vec{H}, \quad \vec{n} = (0, 0, 1) \text{ or } E_x = Z_{021} H_x + Z_{022} H_y; \quad E_y = -Z_{011} H_x - Z_{012} H_y \quad (21)$$

285 The equivalent tensor impedance is obtained using a two-step procedure. (1) We obtain the matrix \hat{b}_j using Eqs. (3a, b) with
 286 the boundary conditions (12) and the procedure (17) – (19) described above. (2) Placing the expressions (21) with tensor
 287 impedance into the left parts and the derivatives $\partial H_{x,y} / \partial z$ in the form (20) into the right parts of Eqs. (4), the analytical
 288 expressions for the components of the tensor impedance are:

$$289 \quad Z_{011} = -\frac{i}{k_0 h} \begin{pmatrix} \frac{\beta_{21}}{k_x^2} \cdot b_{21} - \frac{\beta_{22}}{k_x^2} \cdot (b_{11} - 1) \\ 1 - \beta_{22} \frac{k_x^2}{k_0^2} & 1 - \beta_{22} \frac{k_x^2}{k_0^2} \end{pmatrix},$$

$$290 \quad Z_{012} = -\frac{i}{k_0 h} \begin{pmatrix} \frac{\beta_{21}}{k_x^2} \frac{\partial H_y}{\partial z} \cdot (b_{22} - 1) - \frac{\beta_{22}}{k_x^2} \cdot b_{12} - k_x h \cdot \frac{\beta_{23}}{k_0^2} \\ 1 - \beta_{22} \frac{k_x^2}{k_0^2} & 1 - \beta_{22} \frac{k_x^2}{k_0^2} \end{pmatrix},$$

$$291 \quad Z_{021} = \frac{i}{k_0 h} \begin{pmatrix} (\beta_{11} + \frac{k_x^2}{k_0^2} \frac{\beta_{12} \cdot \beta_{21}}{k_x^2}) \cdot b_{21} - \frac{\beta_{12}}{k_x^2} \cdot (b_{11} - 1) \\ 1 - \beta_{22} \frac{k_x^2}{k_0^2} & 1 - \beta_{22} \frac{k_x^2}{k_0^2} \end{pmatrix},$$

292

$$Z_{022} = \frac{i}{k_0 h} \left(\left(\beta_{11} + \frac{k_x^2}{k_0^2} \frac{\beta_{12} \cdot \beta_{21}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \right) \cdot (b_{22} - 1) - k_x h \cdot \left(\beta_{13} + \frac{k_x^2}{k_0^2} \frac{\beta_{12} \cdot \beta_{23}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \right) - \frac{\beta_{12}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \cdot b_{12} \right) \quad (22)$$

293

294 The proposed method of the transfer of the boundary conditions from the ionosphere F -layer $L_{max} = 250 - 400$ km into the
 295 lower part of the E -layer $L_z = 80 - 90$ km is stable and easily realizable in comparison with some alternative approaches
 296 based on the invariant imbedding methods (Shalashov and Gospodchikov, 2011; Kim and Kim, 2016). The stability of our
 297 method is due to the stability of the Gauss elimination method when the coefficients at the matrix central diagonal are
 298 dominating. The last is valid for the ionosphere with electromagnetic losses where the absolute values of the permittivity
 299 tensor are large. The application of the proposed matrix sweep method in the media without **losses may** require the use of the
 300 Gauss method with the choice of the maximum element, to ensure the stability. However, as our simulations (not presented
 301 here) demonstrated, for the electromagnetic problems in the frequency domain, the simple Gauss elimination and the choice
 302 of the maximal **element give** the same results. The accumulation of errors may occur in evolutionary problems in the time
 303 domain, when the Gauss method should be applied sequentially many times. The use of the independent functions H_x, H_y in
 304 Eqs. (3) seems natural, as well as the transfer (17a), because the impedance conditions are the expressions of the electric E_x ,
 305 E_y through these magnetic components H_x, H_y at the upper boundary of the VLF waveguide $80 - 90$ km. The naturally
 306 chosen direction of the recalculation of the upper boundary conditions from $z = L_{max}$ to $z = L_z$, i.e. from upper layer with
 307 large impedance value to lower altitude layer with relatively small impedance value, provides, at the same time, the stability
 308 of the simulation procedure. The obtained components of the tensor impedance \hat{Z}_0 are small, $|Z_{0\alpha\beta}| \leq 0.1$. This determines
 309 the choice of the upper boundary $z = L_z$ for the effective WGEI. Due to small impedance, EM waves incident from below on
 310 this boundary are reflected effectively back. Therefore, the region $0 \leq z \leq L_z$ indeed can be presented as an effective WGEI.
 311 This waveguide includes not only lower boundary at $L_{ISO} \sim 65 - 75$ km with rather small losses, but also thin dissipative and
 312 anisotropic/gyrotropic layer between 75 and 85 - 90 km.

313 Finally, the main differences and advantages of the proposed tensor impedance method from other methods for
 314 impedance recalculating and in particular invariant imbedding methods (Shalashov and Gospodchikov, 2011; Kim and Kim,
 315 2016) can be summarised as follows:

316 (i) in contrast to invariant imbedding method currently proposed method can be used for direct recalculation of
 317 tensor impedance as it determined analytically, see Eqs. (22).

318 (ii) for the media without non-locality, proposed method does not require to solve integral equation(s).

319 (iii) the proposed method does not require forward and reflected waves. The conditions for the radiation at the upper
 320 boundary $z = L_{max}$ (see Eqs. (12)) are determined through the total field components $H_{x,y}$, which simplify the overall
 321 calculations.

322 (iv) the overall calculation procedure is very effective and computationally stable. Note, that even for the very low-

323 loss systems, the required level of stability can be achieved with modification based on the choice of the maximal element

324 for matrix inversion.

325 3.4 Propagation of Electromagnetic Waves in the Gyrotropic Waveguide 326 and the TIMEB Method

327 Let's use the transverse components of electric E_y and magnetic H_y fields to derive equations for the slow varying
328 amplitudes $A(x,y,z)$, $B(x,y,z)$ of the VLF beams. These components can be represented as:

$$329 \quad E_y = \frac{1}{2} A(x, y, z) \cdot e^{i\omega t - ik_0 x} + c.c., \quad H_y = \frac{1}{2} B(x, y, z) \cdot e^{i\omega t - ik_0 x} + c.c. \quad (23)$$

330 Here we assumed $k_x = k_0$ to reflect beam propagation in the WGEI with the main part in the atmosphere and lower
331 ionosphere (*D*-region) which are similar to free space by its electromagnetic parameters. The presence of a thin anisotropic
332 and dissipative layer belonging to the *E*-region (Guglielmi and Pokhotelov, 1996) of the ionosphere causes, altogether with
333 the impedance boundary condition, the proper z dependence of $B(x,y,z)$. Using (21) and (22), the boundary conditions are
334 determined at the height $z = L_z$ for the slowly varying amplitudes $A(x,y,z)$, $B(x,y,z)$ of the transverse components E_y , H_y . **As it**
335 **follows from Maxwell's equations**, the components E_x and H_x through E_y , H_y in the method of **beams have the form**:

$$336 \quad H_x \approx -\frac{i}{k_0} \frac{\partial E_y}{\partial z}, \quad E_x \approx \gamma_{12} E_y + i \frac{\tilde{\beta}_{33}}{k_0} \frac{\partial H_y}{\partial z} + \tilde{\beta}_{13} H_y \quad (24)$$

337 where $\gamma_{12} = \Delta_0^{-1} (\varepsilon_{13} \varepsilon_{32} - \varepsilon_{12} \varepsilon_{33})$, $\tilde{\beta}_{13} = \Delta_0^{-1} \varepsilon_{13}$, $\tilde{\beta}_{33} = \Delta_0^{-1} \varepsilon_{33}$; $\Delta_0 = \varepsilon_{11} \varepsilon_{33} - \varepsilon_{13} \varepsilon_{31}$. From Eqs. (21) and (24), the boundary
338 conditions for A , B can be defined as:

$$339 \quad A - \frac{i}{k_0} Z_{11} \cdot \frac{\partial A}{\partial z} + Z_{12} \cdot B \approx 0, \quad \gamma_{12} \cdot A + \frac{i}{k_0} Z_{21} \cdot \frac{\partial A}{\partial z} + (\tilde{\beta}_{13} - Z_{22}) \cdot B + \frac{i}{k_0} \tilde{\beta}_{33} \cdot \frac{\partial B}{\partial z} \approx 0 \quad (25)$$

341 The evolution equations for the slowly varying amplitudes $A(x,y,z)$, $B(x,y,z)$ of the VLF beams are derived. The

342 monochromatic beams are considered, when the frequency ω is fixed and the amplitudes do not depend on time t . Looking

343 for the solutions for the EM field as $\vec{E}, \vec{H} \sim \exp(i\omega t - ik_x x - ik_y y)$, Maxwell's equations are:

$$344 \quad \begin{aligned} -ik_y H_z - \frac{\partial H_y}{\partial z} &= ik_0 D_x, & \frac{\partial H_x}{\partial z} + ik_x H_z &= ik_0 D_y, & -ik_x H_y + ik_y H_x &= ik_0 D_z \\ -ik_y E_z - \frac{\partial E_y}{\partial z} &= -ik_0 H_x, & \frac{\partial E_x}{\partial z} + ik_x E_z &= -ik_0 H_y, & -ik_x E_y + ik_y E_x &= -ik_0 H_z \end{aligned} \quad (26)$$

345 Here $D_x = \varepsilon_{11} E_x + \varepsilon_{12} E_y + \varepsilon_{13} E_z$. From Eqs. (21), the equations for E_x , E_z through E_y , H_y are:

346

$$E_x = \frac{1}{\Delta_y} \left\{ [\epsilon_{13} \cdot \epsilon_{32} - (\epsilon_{12} + \frac{k_x k_y}{k_0^2}) \cdot (\epsilon_{33} - \frac{k_y^2}{k_0^2})] E_y + \frac{i}{k_0} (\epsilon_{33} - \frac{k_y^2}{k_0^2}) \frac{\partial H_y}{\partial z} + \frac{k_x}{k_0} \epsilon_{13} \cdot H_y + \frac{i k_y}{k_0^2} \epsilon_{13} \frac{\partial E_y}{\partial z} \right\} \quad (27)$$

$$E_z = \frac{1}{\Delta_y} \left\{ [\epsilon_{31} \cdot (\epsilon_{12} + \frac{k_x k_y}{k_0^2}) - \epsilon_{32} \cdot (\epsilon_{11} - \frac{k_y^2}{k_0^2})] E_y - \frac{i}{k_0} \epsilon_{31} \frac{\partial H_y}{\partial z} - \frac{k_x}{k_0} \cdot (\epsilon_{11} - \frac{k_y^2}{k_0^2}) H_y - \frac{i k_y}{k_0^2} \cdot (\epsilon_{11} - \frac{k_y^2}{k_0^2}) \frac{\partial E_y}{\partial z} \right\}$$

347 In Eq. (27), $\Delta_y \equiv (\epsilon_{11} - \frac{k_y^2}{k_0^2}) \cdot (\epsilon_{33} - \frac{k_y^2}{k_0^2}) - \epsilon_{31} \cdot \epsilon_{13}$. The equations for E_y, H_y obtained from the Maxwell equations are:

348

$$\left(\frac{\partial^2}{\partial z^2} - k_x^2 - k_y^2 \right) E_y + i k_y \left(\frac{\partial E_z}{\partial z} - i k_x E_x - i k_y E_y \right) + k_0^2 D_y = 0; \quad -i k_0 \frac{\partial E_x}{\partial z} + k_x k_0 E_z + k_0^2 H_y = 0 \quad (28)$$

349 After substitution of (27) for E_x, E_z into Eqs. (28), the coupled equations for E_y, H_y can be derived. The follow expansion
350 should be used: $k_x = k_0 + \delta k_x$, $|\delta k_x| \ll k_0$, also $|k_y| \ll k_0$. Then, according to (Weiland and Wilhelmsson, 1977):

351

$$-i \cdot \delta k_x \rightarrow \frac{\partial}{\partial x}, \quad -i \cdot k_y \rightarrow \frac{\partial}{\partial y} \quad (29)$$

352 The expansions should be until the quadratic terms of k_y and the linear terms of δk_x . As a result, parabolic equations (Levy
353 2000) for the **slowly varying amplitudes** A and B are derived. In the lower ionosphere/atmosphere, where the effective
354 permittivity reduces to a scalar $\epsilon(\omega, z)$, they are independent:

355

$$\frac{\partial A}{\partial x} + \frac{i}{2k_0} \left(\frac{\partial^2 A}{\partial y^2} + \frac{\partial^2 A}{\partial z^2} \right) + \frac{i k_0}{2} \cdot (\epsilon - 1) A = 0 \quad (30a)$$

$$\frac{\partial B}{\partial x} + \frac{i}{2k_0} \left(\frac{1}{\beta} \frac{\partial}{\partial z} (\beta \frac{\partial B}{\partial z}) + \frac{\partial^2 B}{\partial y^2} \right) + \frac{i k_0}{2} \cdot (\epsilon - 1) B = 0$$

356

357 Here $\beta \equiv \epsilon^{-1}$. Accounting for the presence of gyrotrropic layer and the tensor impedance boundary conditions at the upper
358 boundary $z = L_z$ of the VLF waveguide, the equations for the slowly varying amplitudes in the general case are:

359

$$\frac{\partial A}{\partial x} + \frac{i}{2k_0} \left(\frac{\partial^2 A}{\partial y^2} + \frac{\partial^2 A}{\partial z^2} \right) + \frac{i k_0}{2} \cdot (\tilde{\epsilon}_{22} - 1) \cdot A + \frac{\gamma_{21}}{2} \frac{\partial B}{\partial z} + \frac{i k_0}{2} \cdot \gamma_{23} B = 0 \quad (30b)$$

$$\frac{\partial B}{\partial x} + \frac{i}{2k_0} \left(\frac{1}{\tilde{\beta}_{11}} \frac{\partial}{\partial z} (\tilde{\beta}_{33} \frac{\partial B}{\partial z}) + \frac{\partial^2 B}{\partial y^2} \right) + \frac{i}{2\tilde{\beta}_{11}} \frac{\partial}{\partial z} (\gamma_{12} A) + \frac{1}{2\tilde{\beta}_{11}} \frac{\partial}{\partial z} (\tilde{\beta}_{13} B) + \frac{i k_0}{2\tilde{\beta}_{11}} \gamma_{32} A + \frac{\tilde{\beta}_{31}}{2\tilde{\beta}_{11}} \frac{\partial B}{\partial z} + \frac{i k_0}{2} \cdot \left(\frac{1}{\tilde{\beta}_{11}} - 1 \right) \cdot B = 0$$

360 In Eqs. (30b),

361

$$\gamma_{12} \equiv \frac{\epsilon_{13} \cdot \epsilon_{32} - \epsilon_{12} \cdot \epsilon_{33}}{\Delta}, \quad \gamma_{21} \equiv \frac{\epsilon_{23} \cdot \epsilon_{31} - \epsilon_{21} \cdot \epsilon_{33}}{\Delta}, \quad \gamma_{23} \equiv \frac{\epsilon_{21} \cdot \epsilon_{13} - \epsilon_{23} \cdot \epsilon_{11}}{\Delta}, \quad \gamma_{32} \equiv \frac{\epsilon_{31} \cdot \epsilon_{12} - \epsilon_{32} \cdot \epsilon_{11}}{\Delta}, \quad \tilde{\beta}_{11} \equiv \frac{\epsilon_{11}}{\Delta}, \quad \tilde{\beta}_{13} \equiv \frac{\epsilon_{13}}{\Delta},$$

$$\tilde{\beta}_{31} \equiv \frac{\epsilon_{31}}{\Delta}, \quad \tilde{\beta}_{33} \equiv \frac{\epsilon_{33}}{\Delta}; \quad \Delta \equiv \epsilon_{11} \cdot \epsilon_{33} - \epsilon_{13} \cdot \epsilon_{31}; \quad \tilde{\epsilon}_{22} \equiv \epsilon_{22} + \frac{\epsilon_{21}(\epsilon_{13}\epsilon_{32} - \epsilon_{12}\epsilon_{33}) + \epsilon_{23}(\epsilon_{31}\epsilon_{12} - \epsilon_{32}\epsilon_{11})}{\Delta}$$

362 Eqs. (30b) are reduced to Eqs. (30a) when the effective permittivity is scalar. At the Earth's surface $z = 0$, the impedance
363 conditions are reduced, accounting for that the medium is isotropic and the conductivity of the Earth is finite, to the form:

364

$$E_y = Z_{0E} H_x, \quad E_x = -Z_{0E} H_y, \quad Z_{0E} \equiv \left(\frac{i\omega}{4\pi\sigma_E} \right)^{1/2} \quad (31a)$$

365

366 Here $\sigma_E \sim 10^8 \text{ s}^{-1}$ is the Earth's conductivity. The boundary conditions (31a) at the Earth's surface, where
 367 $Z_{022} = Z_{021} \equiv Z_{0E}$, $Z_{012} = Z_{021} = 0$, $\beta_{33} = \varepsilon^{-1}(z=0)$, $\gamma_{12} = 0$, $\tilde{\beta}_{13} = 0$, can be rewritten as

368

$$E_y + \frac{i}{k_0} Z_{0E} \frac{\partial E_y}{\partial z} = 0, \quad \frac{i}{\varepsilon(z=0)k_0} \frac{\partial H_y}{\partial z} + Z_{0E} H_y = 0 \quad (31b)$$

369

370 Eqs. (30), combined with the boundary conditions (25) at the upper boundary of the VLF waveguide $z = L_z$, and with the
 371 boundary conditions at the Earth's surface (31b), are used to simulate the VLF wave propagation. The surface impedance of
 372 the Earth has been calculated from the Earth's conductivity, see eq. (31a). The **boundary** conditions to solution of Eqs. (30),
 373 (25), (31b) are chosen in the form

374

$$A(x=0, y, z) = 0, \quad B(x=0, y, z) = B_0 \exp\left(-\left((y-0.5L_y)/y_0\right)^{2n}\right) \exp\left(-\left((z-z_1)/z_0\right)^{2n}\right), \quad n = 2 \quad (32)$$

375

376 In relations (32), z_1 , z_0 , y_0 and B_0 are the vertical position of maximum value, the vertical and transverse characteristic
 377 dimensions of the spatial distribution and the maximum value of H_y , respectively at the input of the system, $x=0$. The
 378 size of the computing region along OY axis is, by the order of value, $L_y \sim 2000$ km. Because the gyrotropic layer is relatively
 379 thin and is placed at the upper part of the VLF waveguide, the beams are excited near the Earth's surface, the wave
 380 diffraction in this gyrotropic layer along OY axis is quite small, i.e. the terms $\partial^2 A/dy^2$, $\partial^2 B/dy^2$ are small there as well.
 381 Contrary to this, the wave diffraction is very important in the atmosphere in the lower part of the VLF waveguide, near the
 382 Earth's surface. To solve the problem of the beam propagation, the method of splitting with respect to physical factors has
 383 been applied (Samarskii 2001). Namely, the problem has been approximated by the finite differences:

383

$$\vec{C} \equiv \begin{pmatrix} A \\ B \end{pmatrix}, \quad \frac{\partial \vec{C}}{\partial x} + \hat{L}_y \vec{C} + \hat{L}_z \vec{C} = 0 \quad (33)$$

384

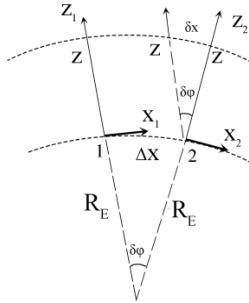
385 In the terms $\hat{L}_y \vec{C}$, the derivatives with respect to y are included, whereas all other terms are included into $\hat{L}_z \vec{C}$. Then the
 386 following fractional steps have been applied, the first one is along y , the second one is along z :

386

$$\frac{\vec{C}^{p+1/2} - \vec{C}^p}{h_x} + \hat{L}_y \vec{C}^{p+1/2} = 0, \quad \frac{\vec{C}^{p+1} - \vec{C}^{p+1/2}}{h_x} + \hat{L}_z \vec{C}^{p+1} = 0 \quad (34)$$

387

388 The region of simulation is $0 < x < L_x = (1000 - 2000)$ km, $0 < y < L_y = (2000 - 3000)$ km, $0 < z < L_z = 80 - 90$ km. The
 389 numerical scheme (34) is absolutely stable. Here h_x is the step along OX axis, $x_p = p h_x$, $p = 0, 1, 2, \dots$. This step has been
 chosen from the conditions of the simulation results independence on the diminishing h_x .



390

391 **Figure 2. The rotation of the local Cartesian coordinate frame at each step along the Earth's surface h_x on a small angle $\delta\varphi \approx$**
 392 **$\Delta x/R_E$, radians, while $\Delta x=h_x$. The following strong inequalities are valid $h_x \ll L_z \ll R_E$. At the Earth's surface $z=0$.**

393

394 On the simulation at each step along OX axis, the correction on the Earth's curvature has been inserted in adiabatic manner
 395 applying the rotation of the local coordinate frame XOZ . Because the step along x is small $h_x \sim 1$ km $\ll L_z$, this correction of
 396 the \vec{C} results in the multiplier $\exp(-ik_0\delta x)$, where $\delta x = z \cdot (h_x/R_E)$, $R_E \gg L_z$ is the Earth's radius (see Fig. 2 and the caption to
 397 this figure). At the distances $x \leq 1000$ km, the simulation results do not depend on the insertion of this correction, whereas at
 398 higher distances some quantitative difference occurs: the VLF beam propagates more closely to the upper boundary of the
 399 waveguide.

400 **3.5 VLF Waveguide Modes and Reflection from the VLF Waveguide Upper
 401 Effective Boundary**

402 In general, our model needs the consideration of the waveguide modes excitations by a current source such as dipole-like
 403 VLF radio source and lightning discharge. Then, the reflection of the waves incident on the upper boundary ($z=L_z$) of the
 404 effective WGEI can be considered. There will be possible to demonstrate that this structure has indeed good enough
 405 waveguiding properties. Then, in the model described in the present paper, the VLF beam is postulated already on the input
 406 of the system. To understand, how such a beam is excited by the, say, dipole antenna near the lower boundary $z=0$ of the
 407 WGEI, the formation of the beam structure based on the mode presentation should be searched. Then the conditions of the
 408 radiation (absence of ingoing waves) (12) can be used as the boundary conditions for the VLF beam radiated to the upper
 409 ionosphere/magnetosphere. Due to a relatively large scale of the inhomogeneity in this region, the complex geometrical
 410 optics (Rapoport et al., 2014) would be quite suitable for modeling a beam propagation, even accounting for the wave
 411 dispersion in magnetized plasma. The proper effective boundary condition, similarly to (Rapoport et al., 2014) would allow
 412 to make relatively accurate matching between the regions, described by the full wave electromagnetic approach with
 413 Maxwell's equations and complex geometrical optics (FWEM-CGO approach). All of these materials are not included in this
 414 paper, but will be delivered in the two future papers. The first paper will be dedicated to the modeling VLF waves
 415 propagating in WGEI based on the field expansion as a set of eigenmodes of the waveguide (the mode presentation

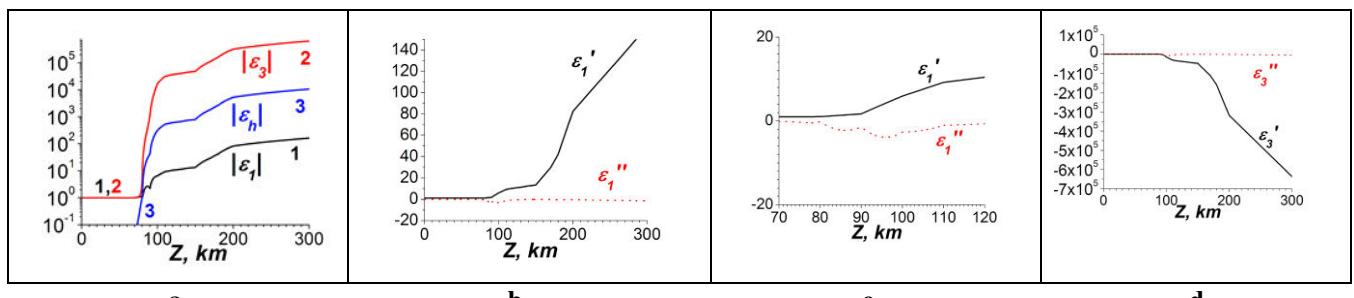
approach). The second paper will deal with the leakage of the VLF beam from the WGEI into the upper ionosphere and magnetosphere and the VLF beam propagation in these media. Here we describe only one result, which concerns the mode excitation in the WGEI, because this result is principally important for the justification of TIMEB method. It was shown that more than five lowest modes of the WGEI are strongly localized in the atmosphere-lower ionosphere. Their longitudinal wavenumbers are close to the corresponding wavenumbers of EM-waves in the atmosphere. This fact convinced that the TIMEB method can be applied to the propagation of the VLF electromagnetic waves in the WGEI.

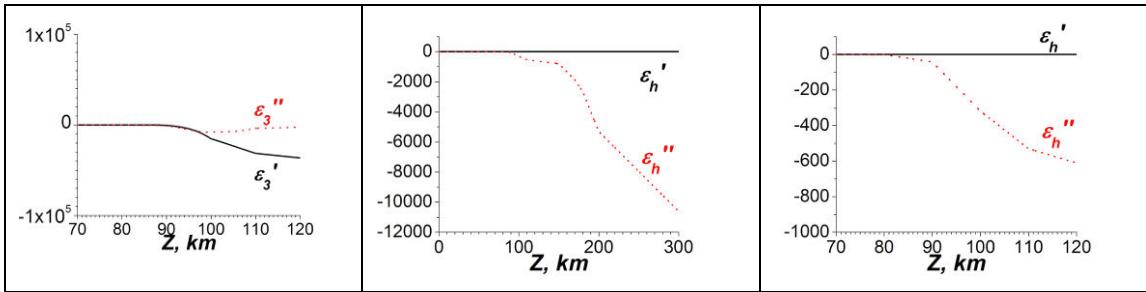
4. Modeling Results

The dependencies of the permittivity components ε_1 , ε_3 , ε_h in the coordinate frame associated with the geomagnetic field \vec{H}_0 are given in Fig. 3. The parameters of the ionosphere used for modeling are taken from (Al'pert 1972; Alperovich and Fedorov 2007; Kelley 2005; Schunk and Nagy 2010; Jursa 1985). The typical results of simulations are presented in Fig. 4. The parameters of the ionosphere correspond to Fig. 3. The angle θ (Fig. 1) is 45° . The VLF frequency is $\omega = 10^5 \text{ s}^{-1}$, $f = \omega/2\pi \approx 15.9 \text{ kHz}$. The Earth's surface is assumed as ideally conductive at the level $z = 0$. The values of EM-field are given in absolute units. The magnetic field is measured in Oersteds (Oe), or Gauss (Gs), $1 \text{ Gs} = 10^{-4} \text{ T}$, whereas the electric field is also in Gs, $1 \text{ Gs} = 300 \text{ V/cm}$.

Note that in the absolute (Gaussian) units the magnitudes of the magnetic field component $|H_y|$ are the same as ones of the electric field component $|E_z|$ in the atmosphere region where the permittivity is $\varepsilon \approx 1$. Below in the Fig. 4 caption, the correspondence between the absolute units and practical SI units is given.

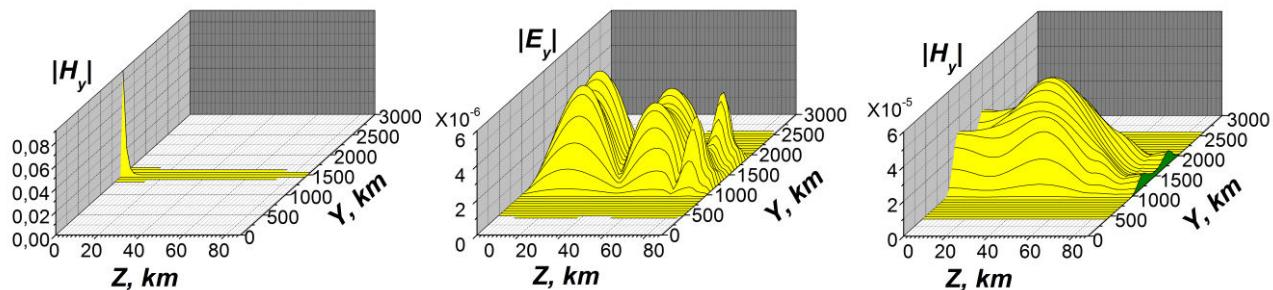
It is seen that the absolute values of the permittivity components increase sharply above $z = 75 \text{ km}$. The behavior of the permittivity components is step-like, as seen from Fig. 3a. Therefore, the results of simulations are tolerant to the choice of the upper wall position of the Earth's surface-ionosphere waveguide. The computed components of the tensor impedance at $z = 85 \text{ km}$ are: $Z_{011} = 0.087 + i0.097$, $Z_{021} = 0.085 + i0.063$, $Z_{012} = -0.083 - i0.094$, $Z_{022} = 0.093 + i0.098$. So, a condition $|Z_{0\alpha\beta}| \leq 0.15$ is satisfied there, which is necessary for the applicability of the boundary conditions (25). The maximum value of the H_y component is $0.1 \text{ Oe} = 10^{-5} \text{ T}$ in Fig. 4a for the initial VLF beam at $x = 0$. This corresponds to the value of E_z component of $0.1 \text{ Gs} = 30 \text{ V/cm}$. At the distance $x = 1000 \text{ km}$ the magnitudes of the magnetic field H_y are of about $3 \cdot 10^{-5} \text{ Oe} = 3 \text{ nT}$, whereas the electric field E_y is of $3 \cdot 10^{-6} \text{ Gs} \approx 1 \text{ mV/cm}$.





442

443 **Figure 3. (a) The vertical dependencies of the modules of components of the permittivity in the frame associated with the**
 444 **geomagnetic field $|\varepsilon_l|, |\varepsilon_3|, |\varepsilon_h|$, curves 1, 2, 3 correspondingly. (b) – (g) The real (corresponding lines with the values denoted by**
 445 **one prime) and imaginary (corresponding lines with the values denoted by two primes) parts of the components $\varepsilon_l, \varepsilon_3, \varepsilon_h$, general**
 446 **and detailed views.**



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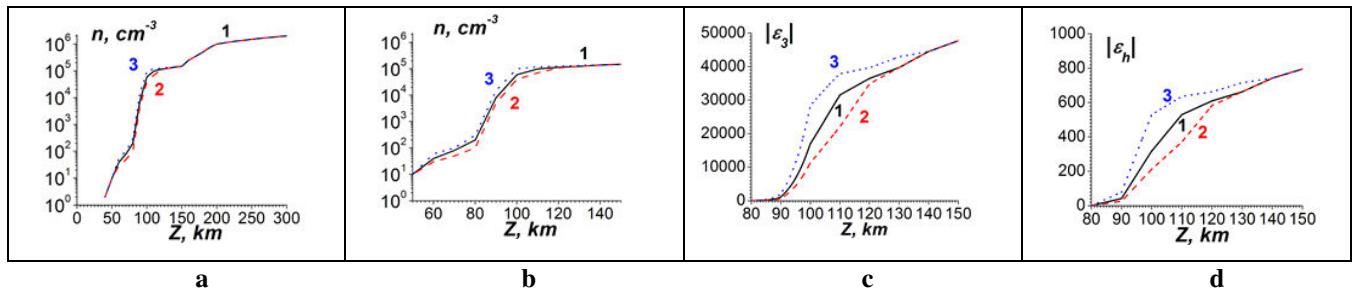
451 **Figure 4. Part a) is the initial distribution of $|H_y|$ at $x = 0$. Parts b), c) are $|E_y|$ and $|H_y|$ at $x = 600$ km. Parts d), e) are $|E_y|$ and $|H_y|$ at**
 452 **$x = 1000$ km. For the electric field maximum value (Fig. d) is $3 \cdot 10^{-6}$ Gs ≈ 1 mV/cm, for the magnetic field maximum value (Fig. e) is**
 453 **$3 \cdot 10^{-5}$ Gs ≈ 3 nT. At the altitudes $z < 75$ km, $|E_z| \approx |H_y|$. $\omega = 1 \cdot 10^5$ s⁻¹; $\theta = 45^\circ$**

454 The wave beams are localized within the WGEI, $0 < z < 75$ km, mainly in the regions with the isotropic permittivity (see
 455 Fig. 4b-e). The mutual transformations of the beams of different polarizations occur near the waveguide upper boundary due
 456 to the anisotropy of the ionosphere within the thin layer $75 \text{ km} < z < 85 \text{ km}$ (Fig. 4b, d). These transformations depend on
 457 the permittivity component values of the ionosphere at the altitude $z > 80$ km and on the components of the tensor

458 impedance. Therefore, the measurements of the phase and amplitude modulations of different EM components near the
 459 Earth's surface can provide information on the properties of the lower and middle ionosphere.

460 In accordance with boundary condition (32), we suppose that when entering the system (at $X = 0$), only one of the two
 461 polarization modes is excited, namely, the TM mode, i.e. at $X = 0$, $H_y \neq 0$; $E_y = 0$ (Fig. 4a). Upon further propagation of the
 462 beam with such boundary conditions at the entrance to the system in a homogeneous isotropic waveguide, the property of the
 463 electromagnetic field described by the relation $H_y \neq 0$; $E_y = 0$ will remain valid. The qualitative effect due to the presence
 464 of gyrotropy (a) in a thin bulk layer near the upper boundary of WGEI and (b) in the upper boundary condition with complex
 465 gyrotropic and anisotropic impedance is as follows. During beam propagation in the WGEI, the TE polarization mode with
 466 the corresponding field components, including E_y , is also excited. This effect is illustrated in Figs. 4 b, d.

467 The magnitude of the E_y component depends on the values of the electron concentration at the altitudes $z = 75 - 100$ km. In
 468 Fig. 5a, b the different dependencies of the electron concentration $n(z)$ are shown (see solid (1), dash (2) and dot (3) lines).
 469 The corresponding dependencies of the component absolute values of the permittivity are shown in Figs. 4c and 4d.



470
 471 **Figure 5. Different profiles of the electron concentration $n(z)$ used in simulations: solid, dash, and dot lines correspond to
 472 undisturbed, decreased and increased concentrations, respectively. (a) the detailed view; (b) general view; (c) and (d) the
 473 permittivity $|\epsilon_3|$ and $|\epsilon_h|$ modules.**

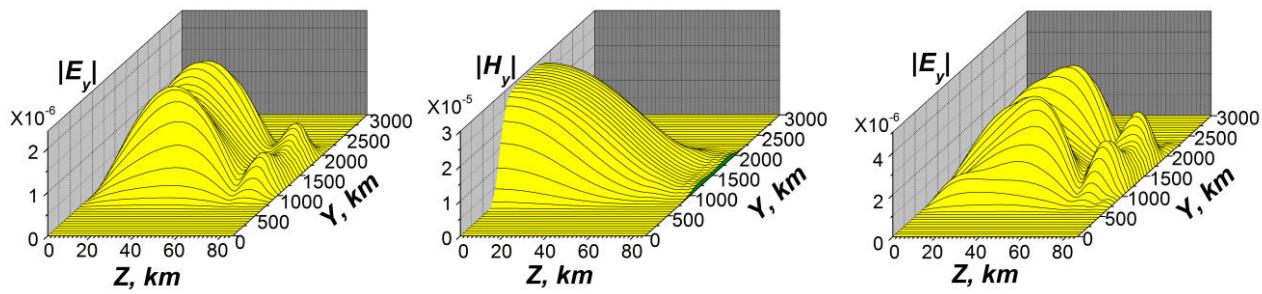
474
 475 The distributions of $|E_y|$, $|H_y|$ at $x = 1000$ km are given in Fig. 6. Results in Figs. 6a and 6b correspond to the solid (1) curve
 476 $n(z)$ in Fig. 5; Figs. 6c and 6d correspond to the dash (2) curve; Figs. 6e and 6f correspond to the dot (3) curve in Fig. 5. The
 477 initial H_y beam is the same and is given in Fig. 4a. The values of the tensor impedance for these three cases are presented in
 478 Table 1.

479
 480 **Table 1. The values of tensor impedance components corresponded to the data shown in Fig. 5.**

Component of the tensor impedance	Z_{011}	Z_{021}	Z_{012}	Z_{022}
Undisturbed concentration (curves 1 in Fig. 5)	$0.088 + i0.098$	$0.085 + i0.063$	$-0.083 - i0.094$	$0.093 + i0.098$
Decreased concentration	$0.114 + i0.127$	$0.107 + i0.079$	$-0.105 - i0.127$	$0.125 + i0.125$

(curves 2 in Fig. 5)				
Increased concentration	$0.067 + i0.0715$	$0.061 + i0.051$	$-0.060 - i0.070$	$0.069 + i0.072$
(curves 3 in Fig. 5)				

481



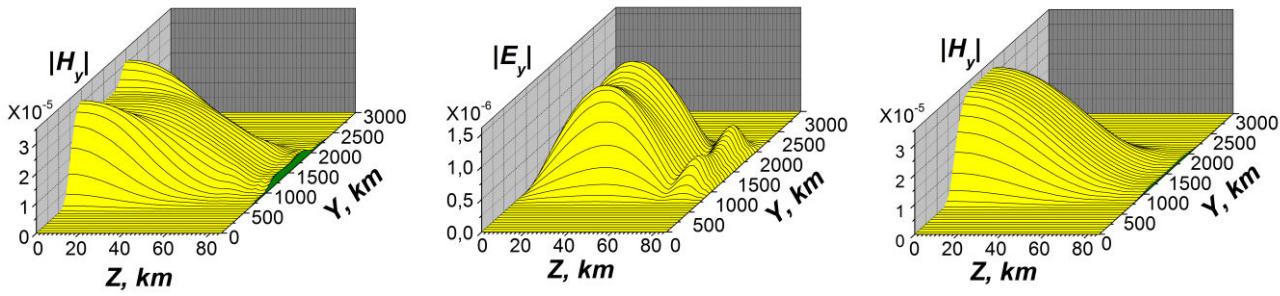
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a

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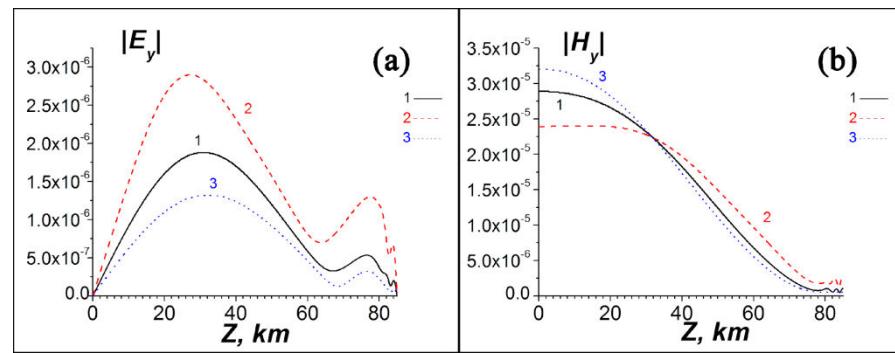
d

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485

486 **Figure 6.** Parts a), c), e) are dependencies of $|E_y|$, parts b), d), f) are dependencies of $|H_y|$ at $x = 1000$ km; $\omega = 1 \cdot 10^5 s^{-1}$; $\theta = 45^\circ$.
487 The initial beams are the same as in Fig. 4, a). Parts a), b) correspond to the solid (1) curves in Fig. 5; parts c), d) are for the dash
488 (2) curves; parts e), f) correspond to the dot (3) curves there.



489

490 **Figure 7.** The dependencies of EM components on the altitude z in the center of the waveguide, $y = 1500$ km, for the different
491 profiles of the electron concentration. The solid (1), dash (2), and dot (3) curves correspond to the different profiles of the electron
492 concentration in Fig. 5, a), b), the same kinds of curves; $\omega = 1 \cdot 10^5 s^{-1}$; $\theta = 45^\circ$

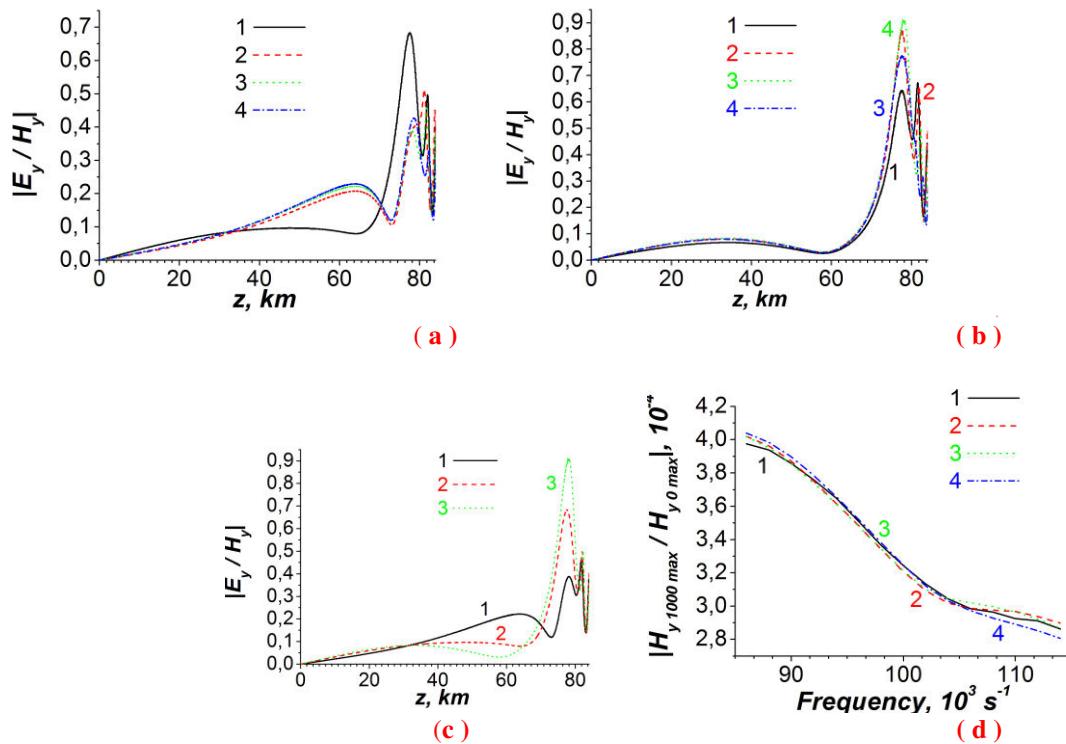
493 The distributions of $|E_y|$, $|H_y|$ on z at $x = 1000$ km in the center of the waveguide, $y = 1500$ km, are given in Fig. 7. These
494 simulations show that the change in the complex tensors of both volume dielectric permittivity and impedance at the lower
495 and upper boundaries of effective WGEI influence remarkably on the VLF losses. The modulation of the electron
496 concentration at the altitudes above $z= 120$ km affects the excitation of E_y component within the waveguide rather weakly.

497 5. Influence of the parameters of WGEI on the polarization transformation and losses of 498 the propagating VLF waves

499 An important effect of the gyrotropy and anisotropy is the corresponding transformation of the field polarization
500 during the propagation in the WGEI, which is absent in the ideal metal planar waveguide without gyrotropy and anisotropy.
501 We will show that this effect is quite sensitive to the carrier frequency of the beam, propagating in the WGEI, inclination of
502 the geomagnetic field and perturbations in the electron concentration, which can vary under the influences of the powerful
503 sources placed “below”, “above” and “inside” the ionosphere. In the real WGEI, the anisotropy and gyrotropy are connected
504 with the volume effect and effective surface tensor impedances at the lower and upper surfaces of the effective WGEI where
505 $z=0$ and $z= L_z$ (Fig. 1). For the corresponding transformation of the field polarization determination, we introduce the
506 characteristic polarization relation $|E_y / H_y| (z; y = L_y / 2; x = x_0)$, taken at the central plane of the beam ($y=L_y$) at the
507 characteristic distance ($x=x_0$) from the beam input/VLF transmitter. The choice of the characteristic polarization parameter
508 ($|E_y/H_y|$) and its dependence on the vertical coordinate (z) is justified by conditions (1) – (3). (1) The WGEI is similar to the
509 ideal planar metallized waveguide because, first, the tensor $\hat{\varepsilon}$ is different from the isotropic \hat{I} only in the relatively small
510 upper part of the WGEI in the altitude range from 75-80 to 85 km (see Fig. 1). Second, both the Earth and ionosphere
511 conductivity are quite high and corresponding impedances are quite low. In particular, the elements of the effective tensor
512 impedance at the upper boundary of WGEI are small, $|Z0\alpha\beta| \leq 0.1$ (see, for example, Table 1). (2) Respectively, the carrier
513 modes of the VLF beam are close to the modes of the ideal metallized planar waveguide. These modes are subdivided into
514 the sets of uncoupled (E_x, H_y, E_z) and (H_x, E_y, H_z) modes. The detailed search of the propagation of the separate eigenmodes
515 of the WGEI is not a goal of this paper, and respectively, will be the subject of the separate paper. (3) Because we have
516 adopted for the initial beam(s) the input boundary conditions in the form (32) (with $H_y \neq 0$, $E_y = 0$), the above mentioned
517 value $|E_y / H_y| (z; y = L_y / 2; x = x_0)$ characterizes the mode coupling and corresponding transformation of the polarization
518 at the distance x_0 from the beam input due to the presence of the volume and surface gyrotropy and anisotropy in the real
519 WGEI. The results presented below are obtained for $x_0 = 1000$ km, that is, by the order of value, a typical distance, for
520 example, between the VLF transmitter and receiver of the European VLF/LF radio network (Biagi et al. 2015). Other
521 parameter characterizing the propagation of the beam in the WGEI, the effective total loss parameter is $|H_{y_{max}}(x=x_0)/$
522 $H_{y_{max}}(x=0)|$. Note that this parameter characterizes both dissipative and diffraction losses. The last are connected with beam
523 spreading in the transverse (y) direction during the propagation in the WGEI.

In Fig. 8 the polarization and loss characteristics dependencies on both the carrier beam frequency and the angle

525 θ between the geomagnetic field and the vertical directions (see Fig. 1) are shown.



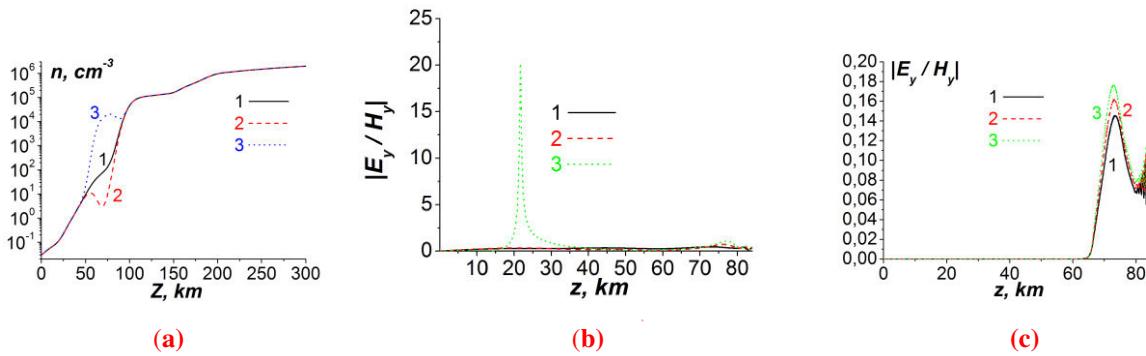
530 **Figure 8. Characteristics of the polarization transformation parameter $|E_y/H_y|$ (a-c) and effective coefficient of the total losses at**
 531 **the distance $x_0=1000$ km from the beam input (d); corresponding altitude dependence of the electron concentration is shown in**
 532 **Fig. 5b line (1); (a, b) and (d) - dependences of the polarization parameter (a, b) and total losses (d) on the vertical coordinate for**
 533 **different angles θ , respectively; black (1), red (2), green (3) and blue (4) curves in Figs. a, b and d correspond to θ equal to**
 534 **$5^\circ, 30^\circ, 45^\circ$ and 60° , respectively; (a) and (b) correspond to frequencies $\omega = 0.86 \cdot 10^5 s^{-1}$ and $\omega = 1.14 \cdot 10^5 s^{-1}$, respectively; (c) -**
 535 **dependence of the polarization parameter on the vertical coordinate for the different frequencies; black (1), red (2) and green (3)**
 536 **lines correspond to the frequencies $0.86 \cdot 10^5$, $1 \cdot 10^5$ and $1.14 \cdot 10^5 s^{-1}$, respectively and $\theta = 45^\circ$.**

537 In Fig. 8a-c the altitude dependence of the polarization parameter $|E_y/H_y|$ exhibits two main maxima in the WGEI. The first
 538 one lies in the gyrotropic region above 70 km, while the second one in the isotropic region of the WGEI. As it is seen from
 539 Fig. a, b, the value of the (larger) second maximum increases, while the value of the first maximum decreases and its
 540 position shifts to the lower altitudes with increasing frequency. At the higher frequency ($\omega = 1.14 \cdot 10^5 s^{-1}$), the larger
 541 maximum of the polarization parameter corresponds to the intermediate value of the angle $\theta = 45^\circ$ (Fig. 8 b); for the lower
 542 frequency ($\omega = 0.86 \cdot 10^5 s^{-1}$), the largest value of the first (higher) maximum corresponds to the almost vertical direction of
 543 the geomagnetic field ($\theta = 5^\circ$, Fig. 8 a). For the intermediate value of the angle ($\theta = 45^\circ$), the largest value of the main
 544 maximum corresponds to the higher frequency ($\omega = 1.14 \cdot 10^5 s^{-1}$) in the considered frequency range (Fig. 8 c). The total losses
 545 increase monotonically with increasing frequency and depend weakly on the value of θ (Fig. 8 d).

546 To model the effect of increasing and decreasing the electron concentration n_e in the lower ionosphere on the
 547 polarization parameter, we have used the following parameterization for the n_e change $\Delta n_e = n_e(z) - n_{0e}(z)$ of the electron
 548 concentration, where $n_{0e}(z)$ is the unperturbed altitude distribution of the electron concentration:

$$549 \Delta n_e(z) = n_{0e}(z)\Phi(z); \Phi(z) = [F(z)] - \frac{(z - z_2)^2}{\Delta z_{12}^2}[F(z_1)] - \frac{(z - z_1)^2}{\Delta z_{12}^2}[F(z_2)]; F(z) = f_{00} \cdot ch^{-2}\{[z - (\frac{z_1 + z_2}{2})]/\Delta z\} \quad (35)$$

550
 551 In Eqs. (35), $\Delta z_{12} \equiv z_2 - z_1$; Δz is the effective width of the electron concentration perturbation altitude distribution. The
 552 perturbation Δn_e is concentrated in the range of altitudes $z_1 \leq z \leq z_2$ and is equal to zero outside this region,
 553 $\Delta n_e(z_1) = \Delta n_e(z_2) = 0$, while $\Phi(z_1) = \Phi(z_2) = 0$.



554
 555
 556 **Figure 9.** (a) Decreased and increased electron concentration (line 2, red color) and (line 3, blue) correspond to $f_{00} = -1.25$ and
 557 $f_{00} = 250$, respectively, relative to reference concentration (line 1, black) with parametrization conditions (see Eqs. (35)): $z_1=50$
 558 km, $z_2=90$ km, $\Delta z = 20$ km; (b) and (c) are the polarization parameter $|E_y/H_y|$ altitude distribution, for decreased and increased
 559 electron concentration, respectively. In (b) and (c) lines 1, 2 and 3 correspond to ω values $0.86 \cdot 10^5 \text{ s}^{-1}$, $1 \cdot 10^5 \text{ s}^{-1}$ and
 560 $1.14 \cdot 10^5 \text{ s}^{-1}$, respectively. Angle θ is equal to 45° .

562 The change in the concentration in the lower ionosphere causes rather nontrivial effect on the parameter of the
 563 polarization transformation $|E_y/H_y|$, Fig. 9 a-c. Note that either increase or decrease in the ionosphere plasma concentration
 564 have been reported as a result of seismogenic phenomena, tsunamis, particle precipitation in the ionosphere due to wave-
 565 particle interaction in the radiation belts (Pulinets et al. 2005; Shinagawa et al. 2013; Arnoldy et al. 1989; Glukhov et al.
 566 1992; Tolstoy et al. 1986) etc. The changes in the $|E_y/H_y|$ due to increase or decrease in electron concentration vary by
 567 absolute values from dozens to thousands percent, as it is seen from the comparison between Figs. 9b, c (lines 3) and Fig. 8c
 568 (line 3), which corresponds to the unperturbed distribution of the ionospheric electron concentration (see also lines 1 in Figs.
 569 5b and 9a). It is even more interesting that in the case of decreasing (Fig. 9 a, curve 2) electron concentration, the main
 570 maximum of $|E_y/H_y|$ appears in the lower atmosphere (at the altitude around 20 km, Fig. 9 b, curve 3, which corresponds to
 571 $\omega = 1.14 \cdot 10^5 \text{ s}^{-1}$). In the case of increasing electron concentration (Fig. 9 a, curve 3) the main maximum of $|E_y/H_y|$ appears
 572 near the E region of the ionosphere (at the altitude around 77 km, Fig. 9 c). The secondary maximum, which is placed, in the

absence of the perturbation of the electron concentration, in the lower atmosphere (Fig. 8 c, curves 2, 3), or mesosphere/ionosphere D region ((Fig. 8 c, curve 1), practically disappears or just is not seen in the present scale, in the case under consideration (Fig. 9 c, curves 1-3).

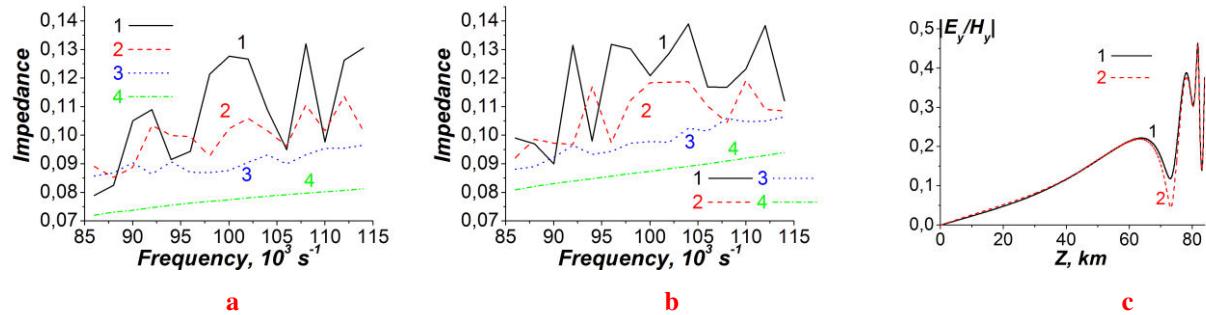
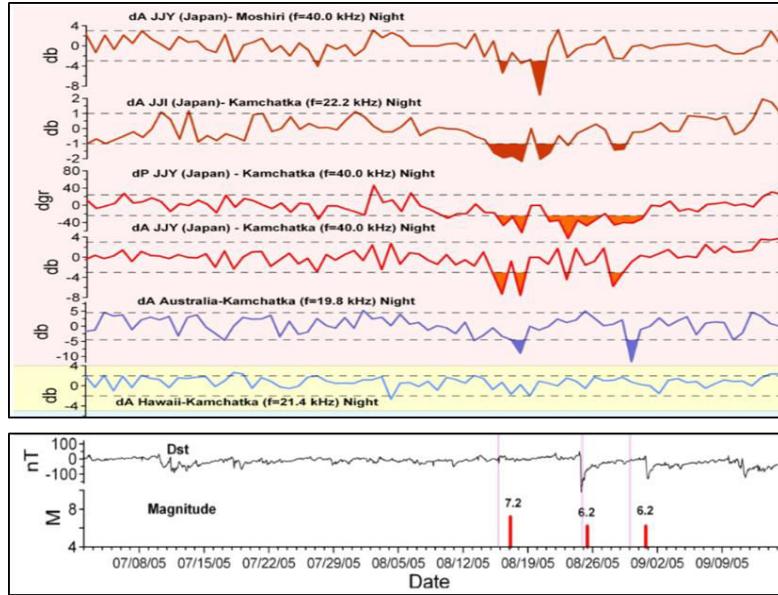


Figure 10. (a, b) Frequency dependencies of the real (a) and imaginary (b) parts of the effective tensor impedance Z_{011} component at the upper boundary ($z=L_z$, see Fig. 1) of the WGEI. Lines 1 (black), 2 (red), 3 (blue) and 4 (green) correspond to $\theta = 5^\circ, 30^\circ, 45^\circ$ and 60° degrees respectively; (c) polarization parameter $|E_y/H_y|$ altitude dependency at the frequency $\omega = 0.86 \cdot 10^5 s^{-1}$ and angle $\theta = 45^\circ$ for the isotropic surface impedance Z_0 at the lower surface of the WGEI equal to 10^4 . Earth conductivity σ equal to $10^9 s^{-1}$, line 1 and $Z_0 = 10^2$ ($\sigma = 10^7 s^{-1}$), line 2.

In Fig. 10, the real (a) and imaginary (b) parts of the surface impedance at the upper boundary of the WGEI have a quasi-periodical character with the amplitude of oscillations occurring around the effective average values (not shown explicitly in Figs. 10 a, b), which decreases with increasing the angle θ . The average $Re(Z_{011})$ and $Im(Z_{011})$ values in general, decrease with increasing angle θ (see Fig. 10a, b). The average values of $Re(Z_{011})$ at $\theta = 5^\circ, 30^\circ, 45^\circ$ and 60° (lines 1-4 in Fig. 10a) and $Im(Z_{011})$ at $\theta = 45^\circ$ and 60° (curves 3, 4 in Fig. 10b) increase with increasing frequency in the frequency range $(0.86-1.14) \cdot 10^5 s^{-1}$. The average $Im(Z_{011})$ value at $\theta = 5^\circ$ and 30° changes in the frequency range $(0.86-1.14) \cdot 10^5 s^{-1}$ non-monotonically with the maximum at $(1-1.1) \cdot 10^5 s^{-1}$. The value of finite impedance at the lower Earth-atmosphere boundary of the WGEI influences on the polarization transformation parameter minimum near the *E* region of the ionosphere (lines 1, 2 in Fig. 10c). The decrease of surface impedance Z_0 at the lower boundary Earth-atmosphere of the WGEI by two orders of magnitude produces the 100% increase of the corresponding $|E_y/H_y|$ minimum at $Z \sim 75$ km (Fig. 10 c).

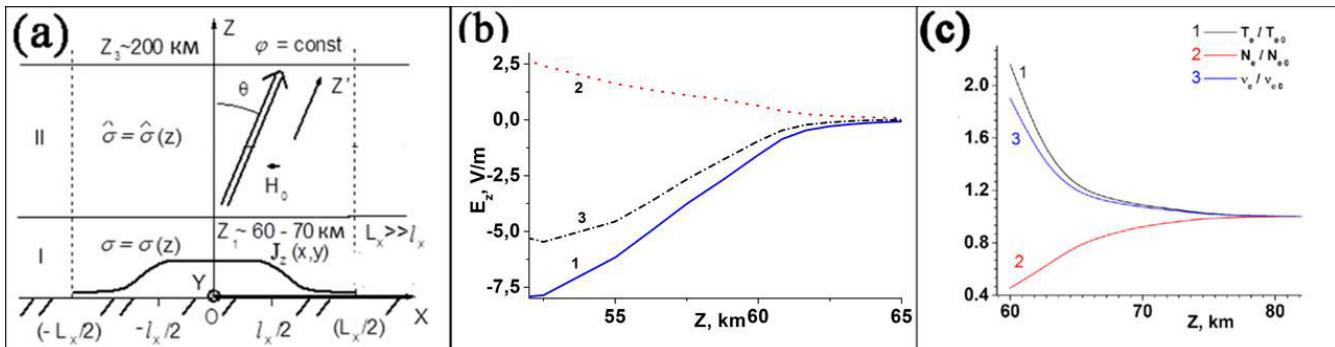
6. Discussion

The observations presented in (Rozhnoi et al., 2015), shows a possibility for seismogenic increasing losses of VLF waves in the WGEI (Fig. 11; see details in (Rozhnoi et al. 2015)). We discuss the qualitative correspondence of our results to these experimental data.



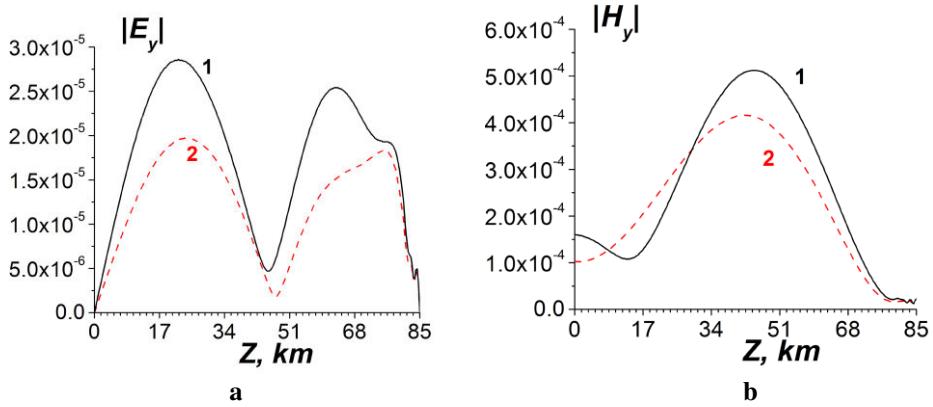
598
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600 **Figure 11.** Averaged residual VLF/LF signals from ground-based observations at the wave paths: JJY-Moshiri, JJI-Kamchatka,
601 JJY-Kamchatka, NWC-Kamchatka, and NPM-Kamchatka. Horizontal dotted lines show 2σ level. The color filled zones highlight
602 values exceeding the -2σ level. In panel below Dst variations and earthquakes magnitude values are shown (from Rozhnoi et al.,
603 2015, see their Fig. 1 but not including the DEMETER data; the work of Rozhnoi et al. (2015) is licensed under a Creative
604 Commons Attribution 4.0 International License (CC BY 4.0)). See other details in (Rozhnoi et al., 2015).



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606 **Figure 12.** Modification of the ionosphere by electric field of seismogenic origin based on the theoretical model (Rapoport et al.
607 2006). (a) Geometry of the model for the determination of the electric field excited by seismogenic current source $J_z(x, y)$ and
608 penetrated into the ionosphere with isotropic (I) and anisotropic (II) regions of the system “atmosphere-ionosphere”. (b) Electric
609 field in the mesosphere in presence of the seismogenic current sources only in the mesosphere (1); in the lower atmosphere (2);
610 both in the mesosphere and in the lower atmosphere (3). (c) Relative perturbations caused by seismogenic electric field,
611 normalized on the corresponding steady-state values in the absence of perturbing electric field, denoted by the index “0”, of
612 electron temperature (T_e/T_{e0}), electron concentration (N_e/N_{e0}), and electron collision frequency (v_e/v_{e0}).



613
614 **Figure 13.** Altitude distributions of the normalized tangential (y) electric (a) and magnetic (b) VLF beam field components in the
615 central plane of the transverse beam distribution ($y=0$) at the distance $x=1000$ km from the input of the system. Line 1 in (a, b)
616 corresponds to the presence of the mesospheric electric current source only with relatively small value of N_e and large v_e . Line 2
617 corresponds to the presence of both mesospheric and near-ground seismogenic electric current sources with relatively large value
618 of N_e and small v_e . Lines 1 and 2 in (a, b) correspond qualitatively to the lines 1 and 3, respectively, in Fig. 12b.

619
620 $\omega=1.5 \cdot 10^5 \text{ s}^{-1}$; $\theta=45^\circ$.

621
622 The modification of the ionosphere due to electric field excited by the near-ground seismogenic current source has
623 been taken into account. In the model (Rapoport et al., 2006), the presence of the mesospheric current source, which
624 followed from the observations (Martynenko et al., 2001; Meek et al., 2004; Bragin, 1974) is also taken into account. It is
625 assumed that the mesospheric current has only the Z-component and is positive, which means that it is directed vertically
626 downward, as is the fair-weather current (curve 1, Fig. 12b). Then suppose that the surface seismogenic current is directed in
627 the same way as the mesospheric current. We first consider the case when the mesospheric current is zero and only the
628 corresponding seismogenic current is present near the earth. The corresponding mesospheric electric field under the
629 condition of a given potential difference between the Earth and the ionosphere (curve 2, Fig. 12 b) is directed opposite to
630 those excited by the corresponding mesospheric current (curve 1, Fig. 12 b). As a result, in the presence of both mesospheric
631 and a seismogenic surface current, the total mesospheric electric field (curve 3, Fig. 12 b) is smaller in absolute value than in
632 the presence of only a mesospheric current (curve 1, Fig. 12 b). It has been shown by Rapoport et al., (2006) that the
633 decrement of losses $|k''|$ for VLF waves in the WGEI is proportional to $|k''| \sim \epsilon'' \sim N_e / v_e$. N_e and v_e decrease and
634 increase, respectively, due to the appearance of a seismogenic surface electric current, in addition to the mesospheric current
635 (curve 3, Fig. 12, b). As a result, the losses increase compared with the case when the seismogenic current is absent and the
636 electric field has a larger absolute value (curve 1, Fig. 12). The increase in losses in the VLF beam, shown in Fig. 13
637 (compare curves 2 and 1 in Figs. 13 a, b) corresponds to an increase in losses with an increase in the absolute value of the
638 imaginary part of the dielectric constant, when a near-surface seismogenic current source appears (curve 3 in Fig. 12 b), in
639 addition to the existing mesospheric current source (curve 2 in Fig. 12 b). This seismogenic increase in losses corresponds
640 qualitatively to the results, presented in (Rozhnoi et al. 2015).

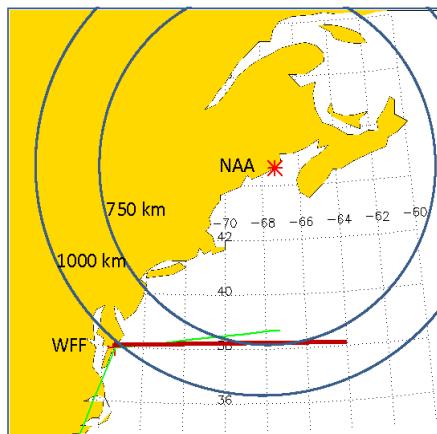
641 The TIMEB is a new method of modeling characteristics of the WGEI. The results of beam propagation in WGEI
642 modeling presented above include the range of altitudes inside the WGEI (see Figs. 4-7). Nevertheless, the TIMEB method
643 described by Eqs. (15)-(19), (22-24), (27), (30) and allows to determine all field components in the range of altitudes
644 $0 \leq z \leq L_{\max}$, where $L_{\max} = 300$ km. The structure and behavior of these eigenmodes in the WGEI and leakage waves will be
645 a subject of separate papers. We present here only the final qualitative result of the simulations. In the range $L_z \leq z \leq L_{\max}$,
646 where $L_z = 85$ km is the upper boundary of the effective WGEI, all field components are (1) at least one order of altitude less
647 than the corresponding maximal field value in the WGEI and (2) field components have the oscillating character along
648 z coordinate and describe the modes, leaking from the WGEI.

649 Let us make a note also on the dependences of the field components in the WGEI on the vertical coordinate (z). The
650 initial distribution of the electromagnetic field with altitude z (Fig. 4a) is determined by the boundary conditions of the beam
651 (see Eq. (32)). The field component includes higher eigenmodes of the WGEI. The higher-order modes experienced quite
652 large losses and practically disappear after beam propagation on 1000 km distance. This determines the change in altitude (z)
653 and transverse (y) distributions of the beam field during propagation along the WGEI. In particular, at the distance $x=600$ km
654 from the beam input (Figs. 4b, c), the few lowest modes of the WGEI along z and y coordinates still persist. At distance
655 $x=1000$ km (Fig. 4d, c; Fig. 6e, f; and Fig. 7a, b), only the main mode persists in the z direction. Note, the described field
656 structure correspond to real WGEI with losses. The gyrotropy and anisotropy causes the volume effects and surface
657 impedance, in distinction to the ideal planar metallized waveguide with isotropic filling (Collin, 2001).

658 The closest approach of the direct investigation of the VLF electromagnetic field profile in the Earth-Ionosphere
659 waveguide was a series of sounding rocket campaigns at mid- and high-latitudes at Wallops Is., VA and Siple Station in
660 Antarctica (Kintner et al., 1983; Brittain et al., 1983; Siefring and Kelly, 1991; Arnoldy and Kintner, 1989), where single-
661 axis E-field and three-axis B-field antennas, supplemented in some cases with in situ plasma density measurements were
662 used to detect the far-field fixed-frequency VLF signals radiated by US Navy and Stanford ground transmitters.

663 The most comprehensive study of the WGEI will be provided by the ongoing NASA VIPER (VLF Trans-Ionospheric
664 Propagation Experiment Rocket) project (PI J. W. Bonnell, UC Berkeley, NASA Grant 80NSSC18K0782). The VIPER
665 sounding rocket campaign is consist of a summer nighttime launch during quiet magnetosphere conditions from Wallops
666 Flight Facility, VA, collecting data through the D, E, and F regions of the ionosphere with a payload carrying the following
667 instrumentation: 2D E- and 3D B-field waveforms, DC-1 kHz; 3D ELF to VLF waveforms, 100 Hz to 50 kHz; 1D wideband
668 E-field measurement of plasma and upper hybrid lines, 100 kHz to 4 MHz; and Langmuir probe plasma density and ion
669 gauge neutral density measurements at a sampling rate of at least tens of Hz. The VIPER project will fly a fully 3D EM field
670 measurement, DC through VLF, and relevant plasma and neutral particle measurements at mid-latitudes through the
671 radiation fields of (1) an existing VLF transmitter (the VLF transmitter Cutler with call sign NAA, the very low
672 frequency (VLF) shore radio station at Cutler, Maine, USA, which transmits, at a frequency of 24 kHz an input power of up
673 to 1.8 megawatts, see Fig. 11) and (2) naturally-occurring lightning transients through and above the leaky upper boundary
674 of the WGEI supported by a vigorous theory and modeling effort in order to explore the vertical and horizontal profile of the

675 observed 3D electric and magnetic radiated fields of the VLF transmitter, and the profile related to the observed plasma and
676 neutral densities. The VLF wave's reflection, absorption, and transmission processes as a function of altitude will be
677 searched making use of the data on the vertical VLF E- and B-field profile.
678



679
680 **Figure 14. Proposed VIPER Trajectory**
681

682 The aim of this experiment will be the investigation of the VLF beams launched by the near-ground source/VLF transmitter
683 with the known parameters and propagating both in the WGEI and leaking from WGEI into the upper ionosphere.
684 Characteristics of these beams will be compared with the theory proposed in the present paper and the theory on leakage of
685 the VLF beams from WGEI, which we will present in the next papers.
686

687 **Conclusions**

688 (1) We have developed the new and highly effective robust method of tensor impedance for the VLF electromagnetic
689 beam propagation in the inhomogeneous waveguiding media – the “tensor impedance method for modeling propagation of
690 electromagnetic beams (TIMEB)” in a multi-layered/inhomogeneous waveguide. The main differences/advantages of the
691 proposed tensor impedance method in comparison with the known method of the impedance recalculating, in particular
692 invariant imbedding methods (Shalashov and Gospodchikov, 2010; Kim and Kim, 2016) are the following: (i) our method is
693 a direct method of the recalculation of tensor impedance, and the corresponding tensor impedance is determined analytically,
694 (see Eq. (22)); (ii) our method applied for the media without non-locality does not need a solution of integral equation(s), as
695 the invariant imbedding method; (iii) the proposed tensor impedance method does not need the revealing the forward and
696 reflected waves. Moreover, even the conditions of the radiation in Eq. (12) at the upper boundary $z=L_{max}$ is determined
697 through the total field components $H_{x,y}$ that makes the proposed procedure technically less cumbersome and practically more
698 convenient.

699 (2) The waveguide includes the region for the altitudes $0 < z < 80 - 90$ km. The boundary conditions are the radiation
700 conditions at $z = 300$ km, which can be recalculated to the lower altitudes as the tensor relations between the tangential
701 components of the EM field. In another words, the tensor impedance conditions have been used at $z = 80 - 90$ km.

702 (3) The application of this method jointly with the previous results of the modification of the ionosphere by seismogenic
703 electric field gives the results, which qualitatively are in agreement with the experimental data on the seismogenic increasing
704 losses of VLF waves/beams propagating in the WGEI.

705 (4) The observable qualitative effect is mutual transformation of different polarizations of the electromagnetic field
706 occur during the propagation. This transformation of the polarization depends on the electron concentration, i.e. the
707 conductivity, of D- and E-layers of the ionosphere at the altitudes 75 – 120 km.

708 (5) Changes in complex tensors of both volume dielectric permittivity and impedances at the lower and upper
709 boundaries of effective WGEI influence remarkably on the VLF losses in the WGEI.

710 (6) An influence is demonstrated on the parameters characterizing the propagation of the VLF beam in the WGEI, in
711 particular, on the parameter of the transformation polarization $|E_y/H_y|$ and tensor impedance at the upper boundary of the
712 effective WGEI, of the carrier beam frequency, inclination of the geomagnetic field and the perturbations in the altitude
713 distribution of the electron concentration in the lower ionosphere

714 (i) The altitude dependence of the polarization parameter $|E_y/H_y|$ has two main maxima in the WGEI: the higher
715 maximum is in the gyrotrropic region above 70 km, while the other is in the isotropic region of the WGEI. The value
716 of the (larger) second maximum increases, while the value of the first maximum decreases and its position shifts to
717 the lower altitudes with increasing frequency. In the frequency range of $\omega = (0.86 - 1.14) \cdot 10^5 \text{ s}^{-1}$, At the higher
718 frequency, the larger maximum polarization parameter corresponds to the intermediate value of the angle $\theta = 45^\circ$;
719 for the lower frequency, the largest value of the first (higher) maximum corresponds to the nearly vertical direction
720 of the geomagnetic field. The total losses increase monotonically with increasing frequency and depend weakly on
721 the value of θ (Fig. 1).

722 (ii) The change in the concentration in the lower ionosphere causes rather nontrivial effect on the parameter of
723 polarization transformation $|E_y/H_y|$. This effect does include the increase and decrease of the maximum value of the
724 polarization transformation parameter $|E_y/H_y|$. The corresponding change of this parameter has large values from
725 dozens to thousands percent. In the case of decreasing electron concentration, the main maximum of $|E_y/H_y|$
726 appears in the lower atmosphere at an altitude of around 20 km. In the case of increasing electron concentration,
727 the main maximum of $|E_y/H_y|$ appears near the E region of the ionosphere (at the altitude around 77 km), while the
728 secondary maximum practically disappears.

729 (iii) The real and imaginary parts of the surface impedance at the upper boundary of the WGEI have a quasi-
730 periodical character with the amplitude of “oscillations” occurring around some effective average values decreases
731 with increasing the angle θ . Corresponding average values of $Re(Z_{II})$ and $Im(Z_{II})$, in general, decrease with

732 increasing angle θ . Average values of $Re(Z_{II})$ for θ equal to $5^\circ, 30^\circ, 45^\circ$ and 60° and $Im(Z_{II})$ corresponding to θ
733 jioequal to 45° and 60° , increase with increasing frequency in the considered frequency range $(0.86-1.14) \cdot 10^5 \text{ s}^{-1}$.
734 The average value of $Im(Z_{II})$ corresponds to θ equal to 5° and 30° , change in the frequency range $(0.86-1.14) \cdot 10^5 \text{ s}^{-1}$
735 non-monotonically, having maximum values around frequency $(1-1.1) \cdot 10^5 \text{ s}^{-1}$.
736 (iv) The value of finite impedance at the lower Earth-atmosphere boundary of the WGEI make quite observable
737 influence on the polarization transformation parameter minimum near the E region of the ionosphere. The decrease
738 of surface impedance Z at the lower boundary Earth-atmosphere of the WGEI in two orders causes the increase of
739 the corresponding minimum value of $|E_y/H_y|$ in $\sim 100\%$.

740 (7) In the range $L_z \leq z \leq L_{\max}$, where $L_z = 85 \text{ km}$ is the upper boundary of the effective WGEI, all field components are
741 (a) at least one order of altitude less than the corresponding maximal value in the WGEI, and (b) field components have the
742 oscillating character (along the z coordinate) and describes the modes, leaking from the WGEI. The detail consideration of
743 the electromagnetic waves leaking from the WGEI will be presented in the separate paper. The initial distribution of the
744 electromagnetic field with z (vertical direction) is determined by the initial conditions on the beam. This field includes
745 higher eigenmodes of the WGEI. The higher-order modes, in distinction to the lower ones, have quite large losses and
746 practically disappear after a beam propagation for 1000 km distance. This circumstance determines the change in altitude (z)
747 distribution of the field of the beam during its propagation along the WGEI. In particular, at the distance $x = 600 \text{ km}$ from
748 the beam input, the few lowest modes of the WGEI along z coordinates are still survived. Further, at $x=1000 \text{ km}$, practically,
749 only the main mode in the z direction remains. This fact reflects in a minimum number of oscillations of the beam field
750 components along z at a given value of x .

751 (8) The proposed propagation of VLF electromagnetic beams in the WGEI model and results will be useful to explore
752 the characteristics of these waves as an effective instrument for diagnostics of the influences on the ionosphere “from above”
753 in the system of Sun-Solar Wind-Magnetosphere-Ionosphere, “from below” from the most powerful meteorological,
754 seismogenic and other sources in the lower atmosphere and lithosphere/Earth, such as hurricanes, earthquakes, tsunamis, and
755 from inside the ionosphere by the strong thunderstorms with lightning discharges, and even from the far space by as gamma-
756 flashes, cosmic rays events.

758 AUTHOR CONTRIBUTION

759 YuR and VG – proposed the idea and concept of the paper, made analytical calculations, have written the initial version of
760 the paper and took part in the revision of the paper; VG developed the code and YuR took part in its verification; YuR, VG,
761 A. Grytsai and A. Culin made numerical modeling; YuR and A.Grytsai provided the administration of the project; VG,
762 YuR, A. Grytsai and Andrey Gulin prepared the Figures; AL and A.Grytsai took part in the writing initial version of the
763 paper and the preparation of the revised version.

764 VF, AR, MS took part in the analysis of the data and preparation the data on VLF propagation in the ionosphere.

765 OA. JB and GM took part in developing the concept of the paper, and writing the initial version of the paper. OA, GM, VF,
766 AG contributed to the preparation of the revised version of the paper.
767 All participants took part in the analysis of the results.

768

769 COMPETING INTERESTS

770 The authors declare that they have no conflict of interest.

771

772 DATA AVAILABILITY

773 The VLF-LF data (Fig. 11) are property of IPE and TUoS groups, and they are not publicly accessible. According to an
774 agreement between all the participants, we cannot make the data openly accessible. Data can be provided under commercial
775 conditions via direct request to rozhnoi@ifz.ru. The ionospheric data used for the modeling the electrodynamics
776 characteristics of the VLF waves in the ionosphere are shown in part in Fig. 5 (namely, altitude distribution of electron
777 concentration). The other data necessary for determination the components of tensor of dielectric permittivity and then the
778 electrodynamics modeling in the accepted simple approximation of the three-component plasma-like ionosphere (including
779 electron, one effective ion and one effective neutral components) and quasi-neutrality are mentioned in Sect. 3.1. The
780 corresponding ionospheric data have been taken from the well-known published handbooks, referred in the paper (Al'pert
781 1972; Alperovich and Fedorov 2007; Kelley 2005; Schunk and Nagy 2010; Jursa 1985).

782

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790 REFERENCES

791 Agapitov, O. V., Artemyev, A. V., Mourenas, D., Kasahara, Y., and Krasnoselskikh, V.: Inner belt and slot region electron
792 lifetimes and energization rates based on AKEBONO statistics of whistler waves, *J. Geophys. Res.: Space Phys.*, 119, 2876–
793 2893, <https://doi.org/10.1002/2014JA019886>, 2014.

794 Agapitov, O. V., Mourenas, D., Artemyev, A. V., Mozer, F. S., Hospodarsky, G., Bonnell, J., and Krasnoselskikh, V.:
795 Synthetic Empirical Chorus Wave Model From Combined Van Allen Probes and Cluster Statistics, *J. Geophys. Res.: Space
796 Phys.*, 123, 2017JA024843, <https://doi.org/10.1002/2017JA024843>, 2018.

797 Alperovich, L. S. and Fedorov, E. N.: *Hydromagnetic Waves in the Magnetosphere and the Ionosphere*, Springer, 426 pp.,
798 2007.

798 Al'pert, Ya.L: *Propagation of Electromagnetic Waves and the Ionosphere* (in Russian), Moscow, Nauka Publ., 1972, 563 pp

799 Arnoldy, R. L. and Kintner, P. M.: Rocket observations of the precipitation of electrons by ground VLF transmitters, *J.*
800 *Geophys. Res.: Space Phys.*, 94, 6825–6832, <https://doi.org/10.1029/JA094iA06p06825>, 1989.

801 Artsimovich, L. A. and Sagdeev, R. Z: Physics of plasma for physicists, ATOMIZDAT Publ. House, Moscow, 320 pp., 1979
802 (in Russian).

803 Artemyev, A. V., Agapitov, O. V., Mourenas, D., Krasnoselskikh, V., and Zelenyi, L. M.: Storm-induced energization of
804 radiation belt electrons: Effect of wave obliquity, *Geophys. Res. Lett.*, 40, 4138–4143, <https://doi.org/10.1002/grl.50837>,
805 2013.

806 Artemyev, A., Agapitov, O., Mourenas, D., Krasnoselskikh, V., and Mozer, F.: Wave energy budget analysis in the Earth's
807 radiation belts uncovers a missing energy, *Nat. Commun.*, 6, <https://doi.org/10.1038/ncomms8143>, 2015.

808 Azadifar, M., Li, D., Rachidi, F., Rubinstein, M., Diendorfer, G., Schulz, W., Pichler, H., Rakov, V. A., Paolone, M.,
809 Pavanello, D.: Analysis of lightning-ionosphere interaction using simultaneous records of source current and 380-km distant
810 electric field, *Journ. Atmos. Solar-Terrestr. Phys.*, 159, 48–56, <https://doi.org/10.1016/j.jastp.2017.05.010>, 2017.

811 Beletskii, N. N., Borysenko, S. A., and Gvozdev, N. I.: The Resonant Interaction of Electromagnetic Waves in a Defect
812 Dielectric Periodic Layered Structure Placed in a Parallel-Plate Waveguide, *Radiophys. Electron.*, 5, 7 pp.,
813 <https://doi.org/10.1615/TelecomRadEng.v75.i6.40>, 2014.

814 Biagi, P. F., Maggipinto, T., and Ermini, A.: The European VLF/LF radio network: current status, *Acta Geod. Geophys.*, 50,
815 109–120, <https://doi.org/10.1007/s40328-014-0089-x>, 2015.

816 Biagi, P. F., Maggipinto, T., Righetti, F., Loiacono, D., Schiavulli, L., Ligonzo, T., Ermini, A., Moldovan, I. A., Moldovan,
817 A. S., Buyuk sarac, A., Silva, H. G., Bezzeghoud ,M., and Contadakis, M. E.: The European VLF/LF radio network to search
818 for earthquake precursors: setting up and natural/man-made disturbances, *Nat. Hazards Earth Syst. Sci.*, 11, 333–341,
819 <https://doi.org/10.5194/nhess-11-333-2011>, 2011.

820 Boudjada, M. Y., Schwingenschuh, K., Al-Haddad, E., Parrot, M., Galopeau, P. H. M., Besser, B., Stangl, G., and Voller,
821 W., Effects of solar and geomagnetic activities on the sub-ionospheric very low frequency transmitter signals received by the
822 DEMETER micro-satellite, *Ann. Geo.*, 55, 1, <https://doi.org/10.4401/ag-5463>, 2012.

823 Bragin, Yu. A., Tyutin Alexander, A., Kocheev, A.A., and Tyutin, Alexei, A.: Direct measurements of the vertical electric
824 field of the atmosphere up to 80 km, *Space Res.*, 12, 279–281, 1974.

825 Brittain, R., Kintner, P. M., Kelley, M. C., Siren, J. C., and Carpenter, D. L.: Standing wave patterns in VLF Hiss, J.
826 *Geophys. Res.: Space Phys.*, 88, 7059–7064, <https://doi.org/10.1029/JA088iA09p07059>, 1983.

827 Buitink, S., Corstanje, A., Enriquez, J. E., Falcke, H., Hörandel ,J. R., Huege, T., Nelles, A., Rachen, J. P., Schellart, P.,
828 Scholten, O., ter Veen, S., Thoudam, S., and Trinh, T. N. G.: A method for high precision reconstruction of air shower X_{\max}
829 using two-dimensional radio intensity profiles, *arXiv:1408.7001v2 [astro-ph.IM]*, 1 Sep 2014.

830 Chevalier, M. W., Inan U. S.: A Technique for Efficiently Modeling Long-Path Propagation for Use in Both FDFD and
831 FDTD, *IEEE Ant. Wireless Propag. Lett.*, 5, 525-528, <https://doi.org/10.1109/LAWP.2006.887551>, 2006

832 Chou, M.-Y., Lin, C. C. H., Yue, J., Chang, L. C., Tsai, H.-F., and Chen, C.-H., Medium-scale traveling ionospheric
833 disturbances triggered by Super Typhoon Nepartak, *Geophys. Res. Lett.*, 44, 7569–7577,
834 <https://doi.org/10.1002/2017gl073961>, 2017.

835 Collin, R. E., *Foundations for Microwave Engineering*. New York: John Wiley & Sons, Inc., 2001.

836 Cummer, S. A., Inan ,U. S., Bell, T. F., and Barrington-Leigh, C. P.: ELF radiation produced by electrical currents in sprites,
837 *Geophys. Res. Lett.*, 25, 1281–1284, <https://doi.org/10.1029/98GL50937>, 1998.

838 Cummer, S. A., Briggs, M. S., Dwyer, J. R., Xiong, S., Connaughton, V., Fishman, G. J., Lu, G., Lyu, F., and Solanki, R.:
839 The source altitude, electric current, and intrinsic brightness of terrestrial gamma ray flashes, *Geophys. Res. Lett.*, 41,
840 <https://doi.org/10.1002/2014GL062196>, 2014.

841 Dwyer, J. R. and Uman, M. A.: The physics of lightning, *Phys. Rep.*, 534, 147–241,
842 <https://doi.org/10.1016/j.physrep.2013.09.004>, 2014.

843 Dwyer, J. R.: The relativistic feedback discharge model of terrestrial gamma ray flashes, *Journ. Geophys. Res.*, 117,
844 A02308, <https://doi.org/10.1029/2011JA017160>, 2012

845 Glukhov, V. S., Pasko, V. P., & Inan, U. S. (1992). Relaxation of transient lower ionospheric disturbances caused by
846 lightning-whistler-induced electron precipitation bursts. *Journal of Geophysical Research*, 97(A11),
847 16971. doi:10.1029/92ja01596

848 Grimalsky, V.V., Kremenetsky, I.A., Rapoport, Yu.G.: Excitation of electromagnetic waves in the lithosphere and their
849 penetration into ionosphere and magnetosphere, *J. of Atmosph. Electricity*, 19, 101-117, 1999 a.

850 Grimalsky, V.V., Kremenetsky, I.A., Rapoport, Yu.G.: Excitation of EMW in the lithosphere and propagation into
851 magnetosphere . In: “Atmospheric and Ionospheric Electromagnetic Phenomena Associated with Earthquakes”. Ed.: M.
852 Hayakawa. (TERRAPUB), Tokyo, 777-787, 1999 b.

853 [Guglielmi A.V., Pokhotelov O.A.: *Geoelectromagnetic waves*. IOP Publ.: Bristol, 1996, 402 P.](#)

854 Gurevich, A V, Zybin, K P.: Runaway breakdown and electric discharges in thunderstorms, *Physics – Uspekhi*, 44, 11, 1119
855 – 1140, <https://doi.org/10.1070/PU2001v044n11ABEH000939>, 2001

856 Gurevich, A V., Karashtin, A N., Ryabov, V A., Chubenko, A L Shepetov: Nonlinear phenomena in the ionospheric
857 plasma. Effects of cosmic rays and runaway breakdown on thunderstorm discharges, *Physics - Uspekhi* 52 (7), 735-745,
858 <https://doi.org/0.3367/UFNe.0179.200907h.0779>, 2009

859 Hapgood, M.: Space Weather, IOP Publ., Bristol, 23 PP., <https://doi.org/10.1088/978-0-7503-1372-8>, 2017

860 Hare, B. M., Scholten, O., Bonardi, A., Buitink, S., Corstanje, A., Ebert, U., Falcke, H., Hörandel, J. R., Leijnse, H., Mitra,
861 P., Mulrey, K., Nelles, A., Rachen, J. P., Rossetto, L., Rutjes, C., Schellart, P., Thoudam, S., Trinh, T. N. G., ter Veen, S.,
862 and Winchen, T., LOFAR Lightning Imaging: Mapping Lightning With Nanosecond Precision, *J. Geophys. Res.:
863 Atmospheres*, 123, 2861–2876, <https://doi.org/10.1002/2017JD028132>, 2018.

864 Hayakawa, M.: Earthquake prediction with radio techniques, Wiley, Singapore, 294 P., 2015.

865 Horne, R. B., Thorne, R. M., Shprits, Y. Y., Meredith, N. P., Glauert, S. A., Smith, A. J., Kanekal, S. G., Baker, D. N.,
866 Engebretson, M. J., Posch, J. L., Spasojevic, M., Inan, U. S., Pickett, J. S., and Decreau, P. M. E.: Wave acceleration of
867 electrons in the Van Allen radiation belts. *Nature*, 437, 227–230, <https://doi.org/10.1038/nature03939>, 2005

868 **Jursa A.S., Ed.: Handbook of Geophysics and the Space Environment Air Force Geophysics Lab., 1042 P, 1985**

869 **Kelley, M. C.: The Earth's Ionosphere Plasma Physics and Electrodynamics, Acad. Press, Amsterdam, 556 P., 2009**

870 Kim, S. and Kim, K.: Invariant imbedding theory of wave propagation in arbitrarily inhomogeneous stratified bi-isotropic
871 media, *J. Opt.* 18, 065605 (9pp), <https://doi.org/10.1088/2040-8978/18/6/065605>, 2016.

872 Kintner, P. M., Brittain, R., Kelley, M. C., Carpenter, D. L., and Rycroft, M. J., In situ measurements of transionospheric
873 VLF wave injection, *J. Geophys. Res.: Space Phys.*, 88, 7065–7073, <https://doi.org/10.1029/JA088iA09p07065>, 1983.

874 Kong, J. A., *Electromagnetic Wave Theory*, New York: John Wiley & Sons, Inc., 1990.

875 Koskinen, H. E. J., *Physics of Space Storms. From the Solar Surface to the Earth*, Springer-Verlag, 419 P., 2011.

876 Kurushin, E. P. and Nefedov E. I.: *Electrodynamics of anisotropic waveguiding structures*, Moscow: Nauka (Science), 1983
877 (in Russian).

878 Kuzichev, I.V., Shklyar, D.R.: On full-wave solution for VLF waves in the near-Earth space, *Journ. Atmos. Solar-Terrestr.*
879 *Phys.*, 72, 1044–1056, <https://doi.org/10.1016/j.jastp.2018.07.002>, 2010

880 Kuzichev, I.V., Vasko, I. Yu., Malykhin, A. Yu., Soto-Chavez, A. R.: On the ionospheric propagation of VLF waves
881 generated by currents in the lower ionosphere, *Journ. Atmos. Solar-Terrestr. Phys.*, 179, 138-148,
882 <https://doi.org/10.1016/j.jastp.2018.07.002>, 2018

883 Lehtinen, N. G., Inan U. S.: Radiation of ELF/VLF waves by harmonically varying currents into a stratified ionosphere with
884 application to radiation by a modulated electrojet, *Journ. Geophys. Res.*, 113, A06301,
885 <https://doi.org/10.1029/2007JA012911>, 2008

886 Lehtinen, N. G., Inan U. S.: Full-wave modeling of transionospheric propagation of VLF waves, *Geophys. Res. Lett.*, 36,
887 L03104, <https://doi.org/10.1029/2008GL036535>, 2009

888 Levy, M.: Parabolic equation methods for electromagnetic wave propagation, *The Inst. of Electrical Eng., Padstow,*
889 Cornwall, 336 P., 2000

890 Lu, G., Zhang, H., Cummer, S. A., Wang, Y., Lyu, F., Briggs, M., Xiong, S., Chen, A.: A comparative study on the lightning
891 sferics associated with terrestrial gamma-ray flashes observed in Americas and Asia, *Journ. Atmos. Solar-Terrestr. Phys.*
892 183, 67–75, <https://doi.org/10.1016/j.jastp.2019.01.001>, 2019

893 Maier, S. A.: *Plasmonics: Fundamentals and Applications*, N. Y.: Springer, 234 p., 2007.

894 Marshall, R. A., Wallace, T.: Finite-Difference Modeling of Very-Low-Frequency Propagation in the Earth-Ionosphere
895 Waveguide, *IEEE Trans. Ant. Propag.*, 65, 7185-7197, <https://doi.org/10.1109/TAP.2017.2758392>, 2017

896 Martynenko, S. I., Rozumenko, V. T., and Tyrnov, O. F.: New possibilities for mesospheric electricity diagnostics, *Adv.*
897 *Space Res.*, 27, 1127–1132, [https://doi.org/10.1016/S0273-1177\(01\)00208-3](https://doi.org/10.1016/S0273-1177(01)00208-3), 2001.

898 Meek, C. E., Manson, A. H., Martynenko, S. I., Rozumenko, V. T., and Tyrnov, O. F.: Remote sensing of mesospheric
899 electric fields using MF radars, *J. Atmos. Sol.-Terr. Phys.*, 66, 881–890, <https://doi.org/10.1016/j.jastp.2004.02.002>, 2004.

900 Mezentsev, A., Lehtinen, N., Ostgaard, N., Perez-Invernón, F. J., and Cummer, S. A.: Spectral characteristics of VLF sferics
901 associated with RHESSI TGFs, *J. Geophys. Res.: Atmospheres*, 123, <https://doi.org/10.1002/2017JD027624>, 2018.

902 Nina, A., Radovanović, M., Milovanović, B., Kovačević, A., Bajčetić, J., and Popović, L. C.: Low ionospheric reactions on
903 tropical depressions prior hurricanes, *Adv. Space Res.*, 60, 1866–1877, <https://doi.org/10.1016/j.asr.2017.05.024>, 2017.

904 Qin, J., Celestin, S., and Pasko, V. P.: Low frequency electromagnetic radiation from sprite streamers, *Geophys. Res. Lett.*,
905 39, L22803, <https://doi.org/10.1029/2012GL053991>, 2012.

906 Patra, S., Spencer, E. Horton, W. and Sojka, J.: Study of Dst/ring current recovery times using the WINDMI model, *Journ.
907 Geophys. Res.*, 116, A02212, <https://doi.org/10.1029/2010JA015824>, 2011

908 Pulinets S and Boyarchuk K (2005) Ionospheric Precursors of Earthquakes, Berlin, Springer, 315P

909 Rapoport, Yu. G., Boardman, A. D., Grimalsky, V. V., Ivchenko, V. M., Kalinich, N: Strong nonlinear focusing of light in
910 nonlinearly controlled electromagnetic active metamaterial field concentrators, *J. Opt. (United Kingdom)*, 16, 0552029–
911 0552038, <https://doi.org/10.1088/2040-8978/16/5/055202>, 2014.

912 Rapoport, Yu. G., Gotnyan, O. E., Ivchenko, V. N., Hayakawa, M., Grimalsky, V. V., Koshevaya, S. V., Juarez, D.:
913 Modeling electrostatic – photochemistry seismoionospheric coupling in the presence of external currents, *Phys. Chem. Earth*,
914 31, 4-9, 437–446, <https://doi.org/10.1016/j.pce.2006.02.010>, 2006

915 Richmond, A. D.: Space Weather Research Prompts Study of Ionosphere and Upper Atmospheric Electrodynamics, *EOS, Trans. , Amer. Geophys. Union*, 77,(11), 101, <https://doi.org/10.1029/96eo00066>, 1996

916 Rozhnoi, A., Solovieva, M., Levin, B., Hayakawa, M., and Fedun, V.: Meteorological effects in the lower ionosphere as
917 based on VLF/LF signal observations, *Nat. Hazards Earth Syst. Sci.*, 14, 2671–2679, <https://doi.org/10.5194/nhess-14-2671-2014>, 2014.

918 Rozhnoi, A., Solovieva, M., Parrot, M., Hayakawa, M., Biagi, P.-F., Schwingenschuh, K., Fedun, V.: VLF/LF signal studies
919 of the ionospheric response to strong seismic activity in the Far Eastern region combining the DEMETER and ground-based
920 observations, *Phys. Chem. Earth*, 85–86, 141–149, <https://doi.org/10.1016/j.pce.2015.02.005>, 2015.

921 **Ruibys, G., & Tolutis, R: Nonreciprocal HF signal transmission by surface helicon. Electronics Letters, 19(8), 273.** <https://doi.org/10.1049/el:19830191>, 1983

922 Samarskii, A.A.: The Theory of Difference Schemes, Marcel Dekker, N.Y., 2001.

923 Sanchez-Dulcet, F., Rodriguez-Bouza M., Silva, H. G., Herraiz, M., Bezzeghoud, M., Biagi, P. F.: Analysis of observations
924 backing up the existence of VLF and ionospheric TEC anomalies before the Mw6.1 earthquake in Greece, January 26, 2014,
925 *Phys. Chem. Earth*, 85–86, 150–166, <https://doi.org/10.1016/j.pce.2015.07.002>, 2015.

926 Scholten, O., Bonardi, A., Buitink, S., Corstanje, A., Ebert, U., Falcke, H., Hörandel, J., Mitra, P., Mulrey, K., Nelles, A.,
927 Rachen, J., Rossetto, L., Rutjes, C., Schellart, P., Thoudam, S., Trinh, G., ter Veen, S., and Winchen, T. Precision study of

931 radio emission from air showers at LOFAR, EPJ Web of Conferences, 136, 02012,
932 <https://doi.org/10.1051/epjconf/201713602012>, 2017.

933 Senior, T. B. A. and J. L. Volakis: Approximate boundary conditions in electromagnetics, London: Institution of Electrical
934 Engineers, 1995.

935 Shalashov, A. G. and Gospodchikov, E. D.: Impedance technique for modeling electromagnetic wave propagation in
936 anisotropic and gyrotrropic media, Physics-Uspekhi, 54, 145–165, <https://doi.org/10.3367/UFNe.0181.201102c.0151>, 2011.

937 Siefring, C. L. and Kelley, M. C.: Analysis of standing wave patterns in VLF transmitter signals: Effects of sporadic E layers
938 and in situ measurements of low electron densities, J. Geophys. Res.: Space Phys., 96, 17813–17826,
939 <https://doi.org/10.1029/91JA00615>, 1991.

940 Shinagawa, H., Tsugawa, T., Matsumura, M., Iyemori, T., Saito, A., Maruyama, T., ... Otsuka, Y. (2013). Two-dimensional
941 simulation of ionospheric variations in the vicinity of the epicenter of the Tohoku-oki earthquake on 11 March 2011.
942 *Geophysical Research Letters*, 40(19), 5009–5013. doi:10.1002/2013gl057627

943 Schunk, R.W. and Nagy, A.F.: *Ionospheres. Physics, Plasma Physics, and Chemistry*, Cambridge, CUP, 628 P., 2010

944 Spiegel, M. R.: Theory and problems of vector analysis and an introduction to tensor analysis, Shnauim Publ., NY, 223 P.,
945 1959.

946 Surkov, V. and Hayakawa, M.: Ultra and Extremely Low Frequency Electromagnetic Fields, Springer, Tokyo, 2014.

947 Tarkhanyan, R. H. and Uzunoglu, N. K., Radiowaves and Polaritons in Anisotropic Media, Weinheim: Wiley-VCH, 210 p.,
948 2006.

949 Thorne, R. M.: Radiation belt dynamics: The importance of wave-particle interactions, Geophys. Res. Lett., 372, 22107,
950 <https://doi.org/10.1029/2010GL044990>, 2010.

951 Tolstoy, A., Rosenberg, T. J., Inan, U. S., & Carpenter, D. L. (1986). Model predictions of subionospheric VLF signal
952 perturbations resulting from localized, electron precipitation-induced ionization enhancement regions. *Journal of*
953 *Geophysical Research*, 91(A12), 13473.doi:10.1029/ja091ia12p13473

954 Tretyakov, S.: Analytical Modeling in Applied Electromagnetics, Artech House, Boston, 2003.

955 Walker, A.D. M.: The Theory of Whistler Propagation, Rev. Geophys. Space Phys., 14, 629-638, 1976

956 Weiland, J. and Wilhelmsson, H.: Coherent Non-Linear Interaction of Waves in Plasmas, Pergamon, London, 1977.

957 Wait J. R.: Electromagnetic waves in stratified media, New York: IEEE Press and Oxford Univ. Press, 1996.

958 Wu, C.-C., Liou, K., Lepping, R. P., Huttig, L., Plunkett, S., Howard, R. A., Socker, D.: The first super geomagnetic storm
959 of solar cycle 24: "The St. Patrick's day event (17 March 2015)", Earth, Planets and Space, 68, 151,
960 <https://doi.org/10.1186/s40623-016-0525-y>, 2016.

961 Yiğit, E., Koucká Knížová, P., Georgieva, K., & Ward, W.: A review of vertical coupling in the Atmosphere–Ionosphere
962 system: Effects of waves, sudden stratospheric warmings, space weather, and of solar activity, *Solar-Terrestrial Physics*, 141,
963 1–12, <https://doi.org/10.1016/j.jastp.2016.02.011>, 2016

964 Yu, Ya., Niu, J., Simpson, J. J.: A 3-D Global Earth-Ionosphere FDTD Model Including an Anisotropic Magnetized

Appendix: the matrix coefficients included into eq. (16)

Here the expressions of the matrix coefficients are presented that are used in the matrix factorization to compute the tensor
969 impedance, see eq. (16).

$$\hat{\alpha}_N^{(0)} = \begin{pmatrix} 1 + \frac{ih_z}{\Delta} (k_1 - \alpha_1 \alpha_2 k_2); & \frac{ih_z}{\Delta} \alpha_2 (k_2 - k_1) \\ \frac{ih_z}{\Delta} \alpha_1 (k_1 - k_2); & 1 + \frac{ih_z}{\Delta} (k_2 - \alpha_1 \alpha_2 k_1) \end{pmatrix}, \quad \hat{\alpha}_N^{(-)} = \begin{pmatrix} -1; & 0 \\ 0; & -1 \end{pmatrix}; \quad \Delta \equiv 1 - \alpha_1 \alpha_2;$$

$$\hat{\alpha}_j^{(-)} = \left\{ \begin{array}{l} \left(\frac{\beta_{22}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \right)_{j-1/2}; \quad - \left(\frac{\beta_{21}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \right)_{j-1/2} + \frac{ik_x h_z}{2} \left(\frac{\beta_{23}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \right)_{j-1} \\ \left(\frac{\beta_{12}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \right)_{j-1/2} + \frac{ik_x h_z}{2} \left(\frac{\beta_{32}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \right)_{j-1}; \quad \left(\beta_{11} + \frac{k_x^2}{k_0^2} \frac{\beta_{12} \cdot \beta_{21}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \right)_{j-1/2} - \frac{ik_x h_z}{2} \left(\beta_{13} + \frac{k_x^2}{k_0^2} \frac{\beta_{12} \cdot \beta_{23}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \right)_{j-1} - \\ \quad - \frac{ik_x h_z}{2} \left(\beta_{31} + \frac{k_x^2}{k_0^2} \frac{\beta_{32} \cdot \beta_{21}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \right)_j \end{array} \right\}$$

$$\hat{\alpha}_j^{(+)} = \left\{ \begin{array}{l} \left(\frac{\beta_{22}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \right)_{j+1/2}; \quad - \left(\frac{\beta_{21}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \right)_{j+1/2} - \frac{ik_x h_z}{2} \left(\frac{\beta_{23}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \right)_{j+1} \\ \left(\frac{\beta_{12}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \right)_{j+1/2} - \frac{ik_x h_z}{2} \left(\frac{\beta_{32}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \right)_{j+1}; \quad \left(\beta_{11} + \frac{k_x^2}{k_0^2} \frac{\beta_{12} \cdot \beta_{21}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \right)_{j+1/2} + \frac{ik_x h_z}{2} \left(\beta_{13} + \frac{k_x^2}{k_0^2} \frac{\beta_{12} \cdot \beta_{23}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \right)_{j+1} + \\ \quad + \frac{ik_x h_z}{2} \left(\beta_{31} + \frac{k_x^2}{k_0^2} \frac{\beta_{32} \cdot \beta_{21}}{1 - \beta_{22} \frac{k_x^2}{k_0^2}} \right)_j \end{array} \right\}$$

$$\hat{\alpha}_j^{(0)} = \begin{pmatrix} -\left(\frac{\beta_{22}}{1-\beta_{22}}\frac{k_x^2}{k_0^2}\right)_{j-1/2} - \left(\frac{\beta_{22}}{1-\beta_{22}}\frac{k_x^2}{k_0^2}\right)_{j+1/2} + k_0^2 h_z^2; & \left(\frac{\beta_{21}}{1-\beta_{22}}\frac{k_x^2}{k_0^2}\right)_{j-1/2} + \left(\frac{\beta_{21}}{1-\beta_{22}}\frac{k_x^2}{k_0^2}\right)_{j+1/2} \\ \left(\frac{\beta_{12}}{1-\beta_{22}}\frac{k_x^2}{k_0^2}\right)_{j-1/2} + \left(\frac{\beta_{12}}{1-\beta_{22}}\frac{k_x^2}{k_0^2}\right)_{j+1/2}; & -\left(\beta_{11} + \frac{k_x^2}{k_0^2}\frac{\beta_{12} \cdot \beta_{21}}{1-\beta_{22}}\frac{k_x^2}{k_0^2}\right)_{j-1/2} - \left(\beta_{11} + \frac{k_x^2}{k_0^2}\frac{\beta_{12} \cdot \beta_{21}}{1-\beta_{22}}\frac{k_x^2}{k_0^2}\right)_{j+1/2} + \\ & + k_0^2 h_z^2 \cdot \left(1 - \beta_{33} \frac{k_x^2}{k_0^2} - \frac{k_x^4}{k_0^4} \frac{\beta_{23} \cdot \beta_{32}}{1-\beta_{22}}\frac{k_x^2}{k_0^2}\right)_j \end{pmatrix}.$$

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