### **RESPONSE TO REVIEWER 2**

Thank you for your valuable comments. We have organized our paper as you suggest.

- 1- You suggest that figure 3-10 and table 1-10 are similar and therefore one gps station is enough to analyze. We have analyzed one gps station which is nearest receiver epicenter of earthquake.
- 2- You have also suggest that teporal and spatial resolution should be taken into account using GIM model. Therefore we have analyzed spatial and temporal analysis as you state
- 3- Discussion section was given as you suggest. Other conclusions were discussed and compared our results.
- 4- Line 41, line 45, line 43-45, 48-50, 57-59 were defined repeatedly
- 5- Line 95 reference was added, line 111 x and y axis and caption was again drawn as you state.
- 6- Equation 2 was edited correctly.
- 7- Fig 2 was drawn again as you state.
- 8- Line 148-149, 164-165, 185-186, 202-203 was organized again.

1	Monitoring potential ionospheric changes caused by Van earthquake (Mw
2	7.2)
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9	ABSTRACT
10	Many scientists from different disciplines have studied earthquakes for many years. As a result
11	of these studies, it has been proposed that some changes take place in the ionosphere layer
12	before, during or after earthquakes, and the ionosphere should be monitored in earthquake
13	prediction studies. This study investigates the changes in the ionosphere created by the
14	earthquake with magnitude of Mw=7.2 in the northwest of the Lake Erçek which is located to
15	the north of the province of Van in Turkey on 23 October 2011 and at 1.41 pm local time (-3
16	UT) with the epicenter of 38.75° N, 43.36° E using the TEC values obtained by the Global
17	Ionosphere Models (GIM) created by IONOLAB-TEC and CODE. In order to see whether the
18	ionospheric changes obtained by the study in question were caused by the earthquake or not,
19	the ionospheric conditions were studied by utilizing indices providing information on solar and
20	geomagnetic activities (F10.7 cm, Kp, Dst).
21	One of the results of the statistical test on the TEC values obtained from the both models,
22	positive and negative anomalies were obtained for the times before, on the day of and after the
23	earthquake, and the reasons for these anomalies are discussed in detail in the last section of the
24	study. As the ionospheric conditions in the analyzed days were highly variable, it was thought
25	that the anomalies were caused by geomagnetic effects, solar activity and the earthquake.
26	Keywords: TEC, Van Earthquake, Ionosphere
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### 36 1. INTRODUCTION

37 The ionosphere is the part of the atmosphere at the altitudes of 60 km to 1,100 km where there are ions and free electrons in considerable amounts that can reflect electromagnetic waves. It 38 39 completely covers the thermosphere, one of the main layers of the atmosphere, but also includes 40 some of the mesosphere and the exosphere. Total Electron Content (TEC), which is defined as The most important parameter that defines 41 42 the ionosphere in space and time is the amount of electrons. This amount varies under the 43 influence of the day night cycle, seasons, geographical location and magnetic storms in the sun. 44 Ionosphere has been monitored directly with some instruments like ionosonde, scatter radar. In 45 addition to that these instruments, GNSS is an effective tool to monitor effectively ionosphere 46 (Li and Parrot, 2017). Total Electron Content (TEC), which is defined as the amount of free 47 electrons along a cylinder with a cross section of 1 m<sup>2</sup>, is a suitable parameter to monitor the 48 changes in the ionosphere in space and time. All signals that contain data that pass through or 49 get reflected from the ionosphere, which is highly irregular and difficult to model, are affected 50 by the structure of this layer. 51 Calculating Total Electron Content (TEC) is a method used directly to investigate the structure 52 of the ionosphere. TEC is represented by the unit of TECU, and one TECU equals to  $10^{16} el/m^2$  (Schaer, 1999). TEC is expressed in two ways: STEC (Slant Total Electron 53

54 Content); the free electron content calculated along the slanted line between the receiver and 55 the satellite, and VTEC (Vertical Total Electron Content); the free electron content calculated 56 along the zenith of the receiver (Langley, 2002).

57 The ionosphere reacts to geomagnetic effect, solar activity, diurnal and seasonal effects, 11

- 58 year-solar-cycle, earthquake, and these factors cause irregularities in the ionosphere
  59 (Namgaladze et al, 2012, Li and Parrot, 2017).
- 60 TEC, which is defined as the number of free electrons on the one square meter area on the line
- 61 followed by a radio wave, is one of the important parameters for examining the structure of the
- 62 ionosphere and the upper atmosphere. With TEC values, it is possible to examine the short and
- 63 long term changes in the ionosphere, ionospheric irregularities and disruptive factors together
- 64 (Erol and Arıkan 2005).
- 65 Ionospheric changes have been studied in more than twenty countries today as precursors of
- 66 earthquakes. Definition of ionospheric anomalies and feasibility studies of seismo-ionospheric
- 67 precursors are still ongoing (Liu et al., 2010; He et al., 2012; Kamogawa and Kakinami, 2013;
- Heki and Enomoto, 2015; Pulinets and Davidenko, 2014; Masci et al., 2015; Yildirim et al.,

69 2016; He and Heki, 2017; Kelley et al., 2017; Rozhnoi et al., 2015; Thomas et al., 2017;

70 Ulukavak and Yalcinkaya 2017).

### 71 2. METHODOLOGY

### 72 2.1 IONOLAB-TEC Method:

The IONOLAB-TEC method developed by the department of Electrical and Electronics
 Engineering of Hacettepe University is a JAVA application that uses the Regularized TEC (D TEI) algorithm (Arikan et al. 2004 ).

In this application, they developed a method that estimates VTEC values by using all GPS signals measured at a period of time in a day. While the measurements taken from the satellites with elevations of  $60^{\circ}$  or higher are used, the measurements from the satellites with elevations of  $10^{\circ}$  to  $60^{\circ}$  are weighted by a Gauss function. The data from satellites with elevations of lower than  $10^{\circ}$  are not included in calculations to reduce multipath effects. In this method raw GPS data was used to determine VTEC value.

# 82 83 2.2 Global Ionosphere Model (GIM):

84 85 Global Ionospheric Maps are published in the IONEX (IONosphere map EXchange) format in 86 a way that covers the entire world. The institutions that produce these maps in the world include CODE (Center for Orbit Determination in Europe, Switzerland), DLR (Fernerkundungstation 87 88 Neustrelitz, Germany), ESOC (European Space Operations Centre, Germany), JPL (Jet Propulsion Laboratory, California), NOAA (National Oceanic and Atmospheric 89 90 Administration, United States), NRCan (National Resources, Canada), ROB (Royal 91 Observatory of Belgium, Belgium), UNB (University of New Brunswick, Canada), UPC 92 (Polytechnic University of Catalonia, Spain), WUT (Warsaw University of Technology, 93 Poland). In this study we used the GIM-TEC values produced by CODE in the IONEX format. 94 In the dates they were analyzed, the temporal resolution of the TEC values was 2 hours, while their positional resolution was 2.5° by latitude and 5° by longitude. In order to calculate TEC 95 96 values for a point whose latitude and longitude is known on the GIM-TEC maps created by 97 CODE using more than 300 GNSS receivers around the world, the 4 TEC values that cover the 98 point and the two-variable interpolation formula are given below. 99  $E_{int}(\lambda_0 + p\Delta\lambda, \beta_0 + q\Delta\beta) = (1-p)(1-q)E_{0.0} + p(1-q)E_{1.0} + q(1-p)E_{0.1} + pqE_{1.1}$ (1)

100 p and q:  $0 \le p, q < 1$ .

101  $\Delta \lambda$  and  $\Delta \beta$ : Longitude and Latitude differences grid widths,

102  $\lambda_0$  and  $\beta_0$ : Initial longitude and latitude values,

103  $E_{0.0}, E_{1.0}, E_{0.1}$  ve  $E_{1.1}$ : TEC values known in neighboring points,



### **3. ANALYSIS TO DETERMINE EARTHQUAKE-RELATED TEC CHANGES**

In order to investigate earthquake-related TEC changes, the TEC values for the stations close to the epicenters, HAKK, MALZ, OZAL stationand TVAN (TUSAGA-Acktive CORS-TR) close to the epicenters GPS station wasstations were analyzed to determine TEC value using the IONOLAB-TEC and GIM-TEC models. The correlation coefficient was obtained for the TEC values from both models between the dates 13.10.2011 and 02.11.2011 for the stations above. In addition to that, spatial analysis was applied to determine distribution characteristics





Biçimlendirilmiş: Normal

### Figure 1. Demonstration of analyzed stationAnalyzed Stations

Figure 1 shows the stations analyzed (represented by red triangles) and the epicenter of the earthquake represented by blue star. For each station, the TEC values with the temporal resolution of two hours obtained from both the IONOLAB-TEC and GIM-TEC models for OZAL station which is nearest station to epicenter of earthquake and the correlation coefficient was computed to explainshowing whether there is a linear relationship between two models. On the other hand, TEC values were also obtained using GIM model to explain spatial changes of ionosphere for IZMI, AFYN, KAYS and BING stations.calculated as below;

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$$r = \frac{1}{n-1} \Sigma\left(\frac{X-\bar{X}}{S_X}\right) \Sigma\left(\frac{Y-\bar{Y}}{S_Y}\right)$$
(2)

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128 In order to determine the outlier values among the TEC values with a two-hour temporal 129 resolution from both models, the TEC values obtained from both models between the dates 130 01.10.2011 and 10.10.2011, which were considered quiet in terms of geomagnetic and solar 131 activity, were used to determine the upper boundary (UB) and the lower boundary (LB). By 132 utilizing the TEC values from both models, the UB and LB values were calculated using the 133 formulae x+3 $\sigma$  and x-3 $\sigma$ . Here, x is the mean TEC value for the relevant epoch and  $\sigma$  is the 134 standard deviation. If the TEC value in any epoch is higher than the upper boundary, it is a 135 positive anomaly. Similarly if it is lower than the lower boundary, it is a negative anomaly. In 136 order to investigate whether the anomalies before, on the day of and after the earthquake were caused by the earthquake or not, we also examined the (Kp\*10), Dst and F10.7 cm indices, 137 138 which provided information on the geomagnetic and solar activity for the days in which 139 anomalies were detected.



Biçimlendirilmiş: Ortadan



Figure 3. GIM-TEC Values for the OZALHAKK Station

	GIM-TEC Anomaly Table for OZAL Station												
<u>Number</u>	DOY	<u>Hour</u>	TEC Difference (TECU)	Type of Anomaly	<u>Number</u>	DOY	<u>Hour</u>	TEC Difference (TECU)	Type of Anomaly				
<u>1</u>	288	<u>2</u>	2.0	Positive	<u>11</u>	<u>295</u>	<u>10</u>	<u>3.3</u>	Positive				
<u>2</u>	<u>288</u>	<u>10</u>	5.7	Positive	<u>12</u>	<u>296</u>	<u>4</u>	<u>1.9</u>	Positive				
<u>3</u>	<u>289</u>	<u>10</u>	2.5	Positive	<u>13</u>	<u>296</u>	<u>10</u>	7.5	Positive				
<u>4</u>	290	<u>10</u>	0.5	Positive	<u>14</u>	<u>297</u>	<u>10</u>	4.1	Positive				
<u>5</u>	292	<u>10</u>	0.8	Positive	<u>15</u>	<u>298</u>	<u>0</u>	0.8	Positive				
<u>6</u>	293	<u>10</u>	<u>5.2</u>	Positive	<u>16</u>	<u>298</u>	<u>2</u>	2.6	Positive				
<u>7</u>	294	<u>8</u>	0.7	Positive	<u>17</u>	<u>298</u>	<u>8</u>	<u>12.2</u>	Positive				
<u>8</u>	294	<u>10</u>	4.0	Positive	<u>18</u>	<u>298</u>	<u>10</u>	<u>11.7</u>	Positive				
<u>9</u>	<u>294</u>	<u>12</u>	<u>10.5</u>	Positive	<u>19</u>	<u>298</u>	<u>12</u>	<u>16.5</u>	Positive				
<u>10</u>	295	8	2.9	Positive	<u>20</u>	298	18	0.8	Positive				

Biçimlendirilmiş: Yazı tipi rengi: Siyah





	IONOLAB-TEC Anomaly Table for OZAL Station												
<u>Number</u>	<u>DOY</u>	<u>Hour</u>	<u>TEC</u> Difference (TECU)	Type of Anomaly	Number	DOY	<u>Hour</u>	TEC Difference (TECU)	Type of Anomaly				
<u>1</u>	<u>288</u>	<u>10</u>	<u>5.1</u>	Positive	<u>9</u>	<u>297</u>	<u>10</u>	<u>6.0</u>	<u>Positive</u>				
<u>2</u>	<u>289</u>	<u>10</u>	<u>1.6</u>	Positive	<u>10</u>	<u>298</u>	<u>0</u>	<u>2.2</u>	<u>Positive</u>				
<u>3</u>	<u>290</u>	<u>10</u>	<u>0.9</u>	Positive	<u>11</u>	<u>298</u>	<u>2</u>	<u>2.4</u>	Positive				
<u>4</u>	<u>292</u>	<u>12</u>	<u>0.6</u>	<u>Positive</u>	<u>12</u>	<u>298</u>	<u>4</u>	<u>4.1</u>	<u>Positive</u>				
<u>5</u>	<u>293</u>	<u>10</u>	<u>3.5</u>	Positive	<u>13</u>	<u>298</u>	<u>6</u>	<u>3.0</u>	<u>Positive</u>				
<u>6</u>	<u>294</u>	<u>12</u>	<u>11.8</u>	<u>Positive</u>	<u>14</u>	<u>298</u>	8	7.3	Positive				
<u>7</u>	<u>295</u>	<u>10</u>	<u>7.4</u>	<u>Positive</u>	<u>15</u>	<u>298</u>	<u>10</u>	<u>13.6</u>	Positive				
<u>8</u>	<u>296</u>	<u>10</u>	<u>9.6</u>	<u>Positive</u>	<u>16</u>	<u>298</u>	<u>12</u>	<u>12.8</u>	Positive				

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	-IONOLAB-TEC Anomaly Table for HAKK Station													
Number	ĐOY	Hour	TEC Difference (TECU)	Type of Anomaly	Number	ĐOY	Hour	TEC Difference (TECU)	Type of Anomaly					
1	<u>287</u>	<u>12</u>	<del>0.4</del>	Positive	9	<del>295</del>	<u>12</u>	<del>7.2</del>	Positive					
2	<del>288</del>	<del>12</del>	<del>9.2</del>	Positive	<del>10</del>	<del>296</del>	<del>12</del>	<del>8.8</del>	Positive					
3	<del>289</del>	<del>12</del>	<del>4.3</del>	Positive	<del>11</del>	<del>297</del>	<del>12</del>	<del>4.6</del>	Positive					
4	<del>290</del>	<u>12</u>	<del>3.8</del>	Positive	<del>12</del>	<u>298</u>	8	<del>16.5</del>	Positive					
5	<del>291</del>	<del>12</del>	<del>4.5</del>	Positive	<del>13</del>	<del>301</del>	<del>12</del>	<del>0.3</del>	Negative					
<del>6</del>	<del>292</del>	<del>12</del>	<del>1.4</del>	Positive	<del>14</del>	<del>302</del>	<del>14</del>	<del>0.9</del>	Negative					
7	<del>293</del>	<del>12</del>	<del>4.2</del>	Positive	<del>15</del>	<del>303</del>	<del>12</del>	<del>0.7</del>	<b>Negative</b>					
8	<del>294</del>	<del>12</del>	<del>10.9</del>	Positive	<del>16</del>	<del>306</del>	<del>10</del>	<del>0.9</del>	Positive					

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-Table 2. <u>OZALHAKK</u> Station IONOLAB-TEC Anomaly Table

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Below are the TEC values for the MALZ station obtained using the GIM-TEC and IONOLAB-TEC methods (Figures 5 and 6).





	GIM-TEC Anomaly Table for MALZ Station													
Number	ÐOY	Hour	TEC Difference (TECU)	Type of Anomaly		Number	<del>DOY</del>	Hour	TEC Difference (TECU)	Type of Anomaly				
1	<del>288</del>	<del>12</del>	<del>3.5</del>	Positive		5	<del>295</del>	<del>12</del>	<del>3.1</del>	Positive				
2	<del>289</del>	<del>12</del>	<del>0.5</del>	Positive		6	<del>296</del>	<del>12</del>	<del>7.9</del>	Positive				
3	<del>293</del>	<del>12</del>	<del>3.9</del>	Positive		7	<del>297</del>	<del>12</del>	4.7	Positive				
4	<del>294</del>	<del>12</del>	<del>8.6</del>	Positive		8	<del>298</del>	8	<del>12.6</del>	Positive				

-Table 3. MALZ Station Global Ionosphere Model Anomaly Table

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	IONOLAB-TEC Anomaly Table for MALZ Station													
Number	ÐO¥	Hour	TEC Difference (TECU)	Type of Anomaly	Number	ÐO¥	Hour	TEC Difference (TECU)	Type of Anomaly					
1	<del>288</del>	<del>12</del>	<del>2.3</del>	Positive	5	<del>296</del>	<del>12</del>	<del>2.5</del>	Positive					
2	<del>293</del>	<del>12</del>	<del>0.4</del>	Positive	<del>6</del>	<del>298</del>	<del>6</del>	<del>8.6</del>	Positive					
3	<del>294</del>	<del>10</del>	7.4	Positive	7	<del>304</del>	θ	<del>0.2</del>	<b>Negative</b>					
4	295	<del>10</del>	3.6	Positive										

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### -Table 4. MALZ Station IONOLAB-TEC Anomaly Table

187 Tables 3 and 4 show the anomalies found as a result of the analysis of the TEC values obtained

by the IONOLAB TEC and GIM TEC methods for the MALZ station. It is believed that the



positive anomaly on days 288 and 289 was caused by moderate (136.9 sfu, 150 sfu) solar

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Biçimlendirilmiş: İki Yana Yasla

Biçimlendirilmiş: Yazı tipi rengi: Otomatik



	Biçimlendirilmiş Tablo
	Biçimlendirilmiş Tablo
	Biçimlendirilmiş: Türkçe
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	Biçimlendirilmiş: Türkçe
	Biçimlendirilmiş: Sola
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	Biçimlendirilmiş: Türkçe
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	Biçimlendirilmiş: Türkçe
	Biçimlendirilmiş: Sola
	Biçimlendirilmiş: Türkçe
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	Biçimlendirilmiş: Türkçe
the second secon	Biçimlendirilmiş: Sola
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			<u>GIM</u> IONO	LAB-TEC Anomaly	1	Table for	<u>AFYN</u>	OZAL S	Station	•
Number	DOY	<u>Hour</u>	TEC Difference (TECU)	Type of Anomaly		Number	DOY	<u>Hour</u>	TEC Difference (TECU)	Type of Anomaly
<u>1</u>	<u>288</u>	<u>10</u>	<u>4.5</u>	<u>Positive</u>		<u>8</u>	<u>296</u>	<u>10</u>	<u>7.1</u>	Positive
<u>2</u>	<u>292</u>	<u>10</u>	<u>2.3</u>	Positive		<u>9</u>	<u>296</u>	<u>12</u>	<u>0.1</u>	Positive
<u>3</u>	<u>293</u>	<u>10</u>	<u>2.2</u>	Positive		<u>10</u>	<u>297</u>	<u>10</u>	<u>3.2</u>	Positive
<u>4</u>	<u>294</u>	<u>8</u>	<u>1.8</u>	Positive		<u>11</u>	<u>298</u>	<u>2</u>	2.3	Positive
<u>5</u>	<u>294</u>	<u>10</u>	<u>6.2</u>	<u>Positive</u>		<u>12</u>	<u>298</u>	<u>8</u>	<u>2.1</u>	Positive
<u>6</u>	<u>295</u>	<u>10</u>	<u>3.3</u>	<u>Positive</u>		<u>13</u>	<u>298</u>	<u>10</u>	<u>12.8</u>	Positive
<u>7</u>	<u>296</u>	<u>4</u>	<u>0.8</u>	Positive		<u>14</u>	<u>298</u>	<u>12</u>	<u>14.2</u>	<u>Positive</u>
15 16			<u>Table</u>	<b>4.</b> AFYN Station C	G	IM-TEC	Anom	aly Ta	<u>ible</u>	
		<sup>70</sup> 60 –			ĸ	AYS Low	ver Bound —	- Upper Bou	nd Mean Observ	ed-TEC
217		30 20 10 286	288 290			298	300	302	304 306	308
.18			rigur	e / GIM-TEC Val	u	es for the	NAI	s stat		
			<u>GI</u>	M-TEC Anomaly T	a	ble for K/	AYS St	<u>ation</u>	TEC	
<u>Number</u>	DOY	<u>Hour</u>	Difference (TECU)	Type of Anomaly		<u>Number</u>	<u>DOY</u>	<u>Hour</u>	Difference (TECU)	Type of Anomaly
<u>1</u>	<u>288</u>	<u>10</u>	<u>4.6</u>	Positive		<u>9</u>	<u>295</u>	<u>10</u>	<u>4.0</u>	Positive
<u>2</u>	<u>289</u>	<u>10</u>	<u>1.2</u>	Positive		<u>10</u>	<u>296</u>	<u>8</u>	<u>1.4</u>	Positive
<u>3</u>	<u>290</u>	<u>10</u>	<u>0.1</u>	<u>Positive</u>		<u>11</u>	<u>296</u>	<u>10</u>	<u>7.8</u>	<u>Positive</u>
<u>4</u>	<u>292</u>	<u>10</u>	<u>2.1</u>	<u>Positive</u>		<u>12</u>	<u>297</u>	<u>10</u>	<u>3.9</u>	<u>Positive</u>
<u>5</u>	<u>293</u>	<u>10</u>	<u>4.0</u>	<u>Positive</u>		<u>13</u>	<u>298</u>	<u>2</u>	<u>4.3</u>	<u>Positive</u>
<u>6</u>	<u>294</u>	<u>8</u>	<u>4.0</u>	Positive		<u>14</u>	<u>298</u>	<u>8</u>	<u>2.9</u>	Positive
<u>7</u>	<u>294</u>	<u>10</u>	<u>8.2</u>	Positive	L	<u>15</u>	<u>298</u>	<u>10</u>	<u>12.1</u>	<u>Positive</u>
<u>8</u>	<u>295</u>	<u>8</u>	0.1	<u>Positive</u>		<u>16</u>	<u>298</u>	<u>12</u>	<u>15.2</u>	<u>Positive</u>
219 220	<u>294</u> <u>295</u>	<u>10</u> <u>8</u>	8.2 0.1 Table	Positive Positive 5. KAYS Station C	G	<u>15</u> <u>16</u> IM-TEC	<u>298</u> 298 Anom	<u>10</u> 12 aly Ta	<u>12.1</u> <u>15.2</u> ble	Positive Positive



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**IONOLAB-TEC Anomaly Table for TVAN Station** 

Number	ĐOY	Hour	TEC Difference (TECU)	Type of Anomaly	Number	ĐOY	Hour	TEC Difference (TECU)	Type of Anomaly
1	<del>288</del>	<del>12</del>	<del>5.1</del>	Positive	<del>10</del>	<del>296</del>	<del>12</del>	<del>3.4</del>	Positive
2	<del>290</del>	<del>12</del>	<del>2.6</del>	Positive	<del>11</del>	<del>297</del>	<del>12</del>	<del>8.5</del>	Positive
3	<del>291</del>	<del>12</del>	<del>2.0</del>	Positive	<del>12</del>	<del>298</del>	<del>10</del>	<del>10.5</del>	Positive
4	<del>292</del>	<del>12</del>	<del>4.0</del>	Positive	<del>13</del>	<del>299</del>	<del>10</del>	<del>2.8</del>	Positive
5	<u>293</u>	<u>12</u>	<u>8.1</u>	Positive	<del>14</del>	<del>302</del>	<u>12</u>	0.7	Negative
<del>6</del>	<del>294</del>	<del>12</del>	<del>5.1</del>	Positive	<del>15</del>	<del>306</del>	<del>10</del>	<del>2.9</del>	Positive
7	<del>295</del>	<del>12</del>	<del>3.2</del>	Positive					

#### Table 8. TVAN Station IONOLAB-TEC Anomaly Table

243 Tables 1, 2, 3, 4, 5, 6, 7 and 8 show the results of the statistical analysis of the TEC values 244 ereated by the IONOLAB-TEC and GIM-TEC methods. The tables also depict the day and hour 245 in which anomalies were observed, and the amount and type of the anomaly. The numbers of 246 anomalies obtained in both models were very close to each other. The F10.7 cm index values 247 between the days 2886 and 292 were 136 sfu, 135.4 sfu, 136.9 sfu, 150 sfu, 151.6 sfu, 145.7 sfu, 146.1 sfu. The index values show that there was usually moderate solar activity. Therefore, 248 249 the anomalies in question may be related to the earthquake or solar activity. The index values for the days 293, 294, 295 and 296 (the day of the earthquake) were 157.8 sfu, 166.3 sfu, 162.5 250 251 sfu and 153.9 sfu respectively. These values indicate strong solar activity. On the other hand, 252 the ionosphere layer was quiet in these days in terms of geomagnetic conditions. As there was 253 strong solar activity, the numbers of anomalies were higher than the numbers in the days 2886-254 292. Since solar activity was moderate in the day 297, the number of anomalies dropped. The 255 solar activity on the day 298 was moderate, but there was strong geomagnetic activity (Dst -256 147 nt, Kp\*10=73). The reason for the high numbers of anomalies on day 298 in both models 257 is believed to be due to geomagnetic activity. This magnetic storm has caused different amount 258 of TEC variation for all stations.

259 As another indicator, we extract  $\Sigma$ ATEC (Totally TEC difference) to determine total amount of

260 <u>anomaly day by day for each analyzed days.</u>

Stations/A	<u>288</u>	<u>289</u>	<u>290</u>	<u>292</u>	<u>293</u>	<u>294</u>	<u>295</u>	<u>296</u>	<u>297</u>	<u>298</u>
<u>nomaly</u>	<u>(<b>Σ</b>AT</u>	<u>(<b>Σ</b>AT</u>	<u>(Σ</u> ΑΤ	<u>(Σ</u> ΑΤ	<u>(Σ</u> ΑΤ	<u>(ΣΑΤ</u>	<u>(ΣΑΤ</u>	<u>(<b>Σ</b>AT</u>	<u>(Σ</u> ΑΤ	<u>(<b>Σ</b>AT</u>
<u>Day</u>	<u>EC)</u>	<u>EC)</u>	<u>EC)</u>	<u>EC)</u>	<u>EC)</u>	<u>EC)</u>	<u>EC)</u>	<u>EC)</u>	<u>EC)</u>	<u>EC)</u>
IZMI-	<u> </u>	0.2	<u> </u>	1.8	0.1	<u>3.9</u>	<u>2</u>	<u>6.2</u>	2.1	<u>28.5</u>
<u>GIM</u>										
<u>AFYN-</u>	4.5	<b>_</b>	<u> </u>	2.3	2.2	<u>8</u>	<u>3.3</u>	<u>8</u>	<u>3.2</u>	<u>31.4</u>
GIM										

KAYS-	<u>4.6</u>	<u>1.2</u>	<u>0.1</u>	<u>2.1</u>	<u>4</u>	<u>12.2</u>	<u>4.1</u>	<u>9.2</u>	<u>3.9</u>	<u>34.5</u>
GIM										
BING-	<u>5.6</u>	<u>2.1</u>	<u>0.4</u>	<u>1.4</u>	<u>5</u>	<u>15.8</u>	<u>5.6</u>	<u>9.6</u>	<u>4.1</u>	<u>39.1</u>
GIM										
OZAL-	<u>7.7</u>	<u>2.5</u>	0.5	<u>0.8</u>	<u>5.2</u>	<u>15.2</u>	<u>6.2</u>	<u>9.4</u>	<u>4.1</u>	<u>44.5</u>
GIM										

Table 7. Total amount of anomaly in TECU for analyzed days

264Table 7 shows total anomaly summary results obtained from analysis results. Positive265anomalies were observed before and after the earthquake and amount of anomaly is nearly equal266to each other in this earthquake. In addition to that,  $\boldsymbol{\Sigma}$ ATEC differences between stations are267also similar to each other for in each analyzed day.

Considering the analyzed days in general <u>for all stations</u>, it may be seen that it is difficult to identify earthquake-related anomalies as the solar activity and geomagnetic conditions before and after the earthquake were not quiet. Therefore, it is believed that the anomalies detected in the stations on days 293-296 may be related to the earthquake and/or solar activity, and the anomalies on days 297 and 298 may be related to the earthquake, solar activity and/or geomagnetic activity.

### 275 4. <u>DISCUSSION AND CONCLUSION</u>

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276 Seismic ionospheric evalutions of Van earthquake have also been studied by many researchers 277 (Arikan et al., 2012; Zolotov et al., 2012; Rolland 2013; Şentürk et al., 2018). (Arikan et al., 278 2012; Zolotov et al., 2012) determined some anomalies before and after the earthquake, but 279 solar and magnetic conditions were not taken into account. On the other hand (Sentürk et al. 280 2018) also obtained abnormal days before and after the earthquake and They evaluated solar 281 activity and magnetic storm conditions for these abnormal days to explain possible causes of 282 anomalies in detail. Some previous studies have also studied on both space weather and 283 earthquake effect in the ionosphere (Yao et al., 2012; Le et al., 2013). They especially state that 284 TEC enhancement may be related to geomagnetic storm and earthquake. 285 (Sentürk et al., 2018) study also shows that there is no obvious anomaly caused only by 286 earthquake. Therefore they suggest that A multidisciplinary study would be useful to identify

287 <u>ionospheric changes as an earthquake precursor under the disturbed space-weather conditions.</u>

- 288 <u>This approach shows that their results agree with our study.</u>
- In the scope of this study, the TEC values for the stations <u>IZMI, AFYN, KAYS, BINGHAKK</u>,
- 290 MALZ, OZAL, TVAN were obtained using the GIM-TEC and TEC values were also obtained
- 291 using GIM-TEC and IONOLAB-TEC methods for OZAL station. In the comparison of the

Biçimlendirilmiş: Normal, Madde işaretleri veya numaralandırma yok Biçimlendirilmiş: Yazı tipi rengi: Siyah obtained values, it was seen that there was high correlation between the TEC values obtained by the two models for OZAL station .- In order to detect earthquake-related TEC changes better, the TEC values created from both models for the period of 13.10.2011-02.11.2011 were used as reference to determine the upper bound and lower bound values. As a result of the statistical test, anomalies were found in all analyzed stations for before, on the day of and after the earthquake. In order to understand whether the anomalies obtained in both models were earthquake-related, the ionospheric conditions, geomagnetic activity and solar activity on the analyzed days were examined using the Kp, Dst and F10.7 cm indices.

Consequently, it was determined that the positive anomalies observed on days 286-292 may be-related to moderate solar activity and/or the earthquake, and the positive anomalies observed on days 293, 294, 295, 296 (day of the earthquake) may be related to strong solar activity and/or the earthquake. Moderate solar activity and strong geomagnetic activity were observed for day 298, so the numbers of anomalies in both models increased dramatically. This increase is considered to be related to geomagnetic activity. The anomaly on day 298 may be related to the earthquake, geomagnetic effects and/or solar activity. The finding that the ionospheric conditions were variable in the analyzed days makes it highly difficult to identify earthquake-related ionospheric changes. Therefore, interdisciplinary study isstudies are needed to determine the earthquake-related part of the change in question.

Biçimlendirilmiş: Sekme durakları: Eskisi 4.23 cm

Biçimlendirilmiş: Türkçe

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