



Transport coefficients in standard Kappa distributed plasmas: a comparative study

Mahmood J. Jwailes, Imad A. Barghouthi, and Qusay S. Atawnah

Department of physics, Al-Quds University, Jerusalem, Palestine

Correspondence: Mahmood J. Jwailes (mahmood.jwailes@students.alquds.edu)

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Abstract. This study systematically derives transport coefficients – electrical conductivity, thermoelectric, diffusion, and mobility – for a Lorentz plasma described by a standard Kappa distribution function. Within the five-moment transport framework, the standard Kappa distribution serves as the zeroth-order function. Momentum and energy collision terms are obtained via the Boltzmann collision integral for Coulomb, hard-sphere, and Maxwell molecule interactions, and incorporated into the momentum equation to formulate generalized Ohm’s and extended Fick’s laws, yielding the transport coefficients. This study also compares the standard Kappa, modified Kappa, and Maxwellian distributions in terms of their influence on plasma behavior. The results show that for velocity-dependent collisions, such as Coulomb collisions, significant differences arise between the standard and modified Kappa distributions. For low kappa parameter κ values, the standard Kappa distribution reduces collision frequency and thermalization, making it suitable for collisionless or weakly collisional plasmas. In contrast, the modified Kappa distribution increases these effects, indicating its relevance for more collisional environments. Consequently, in Coulomb collisions, the standard distribution weakens momentum and energy exchange compared to the Maxwellian case, while the modified distribution enhances them. Transport properties are also affected differently: as κ decreases, the standard distribution enhances conductivity, mobility, diffusion, and thermoelectric effects, whereas the modified distribution reduces conductivity, mobility, and diffusion, with no change in the thermoelectric coefficient.

1 Introduction

Transport processes in plasma can be described using transport equations, which provide a macroscopic representation of the system through velocity moments of the particle distribution function (Schunk, 1977; Schunk and Nagy, 2009; Bittencourt, 2004). These equations are based on linear relationships between the fluxes (e.g., particle flux, heat flux, and current density) and the external forces and gradients (e.g., in density, temperature, and pressure) that drive those fluxes.

The constants of proportionality in these linear relations are known as transport coefficients – namely, the diffusion coefficient, electrical conductivity, mobility coefficient, thermoelectric coefficient, and thermal conductivity – which quantify how particles and energy move through a plasma under the influence of gradients, and applied external forces. Each coefficient characterizes a different aspect of transport, that is, the diffusion coefficient measures the flux of particles driven by spatial variations in density, providing insight into how species spread within the plasma. The mobility coefficient describes how charged particles drift in response to an applied electric field, and it is directly related to the electrical conductivity, which connects the current density to the electric field. The thermoelectric coefficient links electric fields to temperature gradients and characterizes the generation of electric voltages and currents in nonuniform thermal environments. Finally, the thermal conductivity determines the heat flux produced by temperature gradients and governs the rate of thermal energy transport within the plasma (Du, 2013; Wang and Du, 2017; Ebne Abbasi et al., 2017; Ebne Abbasi and Esfandyari-Kalejahi, 2019; Husidic et al., 2021).

For plasmas near thermal equilibrium, the Maxwellian distribution is commonly used to evaluate transport coefficients. However, space and astrophysical plasmas often con-

tain nonthermal particle populations that cause particle velocity distributions to deviate from the Maxwellian form. In these nonthermal environments, such distributions are well fitted by the Kappa velocity distribution functions (Marsch, 2006). Unlike the Maxwellian distribution, the Kappa distributions introduce a power-law tail that decays more slowly than the exponential tail of the Maxwellian, making them able to capture the presence of suprathermal particles, a feature that the Maxwellian distribution fails to describe. The shape of this tail is controlled by the kappa parameter κ , which determines the strength of the high-energy tail: larger κ values approach the Maxwellian limit, while smaller values emphasize suprathermal populations. With typical κ values ranging between 2 and 6, Kappa distributions have been observed across diverse plasma environments, including the solar wind, Earth's magnetosheath, and Jupiter's magnetosphere, supported by direct measurements from satellite missions such as Ulysses, Cluster, and Voyager 2, see Vasyliunas (1968), Pierrard et al. (2001), Maksimovic et al. (1997), Qureshi et al. (2003), Formisano et al. (1973), Collier and Hamilton (1995) for details on these missions and their observations of the Kappa distributions.

Consequently, several studies have extensively investigated transport coefficients in nonequilibrium plasmas using the Kappa velocity distribution functions.

In particular, studies such as Du (2013), Wang and Du (2017), Ebne Abbasi et al. (2017), Ebne Abbasi and Esfandyari-Kalejahi (2019) and Guo and Du (2019) derived diffusion, mobility, electrical conductivity, thermoelectric coefficients, and thermal conductivity based on modified Kappa distributions, which assume a κ -independent effective temperature and therefore produce a stronger low-energy core and enhanced suprathermal tails relative to the Maxwellian distribution.

However, this modified form differs fundamentally from the standard (Olbertian) Kappa distribution, introduced by Olbert (1968) and Vasyliunas (1968), in which the effective temperature is κ -dependent, leading to a weaker thermal core and more pronounced high-energy tails. As a result, the standard Kappa distribution is widely used in space plasma studies and is particularly suitable for modeling collisionless or weakly collisional plasmas, as it accurately fits spacecraft observations and captures the nonthermal heating of suprathermal populations. Beyond observational fitting, Lazar et al. (2015); Lazar et al. (2016) showed that the modified Kappa approach may yield results in linear dispersion and instability analyses that do not always highlight the destabilizing role of suprathermal particles, sometimes indicating reduced growth rates. In contrast, the standard Kappa formalism tends to produce instability thresholds and growth rates that more clearly reflect the destabilizing influence of suprathermal populations in collisionless or weakly collisional plasmas.

This empirical and theoretical success in collisionless plasma regions has motivated the development of transport

coefficients based on the standard Kappa formalism, as done by Husidic et al. (2021).

Nevertheless, the modified Kappa distribution remains valuable from a theoretical perspective, as it provides a self-consistent framework for describing systems in which deviations from equilibrium redistribute particles between the core and suprathermal tails, and with presence of collisional plasmas. Importantly, as highlighted by Husidic et al. (2021), distinctions between the modified Kappa and standard Kappa distributions are crucial because the choice of distribution directly affects the resulting transport coefficients and their physical interpretation.

All reviewed studies employed simplified collision models rather than the full Boltzmann collision integral in evaluating the transport coefficients. The simplest models appear in Wang and Du (2017), Ebne Abbasi and Esfandyari-Kalejahi (2019), and Husidic et al. (2021) which used a Krook-type operator, offering computational simplicity but limited accuracy. More physically based models – such as those proposed by Du (2013) and Guo and Du (2019) – used macroscopic transport equations derived from idealized relaxation assumptions. The most advanced work, presented by Ebne Abbasi et al. (2017), used the Fokker-Planck equation to model Coulomb collisions. While this captures cumulative small-angle scattering and better represents long-range Coulomb forces, it remains an approximation of the Boltzmann collision integral. Thus, all reviewed works share the same limitation: reliance on simplified collision models.

To overcome this limitation, Jwailes et al. (2025) recently introduced a comprehensive re-evaluation of the transport coefficients based on the modified Kappa distribution, using the five-moment transport approximation with the Boltzmann collision integral as the collision model. In this approach, a new transport theory was developed by deriving the five-moment approximation and the corresponding collision terms for various types of collisions for the modified Kappa distribution. The five-moment momentum equation is then linked to the generalized Ohm's law and the extended Fick's law, from which the transport coefficients are determined.

This study is inspired by the work of Husidic et al. (2021) and follows the same methodology introduced by Jwailes et al. (2025). As in Husidic et al. (2021), we focus on evaluating the transport coefficients for the standard Kappa distribution, but we adopt the methodology used in Jwailes et al. (2025), particularly in the formulation of the transport equations, the evaluation of the collision integrals, and the derivation of the transport coefficients. However, in contrast to Husidic et al. (2021), we use the Boltzmann collision integral as our collision model rather than the Krook-type collision model. This substitution is essential for obtaining results that more accurately capture the velocity-dependent interaction dynamics inherent to Kappa-distributed plasmas. The study further examined the transport coefficients derived from both the modified and standard Kappa distributions, highlighting

how the selection of a particular Kappa model can markedly alter the magnitude and trends of these coefficients.

This paper is structured as follows: Section 2 provides a brief review of the Kappa distribution family, introducing the mathematical formulations and the physical interpretation of two different types of suprathermal tail distributions: the standard Kappa and the modified Kappa distribution functions. It also explains how their behaviors differ from that of the Maxwellian distribution. Section 3 presents the theoretical framework of this paper, in which we derive the five-moment approximation and the corresponding collision terms for the standard Kappa velocity distribution function, considering arbitrary drift-velocity and temperature differences between the interacting plasma species. This includes three types of collisions: Coulomb collisions, hard-sphere interactions, and Maxwell-molecule collisions. The section concludes with the derivation of the transport coefficients using the five-moment approximation and the derived collision terms.

Section 4 discusses the derived results presented in Sect. 3 for the standard Kappa distribution and compares them with the corresponding results for both the modified Kappa distribution and the Maxwellian distribution. Three aspects are considered in the comparison: (i) the effective collision frequency and thermalisation rate; (ii) the behavior of the collision terms in the case of Coulomb collisions, with a focus on how collisions affect both the momentum and the energy of the interacting particles; and (iii) the transport coefficients and their dependence on the kappa parameter. The derived formulas are also compared with results from previous studies, highlighting their dependence on the kappa parameter. Finally, Sect. 5 presents the conclusions, summarizing the main findings of this study.

2 Distributions with suprathermal tails

Kappa distributions constitute a broad class of non-Maxwellian velocity distribution functions that effectively describe suprathermal particle populations in space and astrophysical plasmas. Among the various formulations proposed in the literature, two main types are commonly used: the standard Kappa distribution and the modified Kappa distribution. While both distributions share the general objective of characterizing plasmas with high-energy tails, they differ in their mathematical structure, parameter definitions, and physical interpretations.

The concept of the Kappa distribution was first proposed by Olbert (1968) to explain the presence of high-energy particles observed in the solar wind and magnetospheric plasmas, and was subsequently formalized by Vasyliunas (1968), who provided a more rigorous mathematical formulation. This early version is commonly referred to as the Olbertian or standard Kappa distribution (SK). In velocity space, the drifting standard Kappa distribution is given by Lazar and

Fichtner (2021) as,

$$f_s^{\text{SK}}(\mathbf{r}, \mathbf{c}_s, t) = \frac{n_s \xi(\kappa_s)}{\pi^{3/2} w_s^3} \left(1 + \frac{c_s^2}{\kappa_s w_s^2} \right)^{-\kappa_s - 1}, \quad (1)$$

where n_s denotes the number density and w_s is the thermal speed of species s , defined as

$$w_s = \sqrt{\frac{2k_B T_s}{m_s}}, \quad (2)$$

with m_s and T_s being the particle mass and the absolute temperature, respectively, and k_B the Boltzmann constant. The random velocity \mathbf{c}_s is defined in terms of the position \mathbf{r} , velocity \mathbf{v}_s , and the drift velocity $\mathbf{u}_s(\mathbf{r}, t)$ of the species s ,

$$\mathbf{c}_s(\mathbf{r}, \mathbf{v}_s, t) = \mathbf{v}_s - \mathbf{u}_s, \quad (3)$$

The kappa dependent function $\xi(\kappa_s)$ is given by

$$\xi(\kappa_s) = \kappa_s^{-3/2} \frac{\Gamma(\kappa_s + 1)}{\Gamma(\kappa_s - 1/2)}, \quad (4)$$

where κ_s is the kappa parameter that determines the slope of the power-law tails. Within this framework, the effective temperature T_s^κ , obtained via the second velocity moment, depends on the kappa parameter κ_s and is written as,

$$T_s^\kappa = \frac{\kappa_s}{\kappa_s - 3/2} \frac{m_s w_s^2}{2k_B} = \frac{\kappa_s}{\kappa_s - 3/2} T_s. \quad (5)$$

As κ_s increases, the effective temperature decreases until it approaches the Maxwellian temperature T_s (Lazar et al., 2016). This dependence implies that the enhanced presence of suprathermal particles contributes additional energy to the system, effectively heating the plasma. Moreover, the expression for the effective temperature in Eq. (5) imposes a condition on the kappa parameter, namely $\kappa_s > 3/2$; below this value the effective temperature diverges and is therefore undefined (Pierrard and Lazar, 2010).

Decades later, inspired by the principles of non-extensive statistical mechanics introduced by Tsallis (2012). Livadiotis (2017) developed a new theoretical perspective and reformulated the Kappa distribution into what is now known as the modified Kappa distribution (MK). In velocity space, the modified Kappa distribution is given by Livadiotis (2018) and Davis et al. (2023) as,

$$f_s^{\text{MK}}(\mathbf{r}, \mathbf{c}_s, t) = \frac{n_s \eta(\kappa_s)}{\pi^{3/2} w_s^3} \left(1 + \frac{c_s^2}{\kappa_0 w_s^2} \right)^{-\kappa_s - 1}, \quad (6)$$

where w_s is defined as in Eq. (2). The kappa dependent function in this case takes the form

$$\eta(\kappa_s) = \kappa_0^{-3/2} \frac{\Gamma(\kappa_s + 1)}{\Gamma(\kappa_s - 1/2)}, \quad \kappa_0 = \kappa_s - \frac{3}{2}. \quad (7)$$

Here, κ_s is the shape parameter that governs the slope of the suprathermal tails. As before, the condition $\kappa_s > 3/2$ must be

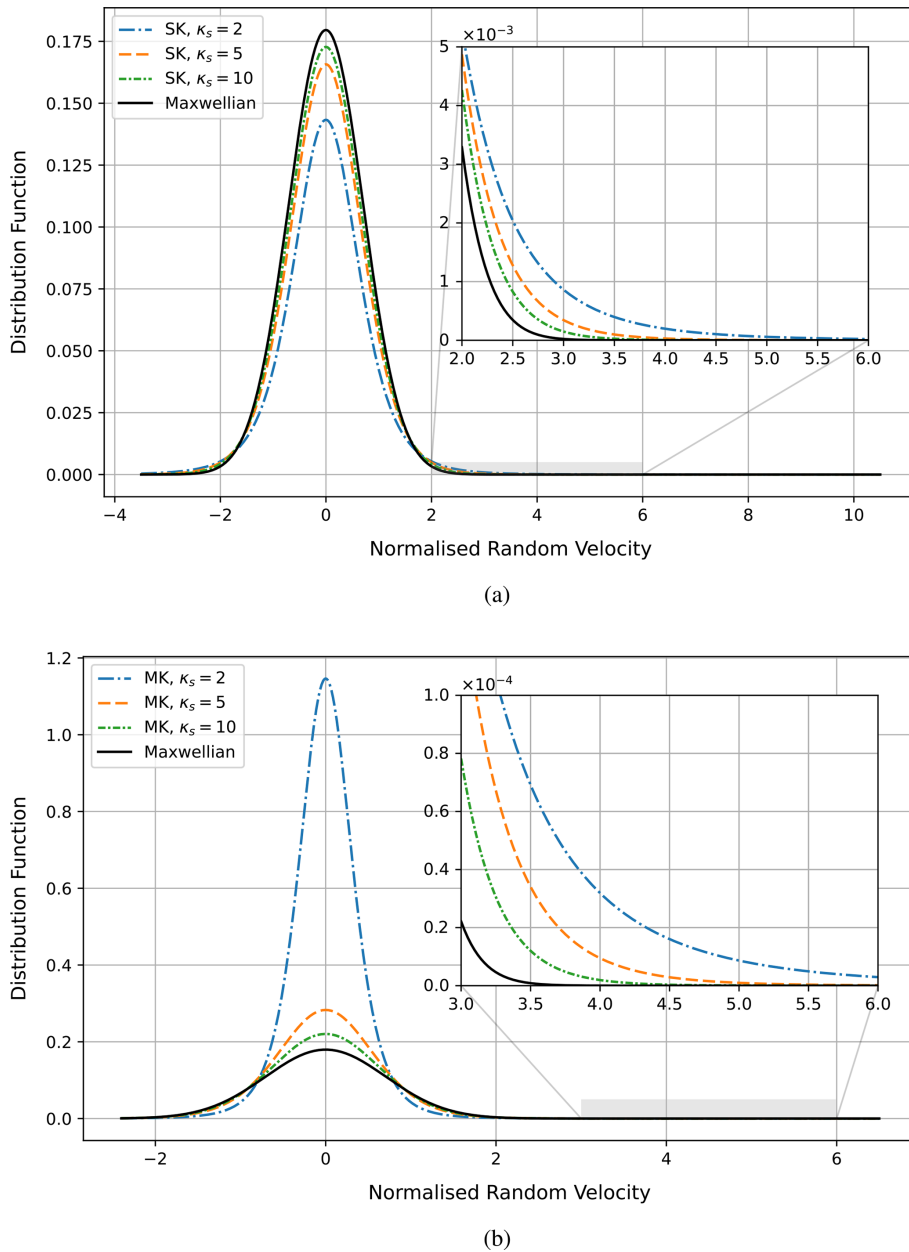


Figure 1. A schematic comparison between (a) standard Kappa, (b) modified Kappa velocity distributions for κ_s values 2, 5, and 10, with the Maxwellian velocity distribution as a functions of the normalised random velocity, c_s/w_s .

satisfied to ensure that the modified Kappa distribution function in Eq. (7) remains well defined. This modified version introduces a stronger thermodynamic basis by decoupling the effective temperature from the kappa parameter κ_s , making it a kappa independent quantity, as given by

$$T_s^\kappa = \frac{\kappa_{0s}}{\kappa_s - 3/2} \frac{m_s w_s^2}{2k_B} = T_s, \quad (8)$$

which is identical to the Maxwellian temperature and remains constant regardless of the value of κ_s (Lazar et al., 2016). This ensures that variations in the high-energy tails do

not change the overall thermal energy of the plasma. In this sense, the modified Kappa distribution maintains the same total thermal energy content as a Maxwellian plasma while redistributing the particles between the core and tail regions.

On the similarity side, both the standard and modified Kappa distributions are used to describe particle populations with suprathermal tails, since both distributions retain a power-law form and exhibit suprathermal tails that are higher than those of the Maxwellian distribution.

Moreover, both distributions reduce to the Maxwellian distribution in the limiting case where κ_s approaches infinity,

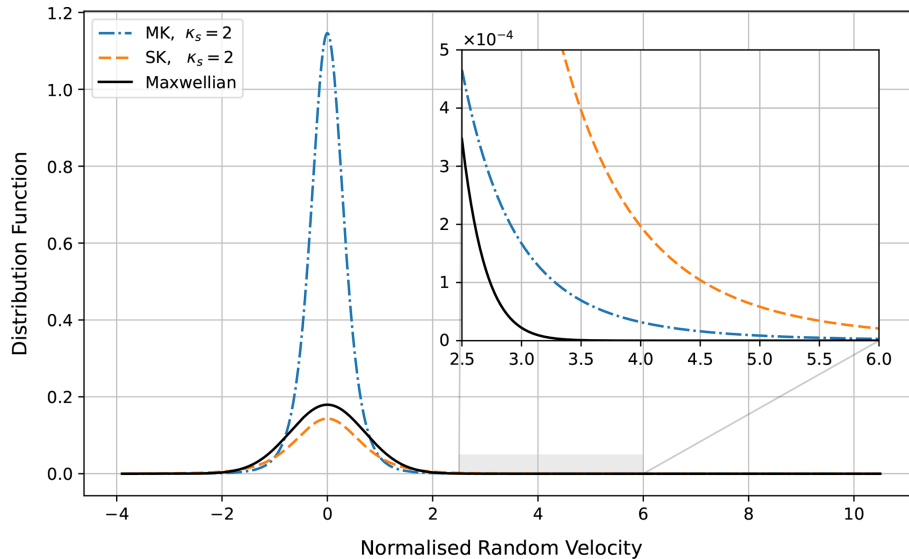


Figure 2. A schematic comparison of the standard Kappa distribution, the modified Kappa distribution for $\kappa_s = 2$, and the Maxwellian velocity distribution as a functions of the normalised random velocity, c_s/w_s .

(Pierrard and Lazar, 2010).

$$\lim_{\kappa_s \rightarrow \infty} f_s^{SK} = \lim_{\kappa_s \rightarrow \infty} f_s^{MK} = \frac{n_s}{\pi^{3/2} w_s^3} \exp\left(-\frac{c_s^2}{w_s^2}\right). \quad (9)$$

This behaviour is illustrated in Fig. 1, where increasing κ_s causes both the standard and modified Kappa distributions to converge smoothly toward the Maxwellian distribution. Although the standard and modified Kappa distributions share this common limiting behavior and exhibit similar qualitative features, they differ in their mathematical formulation and physical interpretation. The mathematical distinction between the two forms is mainly due to their parameterization and normalization. The standard distribution employs κ_s in the energy-dependent term, while the modified version replaces it with $\kappa_s - 3/2$. While this shift may appear minor, it significantly affects the scaling of the velocity distributions, resulting slightly flatter high-energy tails in the modified Kappa distribution compared to the standard Kappa distribution for the same κ_s value. Moreover, in the standard Kappa distribution, the effective temperature of the particles depends on κ_s , making it much higher than the temperature in the Maxwellian case. However, for the modified Kappa distribution, the effective temperature is independent of κ_s , making it identical to the Maxwellian temperature.

These differences are reflected in how the particle velocities are distributed. To illustrate how the two Kappa distributions differ from the Maxwellian distribution, Fig. 2 shows a comparison between the Maxwellian, the modified Kappa, and the standard Kappa distributions. The first thing we can notice is that both the modified and the standard Kappa distributions exhibit enhanced tails compared to the Maxwellian distribution, which is a defining characteristic of Kappa distributions. At the same time, we can also observe differences

in the shape of each distribution, which are directly related to the effective temperature.

In the standard Kappa distribution, the effective temperature T_s^κ is higher than that of both the Maxwellian and the modified Kappa distributions, as shown in Eq. (5). Consequently, the population of high-energy suprathermal particles (i.e., at large velocity magnitudes) is significantly enhanced compared to the other distributions. At the same time, this increase in high-energy particles is accompanied by a reduction in the particle population within the low-energy core (i.e., at small velocity magnitudes). On the other hand, in the modified Kappa distribution, the effective temperature is the same as in the Maxwellian distribution. To maintain this equality in temperature, particles are redistributed between the low-energy core and the high-energy tail without changing the system’s total thermal energy. As a result, the high-energy tail of the modified Kappa distribution is lower than that of the standard Kappa distribution, while the particle population in the low-energy core becomes significantly higher. The mathematical and physical properties of the Maxwellian, standard Kappa, and modified Kappa velocity distribution functions, discussed above, are summarized in Table 1.

Both the standard and modified Kappa distributions are used in different contexts. The standard Kappa distribution is the most commonly used tool in space plasma studies, where it provides excellent fits to spacecraft observations from the solar wind and planetary magnetospheres (Vasyliunas, 1968; Olbert, 1968; Collier and Hamilton, 1995; Collier et al., 1996; Maksimovic et al., 1997; Lazar and Fichtner, 2021). It captures the empirical relationship between suprathermal particle populations and the observed nonthermal heating of plasmas. On the other hand, the modified Kappa distribution, is mainly used in theoretical and statistical modeling,

Table 1. Mathematical and physical comparison of Maxwellian, standard Kappa, and modified Kappa velocity distribution functions.

Feature	Maxwellian (M)	Standard Kappa (SK)	Modified Kappa (MK)
Statistical nature	Thermal equilibrium	Non-equilibrium	Non-equilibrium
Theoretical basis	Boltzmann statistics	Empirical (space data)	Non-extensive statistics
Mathematical form	$f_s^M \propto \exp\left(-\frac{c_s^2}{w_s^2}\right)$	$f_s^{SK} \propto \left(1 + (c_s^2/\kappa_s w_s^2)\right)^{-\kappa_s-1}$	$f_s^{MK} \propto \left(1 + (c_s^2/\kappa_{0s} w_s^2)\right)^{-\kappa_s-1}$
Key shape parameter	–	κ_s	$\kappa_s, \kappa_{0s} = \kappa_s - 3/2$
Parameter constraint	–	$\kappa_s > 3/2$	$\kappa_s > 3/2$
Limit as $\kappa_s \rightarrow \infty$	–	Maxwellian distribution	Maxwellian distribution
Tail behavior	Exponential	Power-law	Power-law
Dominant region	Core-dominated	Tail-dominated	Core-dominated + tail
High-energy population	Lowest	Highest	Intermediate
Low-energy population	Intermediate	Lowest	Highest
Effective temperature, T_{eff}	T_s	$\frac{\kappa_s}{\kappa_s - 3/2} T_s$	T_s
Dependence of T_{eff} on κ_s	Independent of κ_s	Increases as κ_s decreases	Independent of κ_s
Total thermal energy	Baseline	Higher than Maxwellian	Same as Maxwellian

particularly in studies of systems governed by non-extensive entropy, long-range interactions, and quasi-stationary states (Lima et al., 2001; Livadiotis, 2015; Yoon, 2014; Livadiotis, 2017). It provides a self-consistent description of plasma systems that exhibit deviations from classical thermodynamic equilibrium without requiring an increase in thermal energy.

3 Theoretical Formulation

In this section, we present the five-moment approximation of the transport equation system, along with the corresponding collision terms and transport coefficients, using the standard Kappa distribution as the velocity distribution function. The derivation follows the same mathematical framework and analytical steps established in Jwailes et al. (2025). While the full detailed calculations are not repeated here, the essential assumptions and methodological structure remain the same.

3.1 Transport equations

The transport equations describe the spatial and temporal evolution of the physically significant velocity moments, such as number density, drift velocity, temperature, pressure tensor, stress tensor, and heat flow vector. These equations are obtained by multiplying the Boltzmann equation by an appropriate velocity-dependent function and then integrating over the velocity space, as presented in Schunk (1977), Schunk and Nagy (2009), and Bittencourt (2004). However, the general transport equations do not constitute a closed system since each moment equation depends on a higher-order moment. To close the system, the velocity distribution function, is approximated by expanding it into a complete orthogonal series around an appropriate zeroth-order distribu-

tion function $f_s^{(0)}$ (Grad, 1949; Mintzer, 1965). When only the first term of this expansion is retained, the species distribution function f_s , is represented by the zeroth-order function, $f_s^{(0)}$. The general system of transport equations then reduces to the so-called five-moment approximation, in which the stress, heat flux, and all higher-order moments are neglected. At this level of approximation, the properties of each species are described by five parameters: the number density, three components of drift velocity, and temperature. If the chosen zeroth-order distribution function $f_s^{(0)}$ has a stress tensor $\boldsymbol{\tau}_s$ and a heat flux vector \mathbf{q}_s equal to zero, as in the drifting Maxwellian, drifting modified Kappa, and drifting standard Kappa distributions (Scherer et al., 2019), and if the main external forces acting on the charged particles are gravitational and Lorentz forces, the five-moment approximation equations become (Schunk, 1977; Jwailes et al., 2025),

$$\frac{\delta n_s}{\delta t} = \frac{\partial n_s}{\partial t} + \nabla \cdot (n_s \mathbf{u}_s), \tag{10}$$

$$\begin{aligned} \frac{\delta \mathbf{M}_s}{\delta t} &= n_s m_s \frac{D_s \mathbf{u}_s}{Dt} + \nabla p_s^\kappa \\ &\quad - n_s m_s \mathbf{G} - n_s e_s \left(\mathbf{E} + \frac{\mathbf{u}_s \times \mathbf{B}}{c} \right), \end{aligned} \tag{11}$$

$$\frac{\delta E_s}{\delta t} = \frac{3}{2} \frac{D_s p_s^\kappa}{Dt} + \frac{5}{2} p_s^\kappa (\nabla \cdot \mathbf{u}_s). \tag{12}$$

In these equations, the symbol ∇ denotes the gradient in spatial coordinates. The operator D_s/Dt , is defined as

$$\frac{D_s}{Dt} = \frac{\partial}{\partial t} + \mathbf{u}_s \cdot \nabla. \tag{13}$$

The partial pressure associated with the species s can be expressed as

$$p_s^{\kappa} = n_s k_B T_s^{\kappa}, \tag{14}$$

with $n_s(\mathbf{r}, t)$ being the number density and $T_s^{\kappa}(\mathbf{r}, t)$ the effective temperature associated with the standard Kappa distribution given in Eq. (5). The charge and the mass of the species s are denoted by e_s and m_s , respectively. The vectors \mathbf{E} and \mathbf{B} correspond to the electric and magnetic fields, while \mathbf{G} represents the gravitational acceleration. The speed of light in vacuum is c . Finally, the terms appearing on the left-hand side of the five-moment approximation system, Eqs. (10)–(12), are known as the collision terms, or as the transfer collision integrals. These terms represent the rates of change of density, momentum, and energy due to collisions, respectively denoted by $\delta n_s/\delta t$, $\delta \mathbf{M}_s/\delta t$, and $\delta E_s/\delta t$.

The five-moment approximation given in Eqs. (10)–(12) differs from the approximations obtained for the Maxwellian and modified Kappa distributions (Jwailes et al., 2025). In particular, the partial pressure here depends explicitly on the Kappa parameter, κ_s . This dependence arises from the evaluation of the second-order velocity moment, where, for the standard Kappa distribution, the second moment of the random velocity is expressed as

$$\langle c_s^2 \rangle = \frac{3k_B T_s^{\kappa}}{m_s} = \frac{\kappa_s}{\kappa_s - 3/2} \frac{3k_B T_s}{m_s}. \tag{15}$$

In contrast, such a modification does not appear in the modified Kappa distribution, because its temperature is defined consistently with the Maxwellian case. As a result, the second moment retains the familiar Maxwellian form:

$$\langle c_s^2 \rangle = \frac{3k_B T_s}{m_s}. \tag{16}$$

3.2 Collision terms

The collision terms, the left-hand side of Eqs. (10)–(12), depend strongly on the collision model being considered. The appropriate expression for the collision model in the case of binary elastic collisions between particles (i.e., collisions governed by inverse power laws and resonant charge exchange) is the Boltzmann collision integral.

In this study, the Boltzmann collision integral is employed to model three types of collisions: Coulomb collisions, hard-sphere interactions, and Maxwell-molecule collisions. These interactions differ primarily in the range and character of the forces between particles. Coulomb collisions describe long-range electrostatic interactions between charged particles, resulting in cumulative small-angle deflections; in plasmas. In contrast, the hard-sphere model assumes that particles interact only upon direct contact, with instantaneous collisions determined purely by geometry, making it suitable for neutral gases with short-range interactions. Maxwell molecule collisions, based on a soft repulsive force decreasing with

the fifth power of distance, represent an intermediate case with short-range, smooth interactions (Schunk, 1977; Chapman and Cowling, 1990; Schunk and Nagy, 2009).

The general expressions for the collision terms, under the assumption that the velocity distribution functions of both interacting species, s and t , follow drifting standard Kappa distributions, are summarized below for the three types of collisions: Coulomb collisions, hard-sphere interactions, and Maxwell molecule collisions. For Coulomb collisions and hard-sphere interactions, the results are implicitly derived in Jwailes et al. (2025) and can be obtained by setting $\kappa_0 = \kappa$. In the case of Maxwell molecule collisions, the derivation follows the same procedure as in Jwailes et al. (2025), but with modifications in the expectation values.

$$\frac{\delta n_s}{\delta t} = 0, \tag{17}$$

$$\frac{\delta \mathbf{M}_s}{\delta t} = \sum_t n_s m_s v_{st}^{\text{SK}}(\kappa_s, \kappa_t) \Phi(\varepsilon_{st}) \Delta \mathbf{u}_{st}, \tag{18}$$

$$\begin{aligned} \frac{\delta E_s}{\delta t} = \sum_t n_s \left[\frac{3}{2} k_B v_{st,T}^{\text{SK}}(\kappa_s, \kappa_t) \Psi(\varepsilon_{st}) \Delta T_{st}^{\text{SK}} \right. \\ \left. + m_{st} v_{st}^{\text{SK}}(\kappa_s, \kappa_t) \Phi(\varepsilon_{st}) |\Delta \mathbf{u}_{st}|^2 \right], \end{aligned} \tag{19}$$

where the relative drift velocity $\Delta \mathbf{u}_{st}$ and relative temperature difference $\Delta T_{st}^{\text{SK}}$ are defined by

$$\Delta \mathbf{u}_{st} = \mathbf{u}_t - \mathbf{u}_s, \tag{20}$$

$$\Delta T_{st}^{\text{SK}} = H(\kappa_t) T_t - H(\kappa_s) T_s, \tag{21}$$

and the drift-to-thermal speed ratio ε_{st} is given by

$$\varepsilon_{st} = \frac{|\Delta \mathbf{u}_{st}|}{w_{st}}, \quad w_{st} = \sqrt{\frac{2k_B T_{st}}{m_{st}}}, \tag{22}$$

with the reduced mass m_{st} and the reduced temperature T_{st} are defined as

$$m_{st} = \frac{m_s m_t}{m_s + m_t}, \quad T_{st} = \frac{m_s T_t + m_t T_s}{m_s + m_t}. \tag{23}$$

The kappa parameter dependent terms v_{st}^{SK} and $v_{st,T}^{\text{SK}}$ represent, respectively, the effective collision frequency and the thermal equilibration rate (or simply the thermalisation rate) for systems described by the standard Kappa distribution, and they are defined as

$$v_{st}^{\text{SK}}(\kappa_s, \kappa_t) = v_{st} D(\kappa_s, \kappa_t), \tag{24}$$

$$v_{st,T}^{\text{SK}}(\kappa_s, \kappa_t) = 2 \frac{m_{st}}{m_t} v_{st}^{\text{SK}}, \tag{25}$$

where v_{st} denote the effective collision frequency rate for systems governed by the Maxwellian distribution. The functional forms of the factors v_{st} , Φ , Ψ , D , and H depend on

the type of collision considered. The factors v_{st} , Φ , and Ψ are listed in Jwailes et al. (2025) for all three types of collisions – Coulomb collisions, hard-sphere interactions, and Maxwell molecule collisions. The factors D and H , however, are summarized for the three types of collisions as follows, for *Coulomb collisions* and *hard-sphere interactions*, they are defined as:

$$D(\kappa_s, \kappa_t) = \frac{(\kappa_s - 1/2)}{\kappa_s} \frac{(\kappa_t - 1/2)}{\kappa_t}, \tag{26}$$

$$H(\kappa_\alpha) = \frac{\Gamma(\kappa_\alpha) \kappa_\alpha^{1/2}}{\Gamma(\kappa_\alpha + 1/2)}, \quad \alpha = s, t. \tag{27}$$

For *Maxwell molecule collisions*, they are defined as

$$D(\kappa_s, \kappa_t) = 1, \tag{28}$$

$$H(\kappa_\alpha) = \frac{\kappa_\alpha}{(\kappa_\alpha - 3/2)}, \quad \alpha = s, t. \tag{29}$$

A few remarks on the collision terms are worth noting. The collision terms for non-drifting standard Kappa distributions can be obtained by setting the drift velocities of both interacting particles s and t to zero, $\mathbf{u}_s = \mathbf{u}_t = 0$, in Eqs. (17)–(19). The same result holds when the drift velocities of species s and t are equal, i.e., $\mathbf{u}_s = \mathbf{u}_t$. Another point is that, in the limit as κ approaches infinity, $\kappa = \kappa_s = \kappa_t$, the collision terms, Eqs. (17)–(19), exactly recover the same results as those for the Maxwellian distribution (Schunk and Nagy, 2009), with the same definitions of Φ , Ψ , and v_{st} . That is, the effective collision frequency, the thermalisation rate, and the relative temperature difference, which are the terms that the collision terms depend on the kappa parameter through, reduce to their form in the Maxwellian case

$$\lim_{\kappa \rightarrow \infty} v_{st}^{SK}(\kappa, \kappa) = v_{st}, \tag{30}$$

$$\lim_{\kappa \rightarrow \infty} v_{st,T}^{SK}(\kappa, \kappa) = v_{st,T}, \tag{31}$$

$$\lim_{\kappa \rightarrow \infty} \Delta T_{st}^{SK} = T_t - T_s = \Delta T_{st}. \tag{32}$$

Hence,

$$\lim_{\kappa \rightarrow \infty} D(\kappa, \kappa) = \lim_{\kappa \rightarrow \infty} H(\kappa) = 1. \tag{33}$$

With $v_{st,T}$ denoting the thermalisation rate for systems governed by the Maxwellian distribution, defined as

$$v_{st,T} = 2 \frac{m_{st}}{m_t} v_{st}. \tag{34}$$

Obtaining the Maxwellian result provides a consistency check that the derived collision terms are correct, since the standard Kappa distribution reduces to a Maxwellian distribution when the kappa parameter κ approaches infinity, as discussed in Sect. 2.

3.3 Transport coefficients

A Lorentz plasma is a type of plasma characterized by negligible electron–electron collisions compared to electron–ion collisions, allowing the electrons to be treated as moving through a background of nearly stationary ions (Du, 2013). In this setting, and adopting the standard Kappa distribution, the transport coefficients – namely, the electrical conductivity σ_e , thermoelectric coefficient α_e , diffusion coefficient D_e , and mobility coefficient μ_e – can be derived using the five-moment approximation. The procedure starts from the momentum equation with a drifting standard Kappa distribution, Eq. (18), for a simple electron–ion collision. By assuming a steady, low-inertia regime with unmagnetized plasma ($\mathbf{B} = 0$), negligible gravitational effects ($\mathbf{G} = 0$), and negligible ion drift velocity ($\mathbf{u}_i \approx 0$), as appropriate for a Lorentz plasma, while allowing electrons to retain a small but finite drift velocity relative to the thermal speed (so that $\epsilon_{ei} = 0$), the electron momentum equation reduces to the following form (Jwailes et al., 2025),

$$-n_e \mathbf{u}_e = \frac{k_B T_e^\kappa}{m_e v_{ei}^{SK}} \nabla n_e + \frac{n_e k_B}{m_e v_{ei}^{SK}} \nabla T_e^\kappa + \frac{n_e e}{m_e v_{ei}^{SK}} \mathbf{E}. \tag{35}$$

By setting $\nabla n_e = 0$, as in Husidic et al. (2021), and substituting the expression for the effective temperature T_e^κ using Eq. (5), Eq. (35) reduces to the generalized Ohm’s law:

$$\mathbf{E} = \frac{\mathbf{J}_e}{\sigma_e} + \alpha_e \nabla T_e. \tag{36}$$

where \mathbf{E} denotes the electric field and \mathbf{J}_e is the current density, with e being the electron charge and n_e the electron number density. From this, we can identify the electrical conductivity and thermoelectric coefficient as

$$\sigma_e = \frac{n_e e^2}{m_e v_{ei}^{SK}}, \tag{37}$$

$$\alpha_e = -\frac{k_B}{e} \left(\frac{\kappa_e}{\kappa_e - 3/2} \right). \tag{38}$$

Alternatively, by setting $\nabla T_e = 0$, which consequently implies that $\nabla T_e^\kappa = 0$, as in Husidic et al. (2021), we obtain the extended Fick’s law:

$$\mathbf{\Gamma}_e = -D_e \nabla n_e - \mu_e n_e \mathbf{E}. \tag{39}$$

where $\mathbf{\Gamma}_e$ denotes the particle flux density, and the diffusion and mobility coefficients are identified as

$$D_e = \frac{k_B T_e^\kappa}{m_e v_{ei}^{SK}}, \tag{40}$$

$$\mu_e = \frac{e}{m_e v_{ei}^{SK}}. \tag{41}$$

Equations (37), (38), (40), and (41) represent the mathematical forms of the transport coefficients governing electron dynamics in a Lorentz plasma with a standard Kappa distribution. Together, they demonstrate that electrical conductivity,

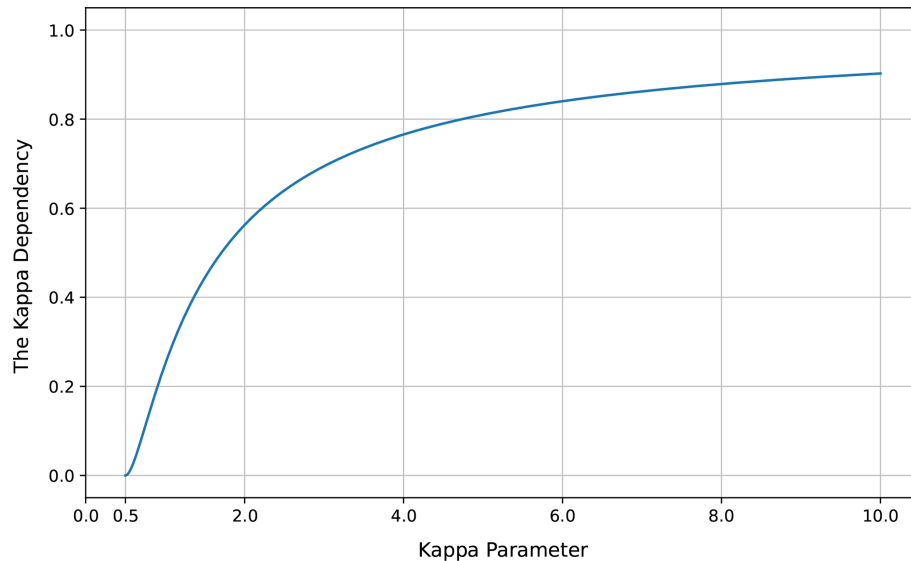


Figure 3. Dependence of the effective collision frequency ν_{st}^{SK} and thermalisation rate $\nu_{st,T}^{\text{SK}}$ on the kappa parameter κ for Coulomb collisions and hard-sphere interactions, as defined in Eq. (43).

thermoelectric, diffusion, and mobility coefficients are controlled primarily by the electron–ion effective collision frequency.

4 Comparison of collision processes and transport coefficients

In this section, we present a comprehensive comparison of the results derived in Sect. 3 for three types of distributions: the standard Kappa, modified Kappa, and Maxwellian distributions. The comparison focuses on three aspects. First, we examine the effective collision frequency and the thermalisation rate. Next, we analyze the collision terms, specifically for Coulomb collisions. Finally, we compare the resulting transport coefficients for each distribution.

4.1 Effective collision frequency and thermalisation rate

The effective collision frequency describes the average rate of how frequently collisions occur, determining the efficiency of momentum transfer within the system, while the thermalisation rate measures how rapidly the system approaches thermal equilibrium through collisions. Both quantities are essential for understanding the exchange of momentum and energy between particles due to collisions. Within the five-moment approximation of the transport equations, these quantities are obtained directly from the momentum and energy collision terms. Expressions for the standard Kappa distribution are given in Eqs. (24) and (25). Corresponding expressions for the modified Kappa distribution can be found in Jwailes et al. (2025), while those for the

Maxwellian distribution are provided in Schunk and Nagy (2009).

Equations (24) and (25) show that, for the standard Kappa distribution, both the effective collision frequency and the thermalisation rate are affected by the kappa-dependent function $D(\kappa_s, \kappa_t)$. This function depends on the kappa parameters κ_s and κ_t of the interacting species s and t and its form varies with the type of collision process considered. In collision processes such as Maxwell molecule interactions, where the collision frequency is independent of particle velocity, the redistribution of particles' velocities introduced by the standard Kappa distribution has no effect. In this case, $D = 1$, and both the effective collision frequency and the thermalisation rate remain identical to the Maxwellian case,

$$\nu_{st}^{\text{SK}} = \nu_{st}, \text{ and } \nu_{st,T}^{\text{SK}} = \nu_{st,T}. \quad (42)$$

In contrast, for collision processes that strongly depend on particle velocity, the standard Kappa distribution significantly affects both the effective collision frequency and the thermalisation rate. This effect becomes particularly evident in processes such as Coulomb collisions and hard-sphere interactions, where the velocity distribution plays a central role. In these cases, the function D varies according to the kappa parameters κ_s and κ_t , as given in Eq. (26). To compare the effective collision frequency and thermalisation rate with the Maxwellian case, and to better understand their behaviour, we consider the special case $\kappa = \kappa_s = \kappa_t$, so that the

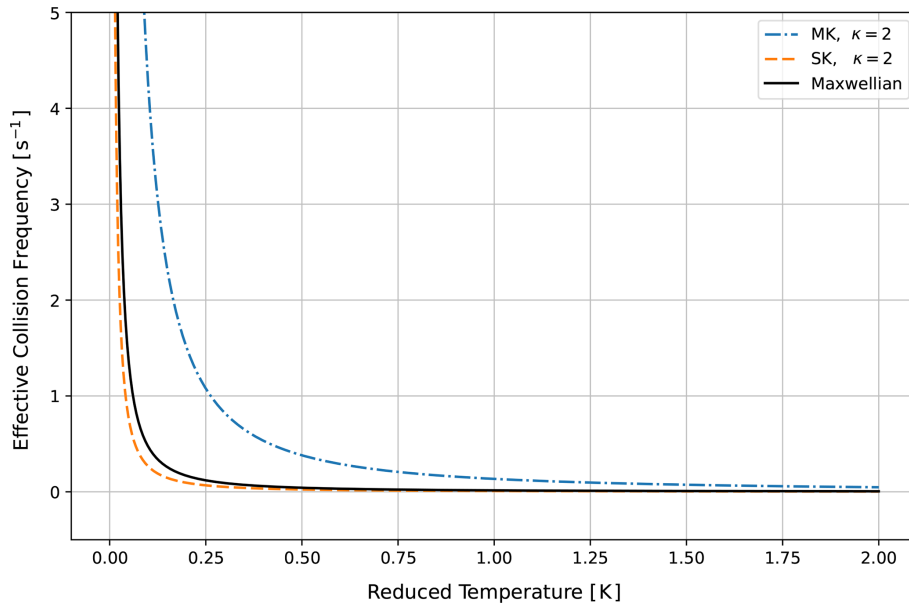


Figure 4. The effective collision frequency as a function of reduced temperature T_{st} for the Maxwellian, modified Kappa, and standard Kappa distributions in case of Coulomb collisions.

expressions, v_{st}^{SK} and $v_{st,T}^{SK}$, reduce to

$$v_{st}^{SK} = v_{st} \left(\frac{\kappa - 1/2}{\kappa} \right)^2, \tag{43}$$

$$v_{st,T}^{SK} = 2 \frac{m_{st}}{m_t} v_{st}^{SK}. \tag{44}$$

Relating Eq. (43) to the Debye length obtained for the standard Kappa distribution in Scherer et al. (2020), we recover exactly the same kappa dependence reported in that work, see Appendix A, thereby confirming the consistency of the present formulation. Equations (43) and (44) show that both the effective collision frequency and the thermalisation rate are reduced at low values of κ and increase as κ increases. As κ goes to infinity, the kappa term in Eq. (43) approaches 1, and the results converge to those of the Maxwellian distribution, as illustrated in Fig. 3. In this figure, we plot the kappa dependency for both the effective collision frequency and the thermalisation rate; in other words, the ratios v_{st}^{SK}/v_{st} and $v_{st,T}^{SK}/v_{st,T}$ as functions of the kappa parameter. This behaviour arises from the redistribution of the particles' velocities in the standard Kappa distribution. As discussed in Sect. 2, low values of κ correspond to a reduction in the population of particles near the core with a small velocity magnitude compared to a Maxwellian distribution. Since collision frequency in Coulomb collisions and hard sphere interactions are inversely proportional to function of velocity, this reduction leads to lower effective collision frequency and thermalisation rates at small κ values.

For the modified Kappa distribution, Jwailes et al. (2025) derived both the effective collision frequency and the ther-

malisation rate. Since both the standard and the modified Kappa distributions primarily redistribute particles' velocities, the effective collision frequency remains unchanged when the collision frequency is independent of velocity. This is the case for Maxwell molecule interaction, for which the collision frequency is constant across Maxwellian, standard Kappa, and modified Kappa distributions

$$v_{st}^{SK} = v_{st}^{MK} = v_{st}, \text{ and } v_{st,T}^{SK} = v_{st,T}^{MK} = v_{st,T}. \tag{45}$$

where v_{st}^{MK} and $v_{st,T}^{MK}$ represent the effective collision frequency and the thermalisation rate, respectively, for systems described by the modified Kappa distribution. For collisions in which the collision frequency depends on particle velocity, such as Coulomb collisions or hard-sphere interactions, the choice of distribution strongly affects the effective collision frequency and the thermalisation rate. As discussed earlier, in the standard Kappa distribution, low values of κ lead to a reduced effective collision frequency compared to the Maxwellian case. However, this is not the case for the modified Kappa distribution, which predicts the opposite behaviour, showing an increased effective collision frequency at low κ values. Figure 4 illustrates this behaviour by showing the effective collision frequency as a function of reduced temperature in the case of Coulomb collision for Maxwellian, standard Kappa, and modified Kappa distributions.

Figure 4, shows that all three distributions exhibit the same general behaviour. However, the standard Kappa distribution shows a lower effective collision frequency compared to the Maxwellian distribution, while the modified Kappa distribution shows a significantly higher effective collisions fre-

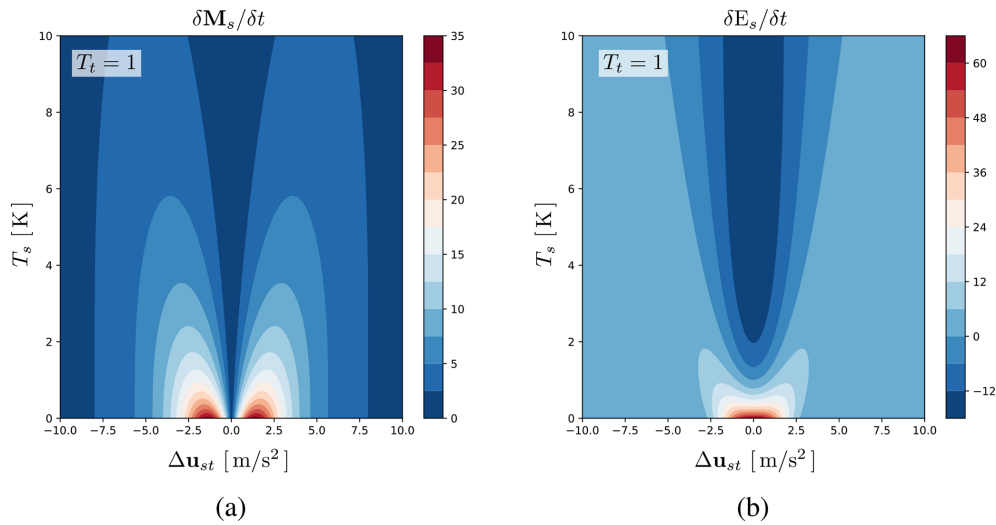


Figure 5. (a) The momentum collision term in units of $[N m^{-3}]$, and (b) the energy collision term in units of $[W m^{-3}]$ for the Maxwellian velocity distribution function in the case of Coulomb collisions.

quency relative to the Maxwellian case. This behaviour arises from the redistribution of particle velocities in the standard and modified Kappa distributions, as discussed in Sect. 2. At low κ values, the low number of particles near the low energy core in the standard Kappa distribution leads to a lower effective collision frequency and thermalisation rate compared to the Maxwellian case. For the modified Kappa distribution, the number of particles near the core with small velocity magnitudes is higher than in the Maxwellian distribution, which increases the collision frequency for Coulomb collisions and hard sphere interactions.

4.2 Collision terms

The collision terms for the five-moment approximation, presented in Eqs. (17)–(19), describe how the density, momentum, and energy for species s change due to collisions.

These terms depend on the number density n_s , drift velocity \mathbf{u}_s , and temperature T_s for species s , as well as on the corresponding parameters of species t , number density n_t , drift velocity \mathbf{u}_t , and temperature T_t . Additionally, two functions of κ_s and κ_t , namely $D(\kappa_s, \kappa_t)$ and $H(\kappa_\alpha)$, $\alpha = s, t$, which contribute to the effective collision frequency, the thermalisation rate, and the relative temperature difference. The particle masses m_s and m_t are constant and remain unchanged throughout the collision process for all types of collisions; as a result, the density collision term vanishes, as shown in Eq. (17).

In the Maxwellian case, both functions $D(\kappa_s, \kappa_t)$ and $H(\kappa_\alpha)$, $\alpha = s, t$, are set equal to one; see Sect. 3.2. The behaviour of the momentum and energy collision terms in this case was studied in detail by Jwailes et al. (2025), providing an explanation for the physical trends shown in Fig. 5a and b. Figure 5a shows the isolines of the magnitude of the mo-

mentum collision term, assuming that the direction of $\Delta\mathbf{u}_{st}$ along the z axis, while Fig. 5b shows the isolines of the corresponding energy collision term. Both figures display the dependence on $\Delta\mathbf{u}_{st}$ and T_s , with all other constants set to 1.0 for simplicity. Assuming identical parameters for all t particles, the summation over t in Eqs. (17)–(19) reduces to multiplication by their number, N_t , which is set to 1000 for the sake of comparison with other cases.

For the momentum collision term, the behavior closely follows the Maxwellian case, with $D(\kappa, \kappa)$ scaling the effective collision frequency, as shown in Figs. 6 and 7. At low κ , the effective collision frequency decreases, as discussed in Sect. 4.1, leading to reduced momentum transfer. To understand how the standard Kappa distribution changes the collision terms, we plot the isolines of the momentum and energy collision terms as functions of $\Delta\mathbf{u}_{st}$ and T_s , as shown in Fig. 6. We assume equal kappa values for both species, s and t , i.e., $\kappa_s = \kappa_t = \kappa$, to allow a direct comparison with the Maxwellian case and under the same conditions as in Fig. 5a. The corresponding cross-sections at $T_s = 0$ are shown in Fig. 7.

For the energy collision term, the function

$$W(\kappa, \kappa) = D(\kappa, \kappa) H(\kappa) \tag{46}$$

appears in the first term of Eq. (19), while $D(\kappa, \kappa)$ contributes to the second term. The overall behavior is similar to the Maxwellian case, with smaller values of κ yielding a smaller energy collision term. Overall, both collision terms increase with increasing κ , converging toward the Maxwellian result.

For the modified Kappa distribution, Jwailes et al. (2025) have studied the behavior of the collision term and compared it with that of the Maxwellian distribution under the same conditions previously applied to the standard Kappa distri-

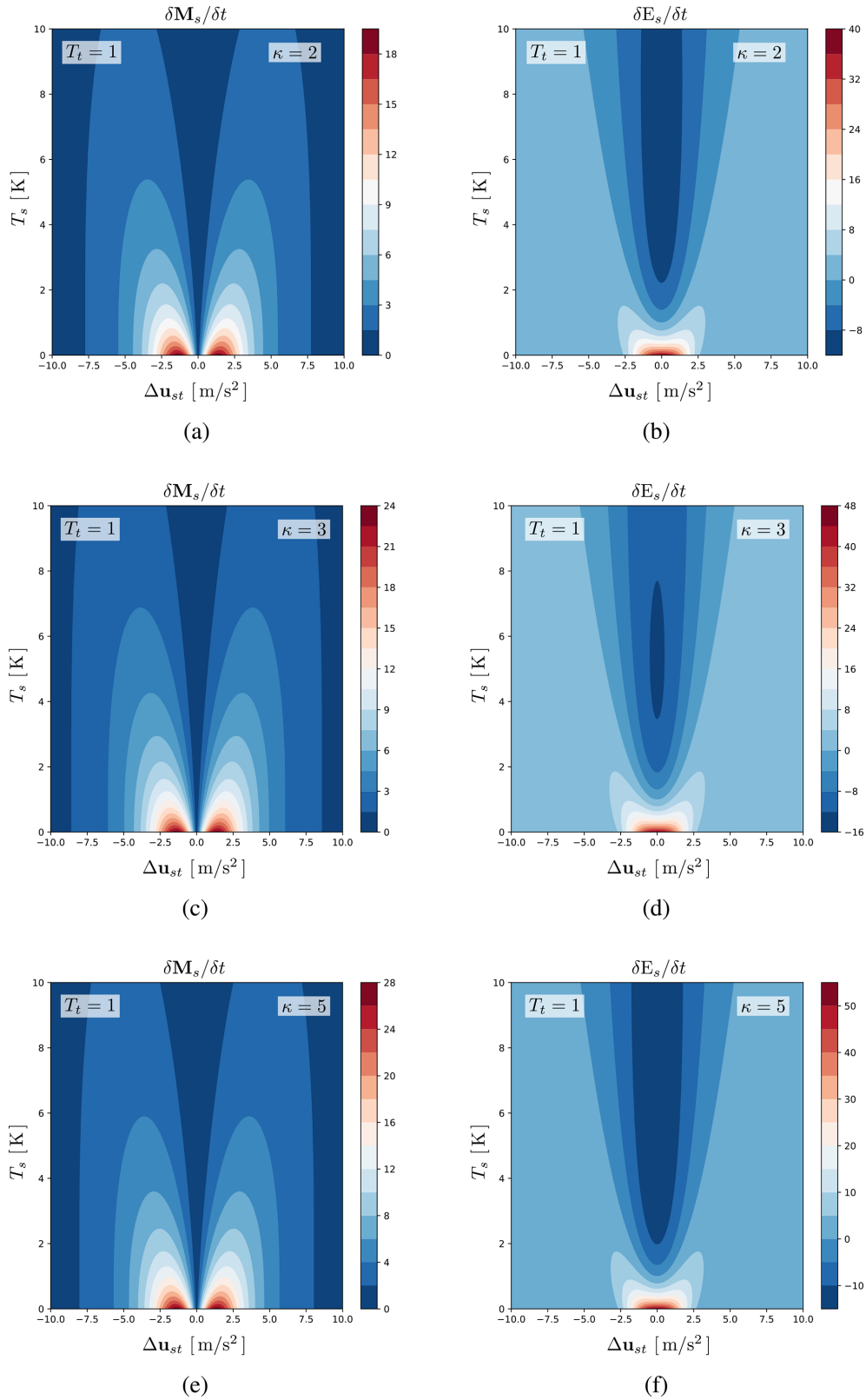


Figure 6. (a, c, e) The momentum collision term in units of [N m⁻³], and (b, d, f) the energy collision term in units of [W m⁻³] for the standard Kappa velocity distribution function in the case of Coulomb collisions at different values of κ : 2, 3, and 5.

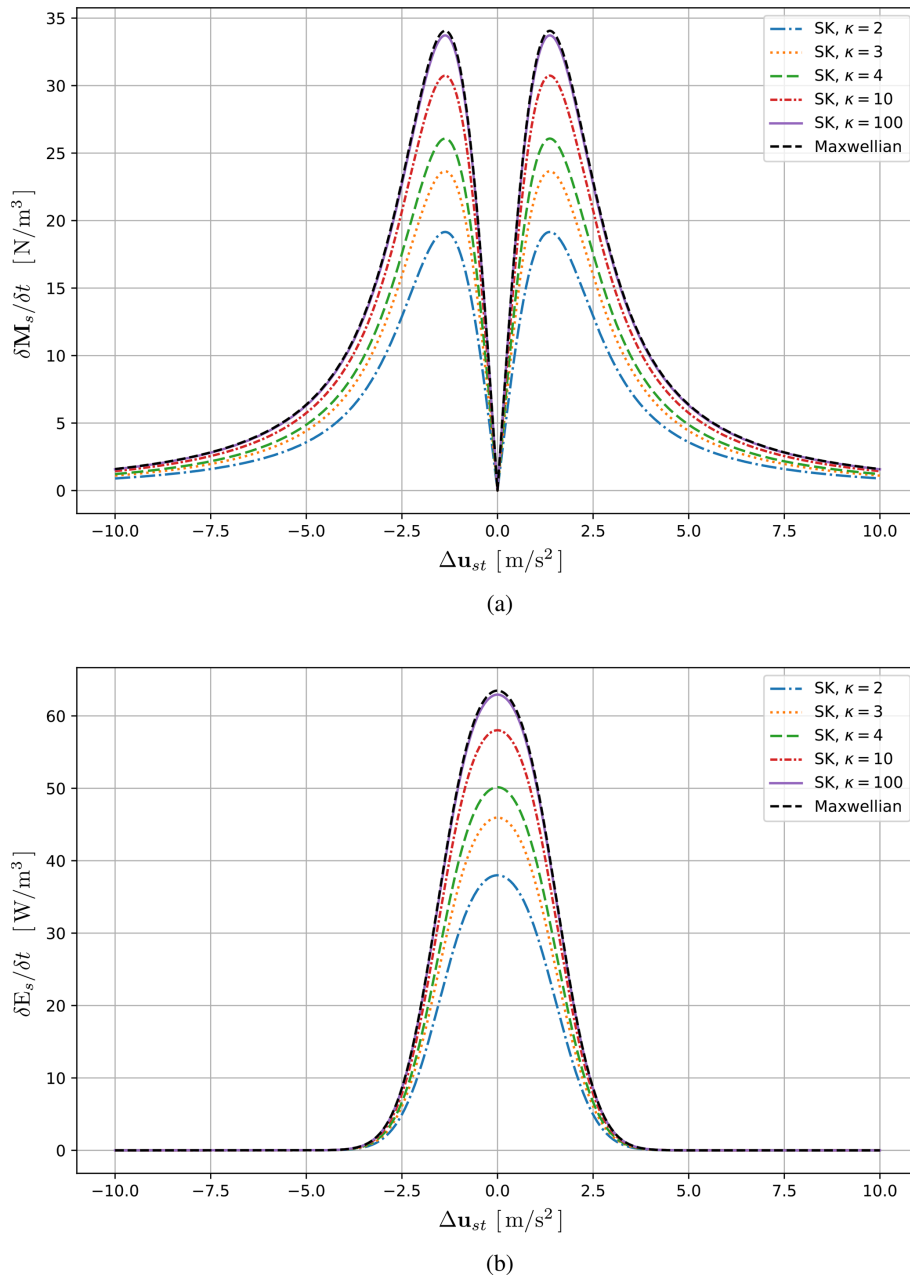


Figure 7. The cross-section of the momentum and energy collision terms for the standard Kappa and Maxwellian velocity distribution functions in the case of Coulomb collisions at $T_s = 0$.

bution. The results show that the collision terms behave similarly to the standard Kappa and the Maxwellian distribution; however, the modified Kappa distribution amplifies the collision terms at low values of κ . That is, collisions have a stronger influence on momentum and energy exchange between particles due to Coulomb interactions, which is the opposite behavior of the standard Kappa distribution discussed earlier. This significant difference is shown in Fig. 8, which presents the cross sections of the momentum and energy collision terms at $T_s = 0$ as functions of $\Delta \mathbf{u}_{st}$. It is clear that,

at the same value of κ , the collision terms in the modified Kappa distribution are much larger than those in both the standard Kappa and the Maxwellian distributions. This behavior is consistent with the results of Sect. 4.1, where we found that the effective collision frequency and the thermalization rate are significantly higher for the modified Kappa distribution than for the standard Kappa distribution, as a result of how the particles distribute near the core.

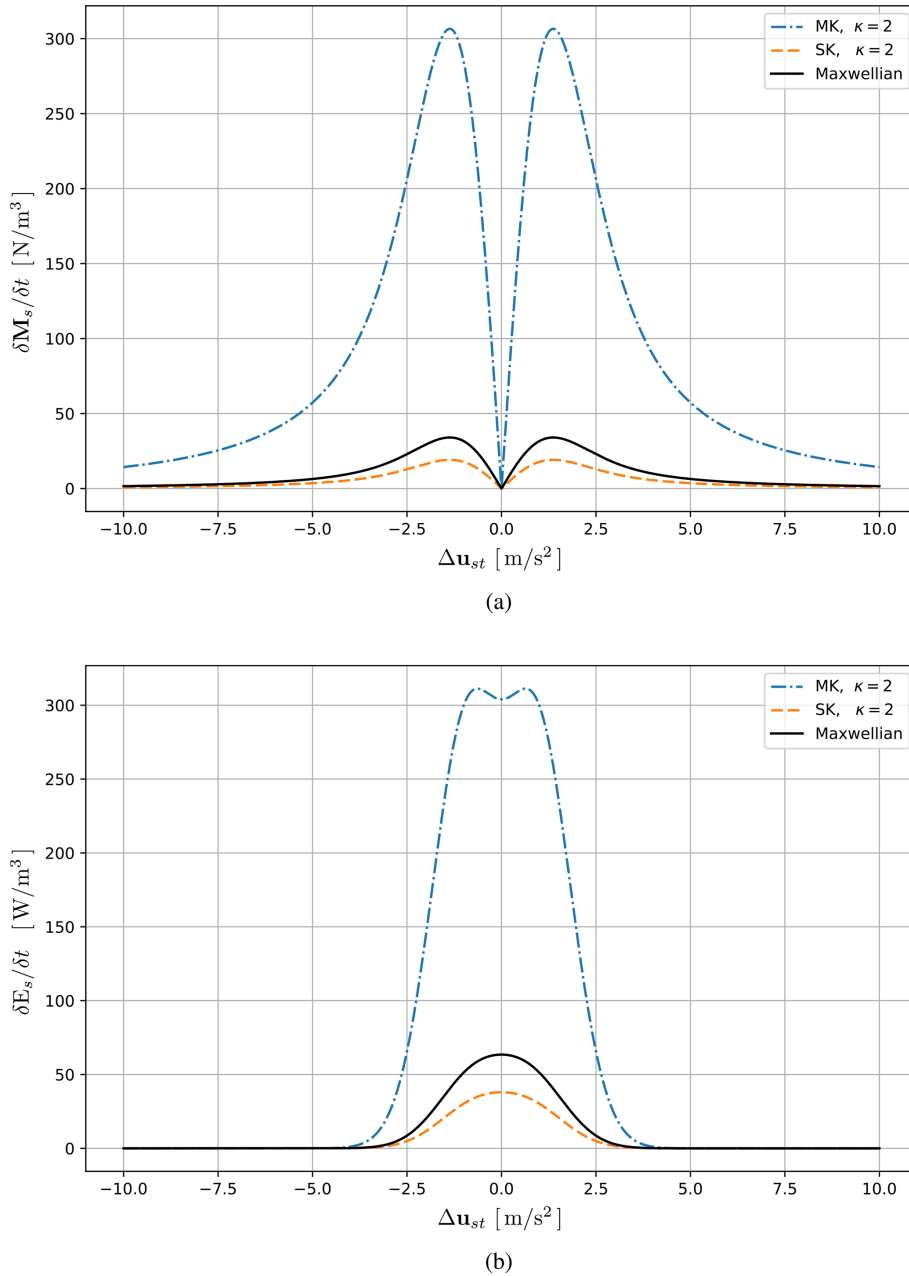


Figure 8. The cross-section of the momentum and energy collision terms for the standard Kappa and Maxwellian velocity distribution functions in the case of Coulomb collisions at $T_S = 0$.

4.3 Transport coefficients

The transport coefficients for the standard Kappa distribution exhibit a clear and systematic dependence on the kappa parameters. This dependence is primarily determined by two factors: the effective electron temperature T_e^K and the electron–ion effective collision frequency ν_{ei}^{SK} .

In particular, the kappa dependence associated with the effective temperature appears in coefficients that are directly related to the system temperature, specifically the thermo-

electric coefficient α_e and the mobility coefficient μ_e . Both of these coefficients are directly influenced by changes in the effective temperature associated with the standard Kappa distribution, where at low values of the kappa parameter, the system temperature increases compared to the Maxwellian case, leading to an increase in both the thermoelectric coefficient α_e and the mobility coefficient μ_e . Because this dependency arises from the effective temperature, it impacts the entire system regardless of the type of interaction between particles (i.e., the effect is the same for all three types of

collisions: Coulomb collisions, hard-sphere interactions, and Maxwell molecules).

The second source of kappa dependence arises from the effective collision frequency, whereas the electrical conductivity, diffusion, and mobility all reflect this dependence through ν_{ei}^{SK} , as these transport coefficients are inversely proportional to the effective collision frequency. As discussed earlier, when $\kappa = \kappa_s = \kappa_r$, the standard Kappa distribution affects the effective collision frequencies of various collision types differently. For Maxwell molecules, the effective collision frequency is identical to the Maxwellian case, and the kappa dependency of the effective collision frequency does not appear in their transport coefficients. However, for Coulomb collisions and hard-sphere interactions, the effective collision frequency decreases as κ decreases, leading to a increase in the transport coefficients at low values of κ compared to the Maxwellian case.

In Fig. 9, we show the kappa dependence of the electrical conductivity, mobility coefficient, thermoelectric coefficient, and diffusion coefficient by plotting the ratios σ_e/σ_e^M , μ_e/μ_e^M , α_e/α_e^M , and D_e/D_e^M as functions of the kappa parameter κ . Here, σ_e^M , μ_e^M , α_e^M , and D_e^M represent, respectively, the electrical conductivity, mobility coefficient, thermoelectric coefficient, and diffusion coefficient for the Maxwellian case, and they are defined as:

$$\sigma_e^M = \frac{ne^2}{m_e \nu_{ei}}, \tag{47}$$

$$\mu_e^M = \frac{k_B T_e}{m_e \nu_{ei}}, \tag{48}$$

$$\alpha_e^M = -\frac{k_B}{e}, \tag{49}$$

$$D_e^M = \frac{e}{m_e \nu_{ei}}. \tag{50}$$

Additionally, the figure also includes a comparison between the present results and those reported by Husidic et al. (2021), highlighting the influence of the kappa parameter on the transport coefficients across different studies.

While the figure shows a different dependence on the kappa parameter between the two studies, the overall behaviour is the same: at low κ values, the transport coefficients become much larger than in the Maxwellian case, and as κ increases, they approach the Maxwellian limit. This indicates that plasmas described by a standard Kappa distribution with smaller κ values conduct electric current more efficiently, allow charged particles to move more easily under electric fields, enable particles to diffuse more rapidly, and convert temperature gradients into electric current more effectively.

Another similarity between the two studies is that they exhibit the same type of kappa dependence. In both cases, the kappa dependence is identical for the electrical conductivity and mobility coefficients. Additionally, in both studies, the coefficients follow the same general order: the diffusion

coefficient shows the strongest kappa dependence, followed by the thermoelectric coefficient, and then the electrical conductivity and mobility coefficients, which have the weakest dependence. This ordering arises because the mobility coefficient depends on kappa through both the effective temperature and the effective collision frequency, whereas the thermoelectric coefficient depends only on the effective temperature. In contrast, the electrical conductivity and diffusion coefficients depend only on the effective collision frequency.

The difference in the kappa dependence between the two studies arises from the collision models used in deriving the transport coefficients. While Husidic et al. (2021) employed a Krook-type collision model, which provides a simplified representation of collisions, our work uses the full Boltzmann collision integral, which offers a more realistic description, particularly for Coulomb collisions.

For the modified Kappa distribution, the transport coefficients were derived by Jwailes et al. (2025).

At first look, we can notice that the kappa dependency due to the effective temperature T_e^κ , which appears in both the thermoelectric and diffusion coefficients for the standard Kappa distribution, Eqs. (38) and (40), does not appear in the case of the modified Kappa distribution. Making the thermoelectric coefficient unchanged compared to the Maxwellian case. This behavior arises from the fact that the modified Kappa distribution has the same effective temperature as the Maxwellian distribution.

On the other hand, the remaining transport coefficients in the case of the modified Kappa distribution, similar to the standard Kappa distribution, are influenced through the effective collision frequency. If the collision frequency is independent of particle velocity, the effective collision frequency remains unchanged, and the transport coefficients are not affected by the collision process between the interacting species. For collisions in which the collision frequency depends on particle velocity – such as Coulomb collisions or hard-sphere interactions – the effective collision frequency is affected by the modified Kappa distribution. As a result, the transport coefficients acquire a kappa parameter dependence, where the transport coefficient at small kappa values becomes smaller than in the Maxwellian case. This behavior is opposite to that of the standard Kappa distribution, where, as mentioned earlier, small kappa values increase the transport coefficient relative to the Maxwellian case.

Figure 10 illustrates this difference by plotting the electrical conductivity as a function of the reduced temperature in the case of Coulomb collisions for the Maxwellian, modified Kappa, and standard Kappa distributions. The figure shows that all three distributions exhibit the same general behavior; however, the standard Kappa distribution yields a higher electrical conductivity than the Maxwellian, while the modified Kappa distribution yields a lower value. This difference arises from the redistribution of particle velocities. At low κ , the standard Kappa distribution contains fewer particles in the low-energy core than the Maxwellian, reducing the

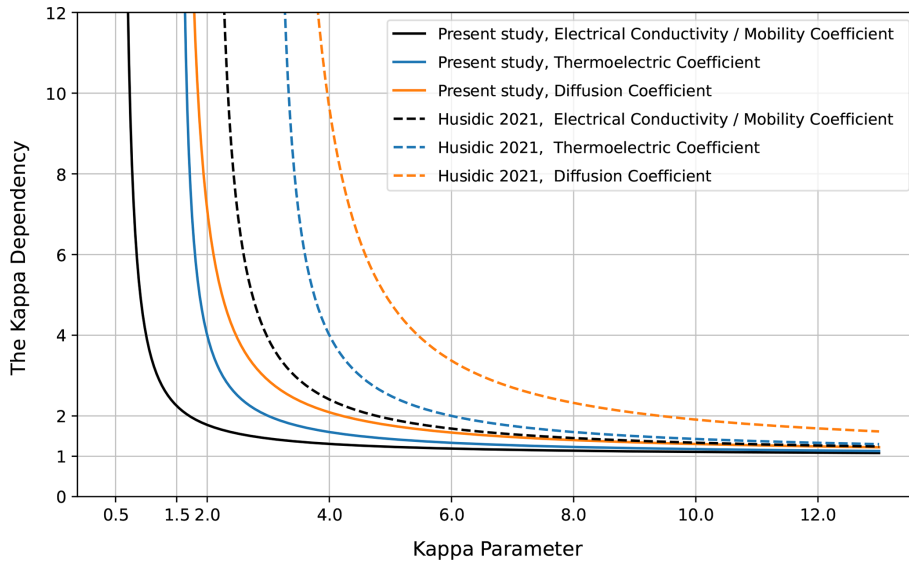


Figure 9. Dependence of the electrical conductivity σ_e , mobility coefficient μ_e , thermoelectric coefficient α_e , and diffusion coefficient D_e on the kappa parameter κ for the standard Kappa distributions in the case of Coulomb collisions.

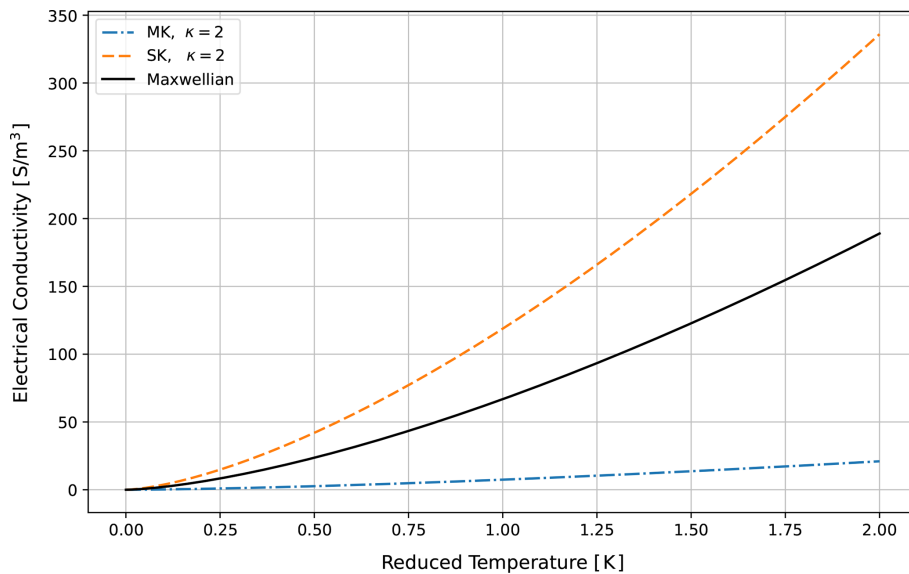


Figure 10. The electrical conductivity as a function of reduced temperature T_{st} for the Maxwellian, modified Kappa, and standard Kappa distributions in case of Coulomb collisions.

collision frequency for interactions inversely proportional to velocity, such as Coulomb collisions and hard sphere interactions. This lowers the effective collision frequency at low κ . In contrast, the modified Kappa distribution increases the population of core particles, leading to higher collision frequencies for these interactions.

The results presented in Fig. 10 are consistent with those reported in previous studies. In particular, the behavior of the standard Kappa distribution shown in the figure agrees with the findings of Husidic et al. (2021), as also illustrated in Fig. 9. For the modified Kappa distribution, similar trends

have been reported in earlier work. As discussed in Jwailes et al. (2025), the results are in good agreement with those of Ebne Abbasi et al. (2017). Both studies indicate that plasmas described by a Maxwellian distribution exhibit higher electrical conductivity than those described by a modified Kappa distribution. The reason for this agreement is that Ebne Abbasi et al. (2017) used the Fokker–Planck collision operator, which models Coulomb collisions similarly to the approach used by Jwailes et al. (2025). However, when comparing Jwailes et al. (2025) with the work of Du (2013), the same behavior is not observed. This difference arises be-

cause the collision model used in that study is not appropriate for modeling Coulomb collisions. In particular, the simplified approach does not adequately account for the enhanced thermal core of the modified Kappa distribution function; instead, the suprathermal tails dominate the trend of the electric conductivity.

A comparison of the electrical conductivity for three distributions: Maxwellian, modified Kappa, and standard Kappa made by Husidic et al. (2021) reveals a different ordering, as the one shown in Fig. 10. The modified Kappa distribution lies between the Maxwellian and the standard Kappa distributions. This is because the simplified Krook-type collision operator does not adequately distinguish between different Kappa distributions. More specifically, it underrepresents the contribution of the enhanced thermal core in the modified Kappa distribution while overemphasizing the role of suprathermal tails. Since the modified Kappa distribution has weaker tails than the standard Kappa distribution but stronger tails than the Maxwellian distribution, its electrical conductivity falls between the two – higher than the Maxwellian case but lower than the standard Kappa case.

Another point of agreement between our results and those of Husidic et al. (2021) concerns the relationship between the electrical conductivity and the mobility coefficient. Specifically, the expressions derived for the electrical conductivity and mobility, given in Eqs. (37) and (41), respectively, for the standard Kappa distribution, satisfy the well-known relation

$$\sigma_e = n_e e \mu_e, \quad (51)$$

which also holds for both the modified Kappa distribution and the Maxwellian case.

In contrast, the behavior of the thermoelectric and diffusion coefficients differs between these distributions. For the modified Kappa distribution and the Maxwellian case, these coefficients satisfy the standard Einstein relation,

$$D_e = \frac{k_B T_e}{e} \mu_e. \quad (52)$$

However, this relation does not hold for the standard Kappa distribution. Instead, a kappa dependent form of the Einstein relation is obtained:

$$D_e = \frac{k_B T_e^\kappa}{e} \mu_e. \quad (53)$$

Here, the influence of the κ parameter appears through the effective temperature. This result is consistent with the findings reported in Husidic et al. (2021). Importantly, as the kappa parameter κ approaches infinity, the effective temperature reduces to the Maxwellian one, and Eq. (53) reduces to the standard Einstein relation in Eq. (52).

The mathematical expressions and the physical trends of the collision terms and transport coefficients for the Maxwellian, standard Kappa, and modified Kappa velocity distribution functions for the three types of collisions:

Coulomb collisions, hard-sphere interactions, and Maxwell molecule interactions, discussed above, are summarized in Tables 2, 3 and 4.

5 Conclusions

For a Lorentz plasma described by a standard Kappa distribution, we derive expressions for the transport coefficients: electrical conductivity, thermoelectric, diffusion, and mobility. The analysis begins with a closed system of transport equations for isotropic plasmas within the five-moment approximation. Transport properties are defined relative to the random velocity of each species, with the velocity distribution function expanded in an orthogonal polynomial series about a drifting standard Kappa distribution. By taking only the first term and neglecting higher order moments yields the five-moment approximation. The corresponding momentum and energy collision terms are evaluated via the Boltzmann collision integral for several interaction types, including Coulomb collisions, hard-sphere interactions, and Maxwell molecule collisions. Under suitable assumptions for an unmagnetized, steady-state plasma, explicit expressions for the transport coefficients for the standard Kappa distribution are obtained from the momentum equation.

The methodology adopted in this study is broadly comparable to that of Jwailes et al. (2025), particularly in terms of the formulation of the transport equations, the evaluation of the collision integrals, and the derivation of the transport coefficients. However, a fundamental distinction between the two studies leads to markedly different physical outcomes. While Jwailes et al. (2025) employed a modified Kappa distribution function, the present work is based on the standard Kappa distribution. These two distributions differ substantially in their statistical representation of plasma particle populations, resulting in distinct plasma responses and transport properties. Although the mathematical forms of the governing equations appear similar, the physical interpretation of the quantities involved depends critically on the specific Kappa distribution adopted. This difference motivates the detailed comparative analysis presented in Sect. 4. That analysis compares three velocity distributions: Maxwellian, standard Kappa, and modified Kappa, across three stages. The first stage examined the effect of the kappa parameter on the effective collision frequency and the thermalisation rate. The second stage focused on how the kappa parameter affects the momentum and energy collision terms for Coulomb collisions. The third stage investigated the impact of the kappa parameter on transport coefficients, including electrical conductivity, diffusion, mobility, and the thermoelectric coefficient. The results of this comparison reveals that the standard Kappa distribution exhibits behavior that is qualitatively different from that of the modified Kappa distribution.

The main distinctions between the results of the standard and modified Kappa distribution functions are summarized

Table 2. Mathematical and physical comparison between the Maxwellian, standard Kappa, and modified Kappa velocity distribution functions for Coulomb collisions.

Feature/Aspect	Maxwellian (M)	Standard Kappa (SK)	Modified Kappa (MK)
Effective collision frequency	v_{st}^{Co}	$v_{st}^{Co} \left(\frac{\kappa - 1/2}{\kappa}\right)^2$	$v_{st}^{Co} \left(\frac{\kappa - 1/2}{\kappa - 3/2}\right)^2$
Effective collision frequency behavior at low κ	Baseline	Lower than Maxwellian	Higher than Maxwellian
Thermalisation rate	$v_{ei,T}^M = 2 \frac{m_{ei}}{m_i} v_{ei}^{Co}$	$v_{ei,T}^M \left(\frac{\kappa - 1/2}{\kappa}\right)^2$	$v_{ei,T}^M \left(\frac{\kappa - 1/2}{\kappa - 3/2}\right)^2$
Thermalisation rate behavior at low κ	Baseline	Lower than Maxwellian	Higher than Maxwellian
Momentum exchange at low κ	Baseline	Lower than Maxwellian	Higher than Maxwellian
Energy exchange at low κ	Baseline	Lower than Maxwellian	Higher than Maxwellian
Thermoelectric coefficient	$\alpha_e^M = -\frac{k_B}{e}$	$\alpha_e^M \left(\frac{\kappa}{\kappa - 3/2}\right)$	α_e^M
Thermoelectric coefficient behavior at low κ	Baseline	Higher than Maxwellian	Same as Maxwellian
Electrical conductivity	$\sigma_e^M = \frac{ne^2}{m_e v_{ei}^{Co}}$	$\sigma_e^M \left(\frac{\kappa}{\kappa - 1/2}\right)^2$	$\sigma_e^M \left(\frac{\kappa - 3/2}{\kappa - 1/2}\right)^2$
Electrical conductivity behavior at low κ	Baseline	Higher than Maxwellian	Lower than Maxwellian
Diffusion coefficient	$D_e^M = \frac{k_B T_e}{m_e v_{ei}^{Co}}$	$D_e^M \left(\frac{\kappa}{\kappa - 3/2}\right) \left(\frac{\kappa}{\kappa - 1/2}\right)^2$	$D_e^M \left(\frac{\kappa - 3/2}{\kappa - 1/2}\right)^2$
Diffusion coefficient behavior at low κ	Baseline	Higher than Maxwellian	Lower than Maxwellian
Mobility coefficient	$\mu_e^M = \frac{e}{m_e v_{ei}^{Co}}$	$\mu_e^M \left(\frac{\kappa}{\kappa - 1/2}\right)^2$	$\mu_e^M \left(\frac{\kappa - 3/2}{\kappa - 1/2}\right)^2$
Mobility coefficient behavior at low κ	Baseline	Higher than Maxwellian	Lower than Maxwellian

Table 3. Mathematical and physical comparison between the Maxwellian, standard Kappa, and modified Kappa velocity distribution functions for hard sphere interaction.

Feature/Aspect	Maxwellian (M)	Standard Kappa (SK)	Modified Kappa (MK)
Effective collision frequency	v_{st}^{HS}	$v_{st}^{HS} \left(\frac{\kappa - 1/2}{\kappa}\right)^2$	$v_{st}^{HS} \left(\frac{\kappa - 1/2}{\kappa - 3/2}\right)^2$
Effective collision frequency behavior at low κ	Baseline	Lower than Maxwellian	Higher than Maxwellian
Thermalisation rate	$v_{ei,T}^M = 2 \frac{m_{ei}}{m_i} v_{ei}^{HS}$	$v_{ei,T}^M \left(\frac{\kappa - 1/2}{\kappa}\right)^2$	$v_{ei,T}^M \left(\frac{\kappa - 1/2}{\kappa - 3/2}\right)^2$
Thermalisation rate behavior at low κ	Baseline	Lower than Maxwellian	Higher than Maxwellian
Momentum exchange at low κ	Baseline	Lower than Maxwellian	Higher than Maxwellian
Energy exchange at low κ	Baseline	Lower than Maxwellian	Higher than Maxwellian
Thermoelectric coefficient	$\alpha_e^M = -\frac{k_B}{e}$	$\alpha_e^M \left(\frac{\kappa}{\kappa - 3/2}\right)$	α_e^M
Thermoelectric coefficient behavior at low κ	Baseline	Higher than Maxwellian	Same as Maxwellian
Electrical conductivity	$\sigma_e^M = \frac{ne^2}{m_e v_{ei}^{HS}}$	$\sigma_e^M \left(\frac{\kappa}{\kappa - 1/2}\right)^2$	$\sigma_e^M \left(\frac{\kappa - 3/2}{\kappa - 1/2}\right)^2$
Electrical conductivity behavior at low κ	Baseline	Higher than Maxwellian	Lower than Maxwellian
Diffusion coefficient	$D_e^M = \frac{k_B T_e}{m_e v_{ei}^{HS}}$	$D_e^M \left(\frac{\kappa}{\kappa - 3/2}\right) \left(\frac{\kappa}{\kappa - 1/2}\right)^2$	$D_e^M \left(\frac{\kappa - 3/2}{\kappa - 1/2}\right)^2$
Diffusion coefficient behavior at low κ	Baseline	Higher than Maxwellian	Lower than Maxwellian
Mobility coefficient	$\mu_e^M = \frac{e}{m_e v_{ei}^{HS}}$	$\mu_e^M \left(\frac{\kappa}{\kappa - 1/2}\right)^2$	$\mu_e^M \left(\frac{\kappa - 3/2}{\kappa - 1/2}\right)^2$
Mobility coefficient behavior at low κ	Baseline	Higher than Maxwellian	Lower than Maxwellian

Table 4. Mathematical and physical comparison between the Maxwellian, standard Kappa, and modified Kappa velocity distribution functions for Maxwell molecule interactions.

Feature/Aspect	Maxwellian (M)	Standard Kappa (SK)	Modified Kappa (MK)
Effective collision frequency	ν_{ei}^{MC}	ν_{ei}^{MC}	ν_{ei}^{MC}
Effective collision frequency behavior at low κ	Baseline	Same as Maxwellian	Same as Maxwellian
Thermalisation rate	$\nu_{ei,T}^M = 2 \frac{m_{ei}}{m_i} \nu_{ei}^{MC}$	$\nu_{ei,T}^M$	$\nu_{ei,T}^M$
Thermalisation rate behavior at low κ	Baseline	Same as Maxwellian	Same as Maxwellian
Momentum exchange at low κ	Baseline	Same as Maxwellian	Same as Maxwellian
Energy exchange at low κ	Baseline	Higher than Maxwellian	Same as Maxwellian
Thermoelectric coefficient	$\alpha_e^M = -\frac{k_B}{e}$	α_e^M	α_e^M
Thermoelectric coefficient behavior at low κ	Baseline	Higher than Maxwellian	Same as Maxwellian
Electrical conductivity	$\sigma_e^M = \frac{n_e e^2}{m_e \nu_{ei}^{MC}}$	$\sigma_e^M \left(\frac{\kappa}{\kappa - 3/2} \right)$	σ_e^M
Conductivity conductivity behavior at low κ	Baseline	Same as Maxwellian	Same as Maxwellian
Diffusion coefficient	$D_e^M = \frac{k_B T_e}{m_e \nu_{ei}^{MC}}$	$D_e^M \left(\frac{\kappa}{\kappa - 3/2} \right)$	D_e^M
Diffusion coefficient behavior at low κ	Baseline	Higher than Maxwellian	Same as Maxwellian
Mobility coefficient	$\mu_e^M = \frac{e}{m_e \nu_{ei}^{MC}}$	μ_e^M	μ_e^M
Mobility coefficient behavior at low κ	Baseline	Same as Maxwellian	Same as Maxwellian

as follows: (1) The standard Kappa distribution introduces an explicit dependence on the kappa parameter κ_s through the partial pressure in the five-moment approximation system, whereas this dependence is absent in the modified Kappa distribution. (2) For velocity-independent interactions, such as Maxwell molecules, the choice of velocity distribution does not affect the collision frequency or the thermalization rate. (3) For velocity-dependent interactions, including Coulomb and hard-sphere collisions, low κ values in the standard Kappa distribution reduce the effective collision frequency, the number of collisions, and the thermalization rate, making it suitable for collisionless or weakly collisional environments, while in the modified Kappa distribution, low κ values increase these quantities, indicating its appropriateness for collisional plasmas. (4) As a result, in the case of a Coulomb collision, the standard Kappa distribution leads to weaker momentum and energy exchange compared to the Maxwellian case, whereas the modified Kappa distribution significantly enhances this exchange. (5) The standard Kappa distribution affects transport coefficients through both the effective temperature and the effective collision frequency, while the modified Kappa distribution does so only via the effective collision frequency. (6) In the Coulomb collision case, a plasma described by the standard Kappa distribution exhibits enhanced transport properties as the value of κ decreases. Specifically, lower κ values lead to more efficient electrical conductivity, greater mobility of charged particles in response to electric fields, faster particle diffusion, and a stronger conversion of temperature gradients into electric current compared to the Maxwellian case. In contrast, when

the plasma is described by the modified Kappa distribution, decreasing κ produces the opposite effect. Electrical conductivity becomes less efficient, the motion of charged particles under electric fields is more restricted, and particle diffusion is slower than in the Maxwellian case. However, the conversion of temperature gradients into electric current remains unchanged, occurring at the same rate as in the Maxwellian distribution.

While this study advances non-Maxwellian transport theory, it has several limitations. The approach relies on the five-moment approximation, retaining only the first expansion term and neglecting higher-order moments that could affect system behavior. It assumes isotropic plasmas, limiting applicability to real space environments, where magnetization and temperature or pressure anisotropies are common. The Coulomb collision cross-section is simplified using a constant logarithm and large-velocity approximation, reducing accuracy at low velocities (Fichtner et al., 1996).

Additionally, the standard Kappa distribution becomes unphysical for $\kappa \leq 3/2$, and therefore quantities derived from it, such as the collision frequency, thermalization rate, and transport coefficients, are only physically meaningful for $\kappa > 3/2$. This restriction is consistent with the mathematical behavior of the kappa dependent terms D and H , which exhibit divergences: in some cases for $\kappa \leq 1/2$, and in others for $\kappa \leq 3/2$.

Future work should address these limitations by developing a transport theory for the standard Kappa distribution via a generalized polynomial expansion, extending the theory to anisotropic plasmas, and incorporating the exact velocity-

dependent Coulomb cross-section. Alongside these theoretical developments, ongoing research is investigating transport coefficients derived from the regularized Kappa distribution (Scherer et al., 2017, 2019), which ensures finite moments and maintains thermodynamic consistency, including a detailed comparison with the results reported by Husidic et al. (2022). Additionally, comparisons with observational data – particularly from the solar wind – are crucial, with efforts focused on evaluating the predictive accuracy of both the Standard Kappa and Modified Kappa distributions against in-situ measurements.

Appendix A

The effective collision frequency for Coulomb interactions in the case of a standard Kappa distribution is given by

$$\nu_{st}^{\text{SK}} = \nu_{st} \left(\frac{\kappa - 1/2}{\kappa} \right)^2, \quad (\text{A1})$$

where ν_{st} is the effective collision frequency in the Maxwellian case.

For Coulomb collisions in a Maxwellian plasma, ν_{st} is inversely proportional to the mean free path λ_{mfp} (Livadiotis, 2019), which represents the average distance a particle travels between successive collisions. Therefore, for a standard Kappa distribution, the effective collision frequency can be expressed as

$$\nu_{st}^{\text{SK}} \propto \frac{1}{\lambda_{\text{mfp}}} \left(\frac{\kappa - 1/2}{\kappa} \right)^2. \quad (\text{A2})$$

The mean free path depends on the Debye length and is proportional to the fourth power of the Debye length divided by the Coulomb logarithm, i.e., (Livadiotis, 2019)

$$\lambda_{\text{mfp}} \propto \frac{\lambda_{\text{D}}^4}{\ln \Lambda}. \quad (\text{A3})$$

Substituting this into the expression for ν_{st}^{SK} , the effective collision frequency scales as

$$\nu_{st}^{\text{SK}} \propto \frac{\ln \Lambda}{\lambda_{\text{D}}^4} \left(\frac{\kappa - 1/2}{\kappa} \right)^2. \quad (\text{A4})$$

This scaling can be rewritten in terms of the Debye length for the standard Kappa distribution, as

$$\nu_{st}^{\text{SK}} \propto \frac{\ln \Lambda}{[\lambda_{\text{D}}^{\text{SK}}]^4}, \quad (\text{A5})$$

with $\lambda_{\text{D}}^{\text{SK}}$ is the Debye length in the case of the standard Kappa distribution, defined as

$$\lambda_{\text{D}}^{\text{SK}} = \lambda_{\text{D}} \left(\frac{\kappa}{\kappa - 1/2} \right)^{1/2}. \quad (\text{A6})$$

Accordingly, the corresponding mean free path in the standard Kappa case is

$$\lambda_{\text{mfp}}^{\text{SK}} \propto \frac{[\lambda_{\text{D}}^{\text{SK}}]^4}{\ln \Lambda}. \quad (\text{A7})$$

Thus, the effective collision frequency can also be expressed in the familiar form

$$\nu_{st}^{\text{SK}} \propto \frac{1}{\lambda_{\text{mfp}}^{\text{SK}}}. \quad (\text{A8})$$

This derivation highlights that the kappa-dependent modification of the Debye length is consistent with the results reported in Scherer et al. (2020), providing an independent validation of the physical robustness and internal consistency of the present theoretical framework.

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References

- Bittencourt, J.: *Fundamentals of Plasma Physics*, Springer New York, 3rd edn., <https://doi.org/10.1007/978-1-4757-4030-1>, 2004.
- Chapman, S. and Cowling, T.: *The Mathematical Theory of Non-uniform Gases: An Account of the Kinetic Theory of Viscosity, Thermal Conduction and Diffusion in Gases*, Cambridge Mathematical Library, Cambridge University Press, ISBN 9780521408448, 1990.
- Collier, M. R. and Hamilton, D. C.: The relationship between kappa and temperature in energetic ion spectra at Jupiter, *Geophys. Res. Lett.*, 22, 303–306, <https://doi.org/10.1029/94GL02997>, 1995.
- Collier, M. R., Hamilton, D. C., Gloeckler, G., Bochsler, P., and Sheldon, R. B.: Neon-20, oxygen-16, and helium-4 densities, temperatures, and suprathermal tails in the solar wind determined with WIND/MASS, *Geophys. Res. Lett.*, 23, 1191–1194, <https://doi.org/10.1029/96GL00621>, 1996.
- Davis, S., Avaria, G., Bora, B., Jain, J., Moreno, J., Pavez, C., and Soto, L.: Kappa distribution from particle correlations in nonequilibrium, steady-state plasmas, *Phys. Rev. E*, 108, 065207, <https://doi.org/10.1103/PhysRevE.108.065207>, 2023.
- Du, J.: Transport coefficients in Lorentz plasmas with the power-law kappa-distribution, *Phys. Plasmas*, 20, 092901, <https://doi.org/10.1063/1.4820799>, 2013.
- Ebne Abbasi, Z. and Esfandyari-Kalejahi, A.: Transport coefficients of a weakly ionized plasma with nonextensive particles, *Phys. Plasmas*, 26, 012301, <https://doi.org/10.1063/1.5051585>, 2019.
- Ebne Abbasi, Z., Esfandyari-Kalejahi, A., and Khaleidi, P.: The collision times and transport coefficients of a fully ionized plasma with superthermal particles, *Astrophys. Space Sci.*, 362, 103, <https://doi.org/10.1007/s10509-017-3081-4>, 2017.
- Fichtner, H., Sreenivasan, S. R., and Vormbrock, N.: Transfer integrals for fully ionized gases, *J. Plasma Phys.*, 55, 95–120, <https://doi.org/10.1017/S0022377800018699>, 1996.
- Formisano, V., Moreno, G., Palmiotto, F., and Hedgecock, P. C.: Solar Wind Interaction with the Earth's Magnetic Field 1. Magnetosheath, *J. Geophys. Res.*, 78, 3714–3730, <https://doi.org/10.1029/JA078i019p03714>, 1973.
- Grad, H.: On the kinetic theory of rarefied gases, *Commun. Pur. Appl. Math.*, 2, 331–407, <https://doi.org/10.1002/cpa.3160020403>, 1949.
- Guo, R. and Du, J.: Transport coefficients of the fully ionized plasma with kappa-distribution and in strong magnetic field, *Physica A*, 523, 156–171, <https://doi.org/10.1016/j.physa.2019.02.011>, 2019.
- Husidic, E., Lazar, M., Fichtner, H., Scherer, K., and Poedts, S.: Transport coefficients enhanced by suprathermal particles in nonequilibrium heliospheric plasmas, *Astron. Astrophys.*, 654, A99, <https://doi.org/10.1051/0004-6361/202141760>, 2021.
- Husidic, E., Scherer, K., Lazar, M., Fichtner, H., and Poedts, S.: Toward a Realistic Evaluation of Transport Coefficients in Non-equilibrium Space Plasmas, *Astrophys. J.*, 927, 159, <https://doi.org/10.3847/1538-4357/ac4af4>, 2022.
- Jwailes, M. J., Barghouthi, I. A., and Atawnah, Q. S.: Transport coefficients in modified Kappa distributed plasmas, *Ann. Geophys.*, 43, 783–802, <https://doi.org/10.5194/angeo-43-783-2025>, 2025.
- Lazar, M. and Fichtner, H.: Kappa Distributions: From Observational Evidences via Controversial Predictions to a Consistent Theory of Nonequilibrium Plasmas, *Astrophys. Space Sci. Library*, Springer International Publishing, ISBN 9783030826239, <https://doi.org/10.1007/978-3-030-82623-9>, 2021.
- Lazar, M., Poedts, S., and Fichtner, H.: Destabilizing effects of the suprathermal populations in the solar wind, *A&A*, 582, A124, <https://doi.org/10.1051/0004-6361/201526509>, 2015.
- Lazar, M., Fichtner, H., and Yoon, P. H.: On the interpretation and applicability of κ -distributions, *A & A*, 589, A39, <https://doi.org/10.1051/0004-6361/201527593>, 2016.
- Lima, J. A. S., Silva, R., and Plastino, A. R.: Nonextensive Thermostatistics and the H Theorem, *Phys. Rev. Lett.*, 86, 2938–2941, <https://doi.org/10.1103/PhysRevLett.86.2938>, 2001.
- Livadiotis, G.: Introduction to special section on Origins and Properties of Kappa Distributions: Statistical Background and Properties of Kappa Distributions in Space Plasmas, *J. Geophys. Res.-Space Phys.*, 120, 1607–1619, <https://doi.org/10.1002/2014JA020825>, 2015.
- Livadiotis, G.: Kappa distributions: Theory and applications in plasmas, Elsevier, <https://www.sciencedirect.com/book/edited-volume/9780128046388/kappadistributions> (last access: 21 May 2026), 2017.
- Livadiotis, G.: Kappa distributions: Thermodynamic origin and Generation in space plasmas, *J. Phys. Conf. Ser.*, 1100, 012017, <https://doi.org/10.1088/1742-6596/1100/1/012017>, 2018.
- Livadiotis, G.: Collision frequency and mean free path for plasmas described by kappa distributions, *AIP Adv.*, 9, 105307, <https://doi.org/10.1063/1.5125714>, 2019.
- Maksimovic, M., Pierrard, V., and Riley, P.: Ulysses electron distributions fitted with Kappa functions, *Geophys. Res. Lett.*, 24, 1151–1154, <https://doi.org/10.1029/97GL00992>, 1997.
- Marsch, E.: Kinetic Physics of the Solar Corona and Solar Wind, *Living Rev. Sol. Phys.*, 3, <https://doi.org/10.12942/lrsp-2006-1>, 2006.
- Mintzer, D.: Generalized Orthogonal Polynomial Solutions of the Boltzmann Equation, *Phys. Fluids*, 8, 1076–1090, <https://doi.org/10.1063/1.1761357>, 1965.
- Olbert, S.: Summary of Experimental Results from M.I.T. Detector on IMP-1, in: *Physics of the Magnetosphere*, edited by: Carovillano, R. L., McClay, J. F., and Radoski, H. R., 641–659, Springer Netherlands, Dordrecht, ISBN 978-94-010-3467-8, https://doi.org/10.1007/978-94-010-3467-8_23, 1968.
- Pierrard, V. and Lazar, M.: Kappa Distributions: Theory and Applications in Space Plasmas, *Solar Phys.*, 267, 153–174, <https://doi.org/10.1007/s11207-010-9640-2>, 2010.
- Pierrard, V., Maksimovic, M., and Lemaire, J.: Core, Halo and Strahl Electrons in the Solar Wind, *Astrophys. Space Sci.*, 277, 195–200, <https://doi.org/10.1023/A:1012218600882>, 2001.
- Qureshi, M. N. S., Pallochchia, G., Bruno, R., Cattaneo, M. B., Formisano, V., Reme, H., Bosqued, J. M., Dandouras, I., Sauvaud, J. A., Kistler, L. M., Möbius, E., Klecker, B., Carlson, C. W., McFadden, J. P., Parks, G. K., McCarthy, M., Korth, A., Lundin, R., Balogh, A., and Shah, H. A.: Solar Wind Particle Distribution Function Fitted via the Generalized Kappa Distribution Function: Cluster Observations, *AIP Conf. Proc.*, 679, 489–492, <https://doi.org/10.1063/1.1618641>, 2003.
- Scherer, K., Fichtner, H., and Lazar, M.: Regularized κ -distributions with non-diverging moments, *Europhys. Lett.*, 120, 50002, <https://doi.org/10.1209/0295-5075/120/50002>, 2017.

- Scherer, K., Lazar, M., Husidic, E., and Fichtner, H.: Moments of the Anisotropic Regularized κ -distributions, *Astrophys. J.*, 880, 118, <https://doi.org/10.3847/1538-4357/ab1ea1>, 2019.
- Scherer, K., Husidic, E., Lazar, M., and Fichtner, H.: The κ -cookbook: a novel generalizing approach to unify κ -like distributions for plasma particle modelling, *Mon. Not. R. Astron. Soc.*, 497, 1738–1756, <https://doi.org/10.1093/mnras/staa1969>, 2020.
- Schunk, R. and Nagy, A.: *Ionospheres: Physics, Plasma Physics, and Chemistry*, Cambridge Atmospheric and Space Science Series, Cambridge University Press, <https://doi.org/10.1017/CBO9780511635342>, 2009.
- Schunk, R. W.: Mathematical structure of transport equations for multispecies flows, *Rev. Geophys.*, 15, 429–445, <https://doi.org/10.1029/RG015i004p00429>, 1977.
- Tsallis, C.: Nonadditive entropy S_q and nonextensive statistical mechanics: Applications in geophysics and elsewhere, *Acta Geophys.*, 60, 502–525, <https://doi.org/10.2478/s11600-012-0005-0>, 2012.
- Vasyliunas, V. M.: A Survey of Low-Energy Electrons in the Evening Sector of the Magnetosphere with OGO 1 and OGO 3, *J. Geophys. Res.*, 73, 2839–2884, <https://doi.org/10.1029/JA073i009p02839>, 1968.
- Wang, L. and Du, J.: The diffusion of charged particles in the weakly ionized plasma with power-law kappa-distributions, *Phys. Plasmas*, 24, 102305, <https://doi.org/10.1063/1.4996775>, 2017.
- Yoon, P. H.: Electron kappa distribution and quasi-thermal noise, *J. Geophys. Res.-Space Phys.*, 119, 7074–7087, <https://doi.org/10.1002/2014JA020353>, 2014.