



## Monitoring potential ionospheric changes caused by Van earthquake (Mw 7.2)

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### ABSTRACT

Many scientists from different disciplines have studied earthquakes for many years. As a result of these studies, it has been proposed that some changes take place in the ionosphere layer before, during or after earthquakes, and the ionosphere should be monitored in earthquake prediction studies. This study investigates the changes in the ionosphere created by the earthquake with magnitude of Mw=7.2 in the northwest of the Lake Erçek which is located to the north of the province of Van in Turkey on 23 October 2011 and at 1.41 pm local time (-3 UT) with the epicenter of 38.75° N, 43.36° E using the TEC values obtained by the Global Ionosphere Models (GIM) created by IONOLAB-TEC and CODE. In order to see whether the ionospheric changes obtained by the study in question were caused by the earthquake or not, the ionospheric conditions were studied by utilizing indices providing information on solar and geomagnetic activities (F10.7 cm, Kp, Dst).

One of the results of the statistical test on the TEC values obtained from the both models, positive and negative anomalies were obtained for the times before, on the day of and after the earthquake, and the reasons for these anomalies are discussed in detail in the last section of the study. As the ionospheric conditions in the analyzed days were highly variable, it was thought that the anomalies were caused by geomagnetic effects, solar activity and the earthquake.

**Keywords:** TEC, Van Earthquake, Ionosphere

### 1. INTRODUCTION



35 The ionosphere is the part of the atmosphere at the altitudes of 60 km to 1,100 km where there  
 36 are ions and free electrons in considerable amounts that can reflect electromagnetic waves. It  
 37 completely covers the thermosphere, one of the main layers of the atmosphere, but also includes  
 38 some of the mesosphere and the exosphere.

39 The most important parameter that defines the ionosphere in space and time is the quantity of  
 40 electrons. This quantity varies under the influence of the day-night cycle, seasons, geographical  
 41 location and magnetic storms in the sun. As it is not possible to measure the quantity of electrons  
 42 in the ionosphere directly, indirect measurement and calculation methods have been developed  
 43 (Li and Parrot, 2018). Total Electron Content (TEC), which is defined as the quantity of free  
 44 electrons along a cylinder with a cross section of  $1 \text{ m}^2$ , is a suitable parameter to monitor the  
 45 changes in the ionosphere in space and time. All signals that contain audio and data that pass  
 46 through or get reflected from the ionosphere, which is highly irregular and difficult to model,  
 47 are affected by the structure of this layer.

48 Calculating Total Electron Content (TEC) is a method used directly to investigate the structure  
 49 of the ionosphere. TEC is represented by the unit of TECU, and one TECU equals to  
 50  $10^{16} \text{ el/m}^2$  (Schaer, 1999). TEC is expressed in two ways: STEC (Slant Total Electron  
 51 Content); the free electron content calculated along the slanted line between the receiver and  
 52 the satellite, and VTEC (Vertical Total Electron Content); the free electron content calculated  
 53 along the zenith of the receiver (Langley, 2002).

54 TEC varies based on positional and temporal variables such as the latitude of the place, seasons,  
 55 solar activity, geomagnetic storms and earthquakes. Ionospheric altitude also differs based on  
 56 geography.

57 TEC, which is defined as the number of free electrons on the one square meter area on the line  
 58 followed by a radio wave, is one of the important parameters for examining the structure of the  
 59 ionosphere and the upper atmosphere. With TEC values, it is possible to examine the short and  
 60 long-term changes in the ionosphere, ionospheric irregularities and disruptive factors together  
 61 (Erol and Arıkan 2005).

62 Ionospheric changes are being studied in more than twenty countries today as precursors of  
 63 earthquakes. Definition of ionospheric anomalies and feasibility studies of seismo-ionospheric  
 64 precursors are still ongoing (Akhoondzadeh et al., 2018; Liu et al., 2010; He et al., 2012;  
 65 Kamogawa and Kakinami, 2013; Heki and Enomoto, 2015; Pulinets and Davidenko, 2014;  
 66 Masci et al., 2015; Yildirim et al., 2016; He and Heki, 2017; Kelley et al., 2017; Rozhnoi et al.,  
 67 2015; Thomas et al., 2017).



## 68 2. METHODOLOGY

### 69 2.1 IONOLAB-TEC Method:

70 The IONOLAB-TEC method developed by the department of Electrical and Electronics  
 71 Engineering of Hacettepe University is a JAVA application that uses the Regularized TEC (D-  
 72 TEI) algorithm (Arikan et al. 2004 ).

73 In this application, they developed a method that estimates VTEC values by using all GPS  
 74 signals measured at a period of time in a day. While the measurements taken from the satellites  
 75 with elevations of  $60^\circ$  or higher are used, the measurements from the satellites with elevations  
 76 of  $10^\circ$  to  $60^\circ$  are weighted by a Gauss function. The data from satellites with elevations of  
 77 lower than  $10^\circ$  are not included in calculations to reduce multipath effects. In this method raw  
 78 GPS data was used to determine VTEC value.

### 79 2.2 Global Ionosphere Model (GIM): 80

81 Global Ionospheric Maps are published in the IONEX (IONosphere map EXchange) format in  
 82 a way that covers the entire world. The institutions that produce these maps in the world include  
 83 CODE (Center for Orbit Determination in Europe, Switzerland), DLR (FERNKUNDUNGSTATION  
 84 Neustrelitz, Germany), ESOC (European Space Operations Centre, Germany), JPL (Jet  
 85 Propulsion Laboratory, California), NOAA (National Oceanic and Atmospheric  
 86 Administration, United States), NRCAN (National Resources, Canada), ROB (Royal  
 87 Observatory of Belgium, Belgium), UNB (University of New Brunswick, Canada), UPC  
 88 (Polytechnic University of Catalonia, Spain), WUT (Warsaw University of Technology,  
 89 Poland). In this study we used the GIM-TEC values produced by CODE in the IONEX format.  
 90 In the dates they were analyzed, the temporal resolution of the TEC values was 2 hours, while  
 91 their positional resolution was  $2.5^\circ$  by latitude and  $5^\circ$  by longitude. In order to calculate TEC  
 92 values for a point whose latitude and longitude is known on the GIM-TEC maps created by  
 93 CODE using more than 300 GNSS receivers around the world, the 4 TEC values that cover the  
 94 point and the two-variable interpolation formula are given below.

$$96 E_{int}(\lambda_0 + p\Delta\lambda, \beta_0 + q\Delta\beta) = (1-p)(1-q)E_{0,0} + p(1-q)E_{1,0} + q(1-p)E_{0,1} + pqE_{1,1} \quad (1)$$

97  $p$  and  $q$ :  $0 \leq p, q < 1$ .

98  $\Delta\lambda$  and  $\Delta\beta$ : Longitude and Latitude differences grid widths,

99  $\lambda_0$  and  $\beta_0$ : Initial longitude and latitude values,

100  $E_{0,0}, E_{1,0}, E_{0,1}$  ve  $E_{1,1}$  : TEC values known in neighboring points,

101  $E_{int}$ : TEC value to be found.

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### 3. ANALYSIS TO DETERMINE EARTHQUAKE-RELATED TEC CHANGES

In order to investigate earthquake-related TEC changes, the TEC values for the stations close to the epicenters, HAKK, MALZ, OZAL and TVAN (TUSAGA-Aktive CORS-TR) GPS stations were analyzed to determine TEC value using the IONOLAB-TEC and GIM-TEC models. The correlation coefficient was obtained for the TEC values from both models between the dates 13.10.2011 and 02.11.2011 for the stations above.



**Figure 1.** Analyzed Stations

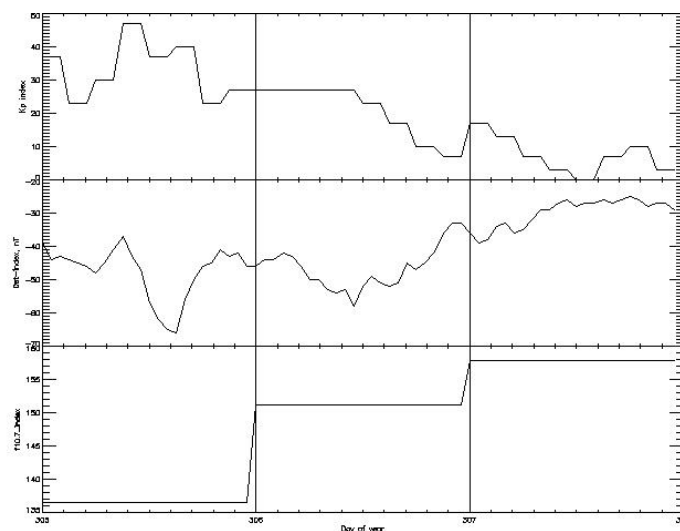
Figure 1 shows the stations analyzed (represented by red triangles) and the epicenter of the earthquake represented by blue star. For each station, the TEC values with the temporal resolution of two hours obtained from both the IONOLAB-TEC and GIM-TEC models and the correlation coefficient showing whether there is a linear relationship between two values were calculated as below;

$$r = \frac{\Sigma(xy) - (\Sigma x)(\Sigma y)/n}{\sqrt{(\Sigma x^2 - (\Sigma x)^2/n)(\Sigma y^2 - (\Sigma y)^2/n)}} \quad (2)$$

In order to determine the outlier values among the TEC values with a two-hour temporal resolution from both models, the TEC values obtained from both models between the dates 01.10.2011 and 10.10.2011, which were considered quiet in terms of geomagnetic and solar activity, were used to determine the upper boundary (UB) and the lower boundary (LB). By

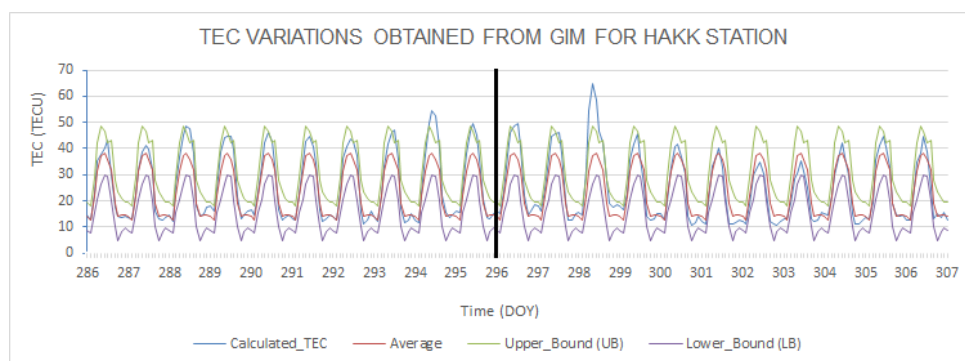


utilizing the TEC values from both models, the UB and LB values were calculated using the formulae  $x+3\sigma$  and  $x-3\sigma$ . Here,  $x$  is the mean TEC value for the relevant epoch and  $\sigma$  is the standard deviation. If the TEC value in any epoch is higher than the upper boundary, it is a positive anomaly. Similarly if it is lower than the lower boundary, it is a negative anomaly. In order to investigate whether the anomalies before, on the day of and after the earthquake were caused by the earthquake or not, we also examined the ( $Kp*10$ ), Dst and F10.7 cm indices, which provided information on the geomagnetic and solar activity for the days in which anomalies were detected.

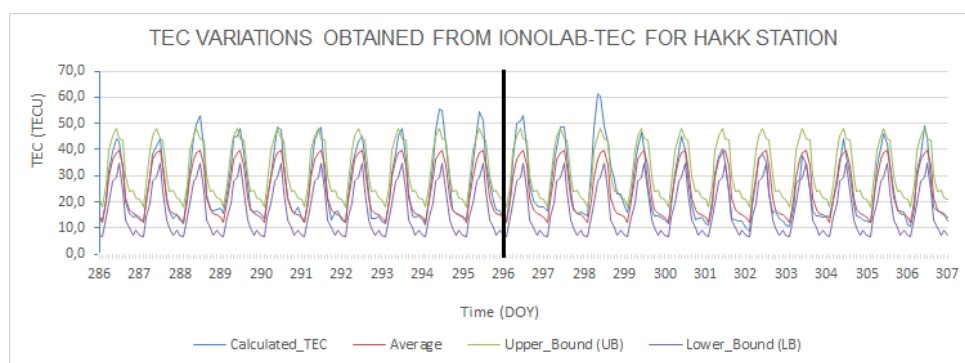


**Figure 2.** The Chart for the Dates 01-03.11.2011 in ( $Kp*10$ ), Dst and F10.7 cm Indices (URL-1)

Figures 2 shows that the ( $Kp*10$ ), Dst and F10.7 cm indices that provide information on geomagnetic and solar activity in October and on the first three days of November. Below are the TEC values for the HAKK station for the dates 13.10.2011-02.11.2011 obtained using the GIM-TEC and IONOLAB-TEC methods.



**Figure 3.** GIM-TEC Values for the HAKK Station



**Figure 4.** IONOLAB-TEC Values for the HAKK Station

The correlation coefficient  $r$  between the TEC values calculated by both methods for the HAKK station was 0.98 indicating a strong positive relationship. The anomaly tables for this station are provided below (Tables 1 and 2).

GIM-TEC Anomaly Table for HAKK Station									
Number	DOY	Hour	TEC Difference (TECU)	Type of Anomaly	Number	DOY	Hour	TEC Difference (TECU)	Type of Anomaly
1	286	12	1.0	Positive	7	294	12	10.5	Positive
2	288	12	5.7	Positive	8	295	12	7.3	Positive
3	289	12	2.5	Positive	9	296	12	7.5	Positive
4	290	12	0.5	Positive	10	297	12	4.1	Positive
5	292	12	0.8	Positive	11	298	8	16.5	Positive
6	293	12	5.2	Positive					

**Table 1.** HAKK Station Global Ionosphere Model Anomaly Table



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IONOLAB-TEC Anomaly Table for HAKK Station									
Number	DOY	Hour	TEC Difference (TECU)	Type of Anomaly	Number	DOY	Hour	TEC Difference (TECU)	Type of Anomaly
1	287	12	0.4	Positive	9	295	12	7.2	Positive
2	288	12	9.2	Positive	10	296	12	8.8	Positive
3	289	12	4.3	Positive	11	297	12	4.6	Positive
4	290	12	3.8	Positive	12	298	8	16.5	Positive
5	291	12	4.5	Positive	13	301	12	0.3	Negative
6	292	12	1.4	Positive	14	302	14	0.9	Negative
7	293	12	4.2	Positive	15	303	12	0.7	Negative
8	294	12	10.9	Positive	16	306	10	0.9	Positive

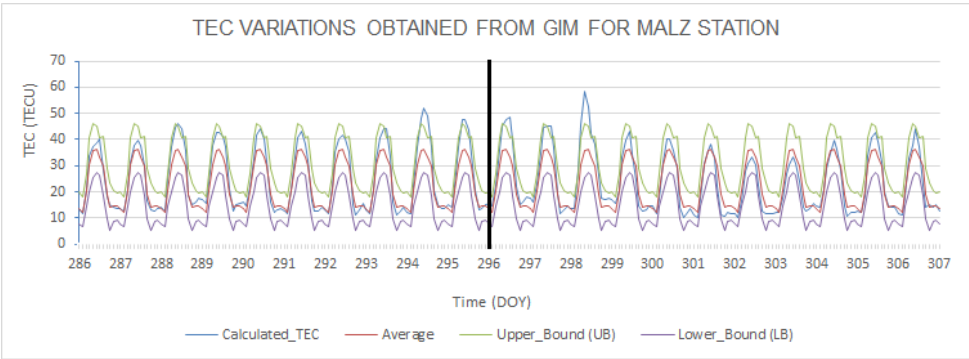
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**Table 2.** HAKK Station IONOLAB-TEC Anomaly Table

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156 Below are the TEC values for the MALZ station obtained using the GIM-TEC and IONOLAB-  
157 TEC methods (Figures 5 and 6).

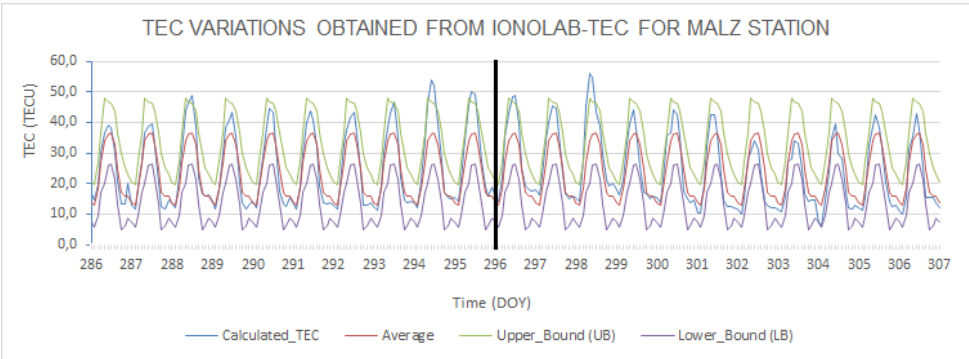
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**Figure 5.** GIM-TEC Values for the MALZ Station



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**Figure 6.** IONOLAB-TEC Values for the MALZ Station





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164 The correlation coefficient  $r$  between the TEC values calculated by both methods for the MALZ  
 165 station was 0.98 indicating also a strong positive relationship. The anomaly tables for this  
 166 station are provided below (Tables 3 and 4).

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GIM-TEC Anomaly Table for MALZ Station									
Number	DOY	Hour	TEC Difference (TECU)	Type of Anomaly	Number	DOY	Hour	TEC Difference (TECU)	Type of Anomaly
1	288	12	3.5	Positive	5	295	12	3.1	Positive
2	289	12	0.5	Positive	6	296	12	7.9	Positive
3	293	12	3.9	Positive	7	297	12	4.7	Positive
4	294	12	8.6	Positive	8	298	8	12.6	Positive

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**Table 3.** MALZ Station Global Ionosphere Model Anomaly Table

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IONOLAB-TEC Anomaly Table for MALZ Station									
Number	DOY	Hour	TEC Difference (TECU)	Type of Anomaly	Number	DOY	Hour	TEC Difference (TECU)	Type of Anomaly
1	288	12	2.3	Positive	5	296	12	2.5	Positive
2	293	12	0.4	Positive	6	298	6	8.6	Positive
3	294	10	7.4	Positive	7	304	0	0.2	Negative
4	295	10	3.6	Positive					

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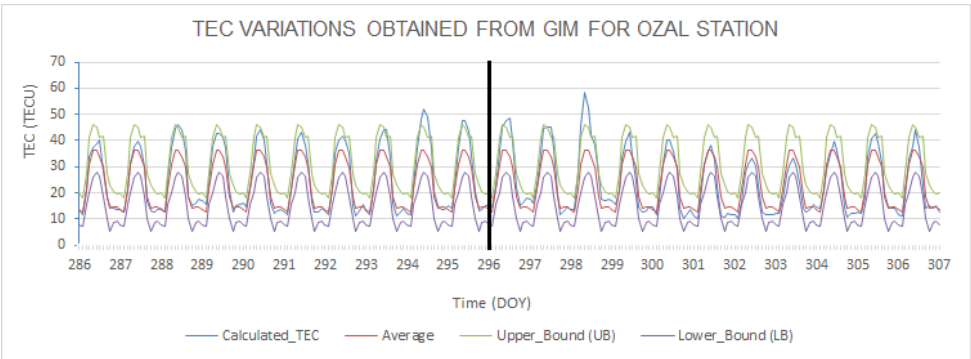
**Table 4.** MALZ Station IONOLAB-TEC Anomaly Table

172 Tables 3 and 4 show the anomalies found as a result of the analysis of the TEC values obtained  
 173 by the IONOLAB-TEC and GIM-TEC methods for the MALZ station. It is believed that the  
 174 positive anomaly on days 288 and 289 was caused by moderate (136.9 sfu, 150 sfu) solar  
 175 activity. It is also believed that the anomalies on the days 293, 294, 295 and 296 were caused  
 176 by strong (157.8 sfu, 166.3 sfu, 162.5 sfu, 153.9 sfu) solar activity.

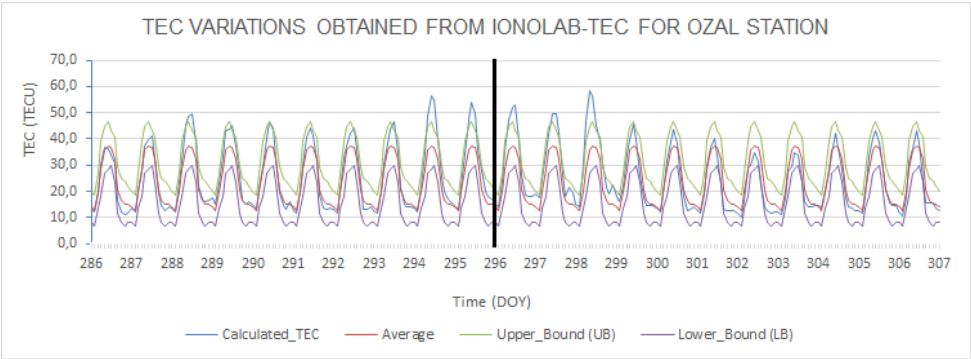
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178 Below are the TEC values for the OZAL station obtained using the GIM-TEC and IONOLAB-  
 179 TEC methods for the dates 13 October – 02 November (Figures 7 and 8).





**Figure 7.** GIM-TEC Values for the OZAL Station



**Figure 8** IONOLAB-TEC Values for the OZAL Station

The correlation coefficient  $r$  between the TEC values calculated by both methods for the OZAL station was 0.98 demonstrating a strong positive relationship. The anomaly tables for this station are provided below (Tables 5 and 6).

GIM-TEC Anomaly Table for OZAL Station									
Number	DOY	Hour	TEC Difference (TECU)	Type of Anomaly	Number	DOY	Hour	TEC Difference (TECU)	Type of Anomaly
1	288	12	2.8	Positive	5	296	12	7.2	Positive
2	293	12	3.2	Positive	6	297	12	4.0	Positive
3	294	12	7.9	Positive	7	298	8	12.4	Positive
4	295	12	2.4	Positive					

**Table 5.** OZAL Station Global Ionosphere Model Anomaly Table



IONOLAB-TEC Anomaly Table for OZAL Station									
Number	DOY	Hour	TEC Difference (TECU)	Type of Anomaly	Number	DOY	Hour	TEC Difference (TECU)	Type of Anomaly
1	288	12	6.1	Positive	7	295	10	7.4	Positive
2	289	12	1.6	Positive	8	296	12	9.6	Positive
3	290	12	0.9	Positive	9	297	12	6.0	Positive
4	293	12	3.5	Positive	10	298	8	13.6	Positive
5	292	12	0.6	Positive	11	301	14	1.2	Negative
6	294	12	11.8	Positive	12	302	14	1.4	Negative

Table 6. OZAL Station IONOLAB-TEC Anomaly Table

Below are the TEC values for the TVAN station obtained using the GIM-TEC and IONOLAB-TEC methods (Figures 9, 10).

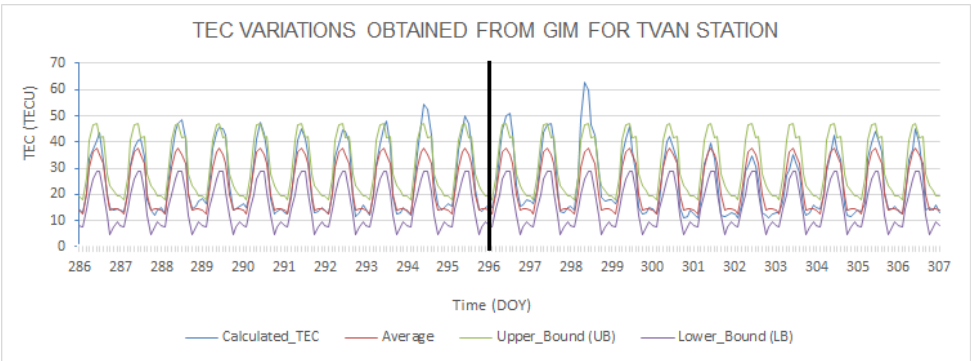


Figure 9. GIM-TEC Values for the TVAN Station

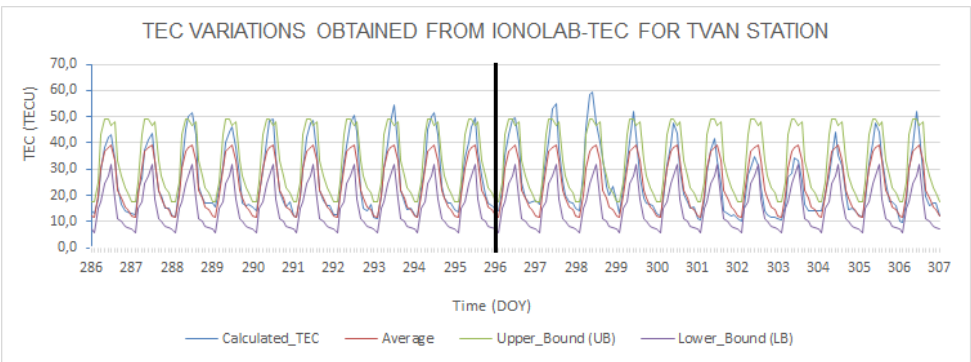


Figure 10. IONOLAB-TEC Values for the TVAN Station



The correlation coefficient between the TEC values calculated by both methods for the TVAN station was 0.98 representing a strong positive relationship. The anomaly tables for this station are provided below (Tables 7 and 8).

GIM-TEC Anomaly Table for TVAN Station									
Number	DOY	Hour	TEC Difference (TECU)	Type of Anomaly	Number	DOY	Hour	TEC Difference (TECU)	Type of Anomaly
1	286	12	2.1	Positive	10	294	12	11.0	Positive
2	288	12	7.0	Positive	11	295	12	5.4	Positive
3	289	12	3.5	Positive	12	296	12	9.3	Positive
4	290	12	1.8	Positive	13	297	12	5.5	Positive
5	292	12	2.8	Positive	14	298	8	16.5	Negative
6	293	12	6.4	Positive					

**Table 7.** TVAN Station Global Ionosphere Model Anomaly Table

IONOLAB-TEC Anomaly Table for TVAN Station									
Number	DOY	Hour	TEC Difference (TECU)	Type of Anomaly	Number	DOY	Hour	TEC Difference (TECU)	Type of Anomaly
1	288	12	5.1	Positive	10	296	12	3.4	Positive
2	290	12	2.6	Positive	11	297	12	8.5	Positive
3	291	12	2.0	Positive	12	298	10	10.5	Positive
4	292	12	4.0	Positive	13	299	10	2.8	Positive
5	293	12	8.1	Positive	14	302	12	0.7	Negative
6	294	12	5.1	Positive	15	306	10	2.9	Positive
7	295	12	3.2	Positive					

**Table 8.** TVAN Station IONOLAB-TEC Anomaly Table

Tables 1, 2, 3, 4, 5, 6, 7 and 8 show the results of the statistical analysis of the TEC values created by the IONOLAB-TEC and GIM-TEC methods. The tables also depict the day and hour in which anomalies were observed, and the quantity and type of the anomaly. The numbers of anomalies obtained in both models were very close to each other. The F10.7 cm index values between the days 286 and 292 were 136 sfu, 135.4 sfu, 136.9 sfu, 150 sfu, 151.6 sfu, 145.7 sfu, 146.1 sfu. The index values show that there was usually moderate solar activity. Therefore, the anomalies in question may be related to the earthquake or solar activity. The index values for the days 293, 294, 295 and 296 (the day of the earthquake) were 157.8 sfu, 166.3 sfu, 162.5 sfu and 153.9 sfu respectively. These values indicate strong solar activity. On the other hand, the ionosphere layer was quiet in these days in terms of geomagnetic conditions. As there was strong solar activity, the numbers of anomalies were higher than the numbers in the days 286-



220 292. Since solar activity was moderate in the day 297, the number of anomalies dropped. The  
 221 solar activity on the day 298 was moderate, but there was strong geomagnetic activity (Dst -  
 222 147 nt,  $K_p \cdot 10 = 73$ ). The reason for the high numbers of anomalies on day 298 in both models  
 223 is believed to be due to geomagnetic activity. Considering the analyzed days in general, it may  
 224 be seen that it is difficult to identify earthquake-related anomalies as the solar activity and  
 225 geomagnetic conditions before and after the earthquake were not quiet. Therefore, it is believed  
 226 that the anomalies detected in the stations on days 293-296 may be related to the earthquake  
 227 and/or solar activity, and the anomalies on days 297 and 298 may be related to the earthquake,  
 228 solar activity and/or geomagnetic activity.

#### 229 4. CONCLUSION

230 In the scope of this study, the TEC values for the stations HAKK, MALZ, OZAL, TVAN were  
 231 obtained using the GIM-TEC and IONOLAB-TEC methods. In the comparison of the obtained  
 232 values, it was seen that there was high correlation between the TEC values obtained by the two  
 233 models. In order to detect earthquake-related TEC changes better, the TEC values created from  
 234 both models for the period of 13.10.2011-02.11.2011 were used as reference to determine the  
 235 UB and LB values. As a result of the statistical test, anomalies were found in all analyzed  
 236 stations for before, on the day of and after the earthquake. In order to understand whether the  
 237 anomalies obtained in both models were earthquake-related, the ionospheric conditions,  
 238 geomagnetic activity and solar activity on the analyzed days were examined using the  $K_p$ , Dst  
 239 and F10.7 cm indices.

240 Consequently, it was determined that the positive anomalies observed on days 286-292 may be  
 241 related to moderate solar activity and/or the earthquake, and the positive anomalies observed  
 242 on days 293, 294, 295, 296 (day of the earthquake) may be related to strong solar activity and/or  
 243 the earthquake. Moderate solar activity and strong geomagnetic activity were observed for day  
 244 298, so the numbers of anomalies in both models increased dramatically. This increase is  
 245 considered to be related to geomagnetic activity. The anomaly on day 298 may be related to the  
 246 earthquake, geomagnetic effects and/or solar activity. The finding that the ionospheric  
 247 conditions were variable in the analyzed days makes it highly difficult to identify earthquake-  
 248 related ionospheric changes. Therefore, interdisciplinary studies are needed to determine the  
 249 earthquake-related part of the change in question.

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## 254 REFERENCES

- 255 Arikan, F., Erol, C. B., Arikan, O.: Regularized Estimation of Vertical Total Electron Content  
 256 from GPS Data for a Desired Time Period, *Radio Science*, 39:RS6012, 2004.
- 257 Erol, C.B., Arikan, F.: Statistical Characterization of the Ionosphere Using GPS Signals., *J. of*  
 258 *Electromagnetic Waves and Appl.*, Vol.19, No:3, 2005.
- 259 He, L., and Heki, K.: Ionospheric anomalies immediately before Mw 7.0-8.0  
 260 earthquakes. *Journal of Geophysical Research: Space Physics*, 2017.
- 261 He, L., Wu, L., Pulinets, S., Liu, S., Yang, F. A.: Nonlinear background removal method for  
 262 seismo-ionospheric anomaly analysis under a complex solar activity scenario: A case  
 263 study of the M9.0 Tohoku earthquake. *Advances in Space Research*, 50(2), 211-220,  
 264 2012.
- 265 Heki, K. and Enomoto, Y.: Mw dependence of the preseismic ionospheric electron  
 266 enhancements. *Journal of Geophysical Research: Space Physics*, 120(8), 7006-7020,  
 267 2015.
- 268 Kelley, M. C., Swartz, W. E., Heki, K.: Apparent ionospheric total electron content variations  
 269 prior to major earthquakes due to electric fields created by tectonic stresses. *Journal*  
 270 *of Geophysical Research: Space Physics*, 2017.
- 271 Kamogawa, M. and Kakinami, Y.: Is an ionospheric electron enhancement preceding the 2011  
 272 Tohoku-Oki earthquake a precursor?. *Journal of Geophysical Research: Space*  
 273 *Physics*, 118(4), 1751-1754, 2013.
- 274 Langley R. B.: Monitoring the Ionosphere and Neutral Atmosphere with GPS Division of  
 275 Atmospheric and Space Physics Workshop, Fredericton, N.B., 2002.
- 276 Li, M. and Parrot, M.: Statistical analysis of the ionospheric ion density recorded by DEMETER  
 277 in the epicenter areas of earthquakes as well as in their magnetically conjugate point  
 278 areas. *Advances in Space Research*, 2017.
- 279 Liu, J. Y., Chen, C. H., Chen, Y. I., Yang, W. H., Oyama, K. I., Kuo, K. W.: A statistical study  
 280 of ionospheric earthquake precursors monitored by using equatorial ionization  
 281 anomaly of GPS TEC in Taiwan during 2001–2007. *Journal of Asian Earth*  
 282 *Sciences*, 39(1-2), 76-80, 2010.
- 283 Masci, F., Thomas, J. N., Villani, F., Secan, J. A., Rivera, N.: On the onset of ionospheric  
 284 precursors 40 min before strong earthquakes. *Journal of Geophysical Research: Space*  
 285 *Physics*, 120(2), 1383-1393, 2015.
- 286 Pulinets S. A.: Strong earthquakes prediction possibility with the help of top side sounding from  
 287 satellites. *Advances in Space Research* 21(3): 455–458, 1998.



- 288 Pulinets, S. and Davidenko, D.: Ionospheric precursors of earthquakes and global electric  
 289 circuit. *Advances in Space Research*, 53(5), 709-723, 2014.
- 290 Rozhnoi, A., Solovieva, M., Parrot, M., Hayakawa, M., Biagi, P. F., Schwingenschuh, K.,  
 291 Fedun, V.: VLF/LF signal studies of the ionospheric response to strong seismic activity  
 292 in the Far Eastern region combining the DEMETER and ground-based  
 293 observations. *Physics and Chemistry of the Earth, Parts A/B/C*, 85, 141-149, 2015.
- 294 Schaer S.: Mapping and Predicting the Earth's Ionosphere Using the Global Positioning System,  
 295 Doktora Tezi, University of Bern, İsviçre, 1999.
- 296 Thomas, J. N., Huard, J., Masci, F.: A statistical study of global ionospheric map total electron  
 297 content changes prior to occurrences of  $M \geq 6.0$  earthquakes during 2000–  
 298 2014. *Journal of Geophysical Research: Space Physics*, 122(2), 2151-2161, 2017.
- 299 Yildirim, O., Inyurt, S., Mekik, C.: Review of variations in  $M_w < 7$  earthquake motions on  
 300 position and TEC ( $M_w = 6.5$  Aegean Sea earthquake sample). *Natural Hazards and*  
 301 *Earth System Sciences*, 16(2), 543-557, 2016.
- 302 URL-1 <https://omniweb.gsfc.nasa.gov/form/dx1.html>