

# Stratospheric semi-decadal oscillations in NCEP data

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Abstract. An analysis of the National Centers for Environmental Prediction (NCEP)/National Center for Atmospheric Research (NCAR) data is presented to provide a more complete description of the stratospheric 5-year semi-decadal (SD) oscillation (Mayr et al., 2007). The zonal-mean temperature and zonal wind data from the Atmospheric Research R-1 analysis are employed, covering the years from 1962 to 2002 in the altitude range from 10 to 30 km. For diagnostic purposes, the data are separated into the hemispherically symmetric and anti-symmetric components, and spectral analysis is applied to identify the signatures of the SD oscillations. Through the synthesis or filtering of spectral features, the SD modulations of the annual oscillation (AO) and quasi-biennial oscillation (QBO) are delineated. In agreement with the earlier findings, the magnitude of the SD oscillation is more pronounced when the 30-month QBO dominates during the years from 1975 to 1995. This is consistent with results from a numerical model, which shows that such a OBO generates the SD oscillation through interaction with the 12-month AO. In the zonal winds, the SD oscillation in the NCEP data is confined to equatorial latitudes, where it modulates the symmetric AO and QBO by about 5 m/s below 30 km. In the temperature data, the effect is also seen around the equator, but it is much larger at polar latitudes where the SD oscillation produces variations as large as 2 K. Our data analysis indicates that the SD oscillation is mainly hemispherically symmetric, and it appears to originate at equatorial latitudes where most of the energy resides.

**Keywords.** Meteorology and atmospheric dynamics (Climatology; Middle atmosphere dynamics; Waves and tides)

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# 1 Introduction

The equatorial region of the middle atmosphere is controlled to a large extent by wave mean flow interactions that involve the mean zonal circulation. It was demonstrated by Lindzen and Holton (1968), Holton and Lindzen (1972) and others (e.g. Plumb, 1977; Dunkerton, 1985) that the stratospheric quasi-biennial oscillation (QBO), discovered by Reed (1965) and discussed by Pascoe et al. (2005) and Baldwin et al. (2001), can be generated with eastward propagating Kelvin waves and westward propagating Rossby gravity waves. The theory of Lindzen and Holton was applied by Dunkerton (1979) and Hamilton (1986) to simulate also the semi-annual oscillation (SAO) of the zonal circulation, which dominates in the upper stratosphere and lower mesosphere at around 50 km (Hirota, 1980). While it is generally accepted that wave mean flow interactions are required to generate the above equatorial oscillations, more recent modeling studies have shown that small-scale gravity waves (GW) are also likely involved (e.g. Hitchman and Leovy, 1988; Mengel et al., 1995; Takahashi, 1999; Hamilton et al., 2001; Giorgetta et al., 2002). That GWs are of central importance for the dynamics of the upper mesosphere had earlier been recognized by Lindzen (1981) and Dunkerton (1982).

In their seminal theory for the QBO, Lindzen and Holton (1968) emphasized that wave interactions are very effective in generating flow oscillations at equatorial latitudes. Since the Coriolis force vanishes at low latitudes (at the equator exactly), the meridional winds are barely involved in redistributing the flow momentum, and the wave source is balanced mainly by viscous dissipation. Away from the equator, the meridional circulation increasingly comes into play to dissipate and redistribute the zonal flow momentum. As a result, the QBO zonal winds peak at the equator, as observed, even in models that apply a latitude independent wave source (e.g. Mengel et al., 1995); and Haynes (1998) provided a detailed discussion of the process involved.

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For the wave forcing to be so effective at equatorial latitudes, it is essential that the zonal flow is hemispherically symmetric, and the observed QBO and SAO show that to first order. Due to critical level absorption, the wave momentum source varies with the flow velocity. Under symmetric conditions, the waves propagating up then reinforce each other to produce a coherent oscillation around the equator. When the flow is hemispherically anti-symmetric in contrast, the opposing momentum sources in the two hemispheres tend to cancel each other across the equator in the presence of horizontal diffusion.

In a recent study of the solar cycle (SC) effect (Mayr et al., 2005), it was shown that the model produced a hemispherically symmetric annual oscillation (AO), which is modulated by the SC and is confined in the zonal winds to low latitudes like the QBO and SAO. The symmetric AO is induced in the model by the anti-symmetric component of the SC forcing, and it is amplified by wave mean flow interactions around the equator.

Observed annual oscillations (AO) at low latitudes have been documented for the temperature variations (e.g. Reed and Vlcek, 1969; Pawson and Fiorino, 1998) and the zonal winds (e.g. Reddy and Reddi, 1986; Reddy et al., 1986; Reddy and Vijayan, 1993; Fleming et al., 1990). Driven to first order by the seasonal and hemispherical differences in heating, the AO is anti-symmetric to first order. This produces significant variations across the equator to mask the weaker symmetric AO that can develop at low latitudes.

Mayr et al. (2007) carried out an analysis of data supplied by the National Centers for Environmental Prediction (NCEP), which revealed a pronounced AO that is hemispherically symmetric. The NCEP data also show that this particular AO is modulated by a 5-year semi-decadal (SD) oscillation, which appears to be generated by a 30-month QBO as predicted in a 2-D modeling study (Mayr et al., 2000). Another possible source of SD oscillations is the El Niño/Southern Oscillation (ENSO), which can generate inter-annual climate variations with periods from 3 to 8 years (e.g. Wang, 1995; Trenberth, 1997; Trenberth et al., 2005).

In the present paper, we expand on the earlier analysis of NCEP data by documenting more fully the SD modulations of the AO and QBO zonal winds and temperatures. In Sect. 2, we briefly describe the data and the analysis. In Sect. 3, we discuss the spectra that describe the hemispherically symmetric and anti-symmetric components of the zonal wind and temperature data; and we show model results to provide understanding. In Sect. 3, we discuss the inferred SD modulation of the AO, which appears only in the symmetric component and extends in the temperature with significant amplitudes to the polar region. In Sect. 4, we discuss the SD modulation of the QBO that reveals similar properties. In Sect. 5, we summarize the results and present the conclusion.

# 2 NCEP reanalysis R-1 data

In our study, we employ the data from the National Centers for Environmental Prediction (NCEP)/National Center from the Atmospheric Research R-1 reanalysis (Kalney et al., 1996), which cover the years from 1958 to the present. The R-1 data extend from the surface to 10 hPa (approximately 31 km) and are spaced at 2.5 degrees in latitude. We use the zonal-mean zonal winds and temperatures.

As discussed by Kalney et al. (1996) and Kalney (2003), the R-1 data represent mainly balloon-borne radiosonde observations, which have been assimilated in GCM simulations. Nadir sounding satellite measurements introduced after 1978 improved the global coverage, but Huesmann and Hitchman (2001, 2003) have shown that the observations produced a jump in the NCEP temperatures in particular near the tropopause and at around 31 km. The NCEP data have been applied in a wide range of scientific studies, dealing for example with anthropogenic effects on the Earth' climate (Levitus et al., 2001), the solar cycle effect in the QBO temperature variations of the lower stratosphere (Salby and Callaghan, 2006), the longitudinal variations of the annular mode in the Northern Hemisphere (Thompson and Wallace, 2001), and the variability of atmospheric heat sources (Yanai and Tomita, 1998), to mention a few.

In our analysis, we delineate the NCEP data in terms of their hemispherically symmetric and anti-symmetric components. The data in the Northern and Southern Hemispheres,  $D_n$  and  $D_s$ , respectively, are split up into the symmetric (S) and anti-symmetric (A) elements by computing, respectively,  $DS=(D_n+D_s)/2$  and  $DA_n=(D_n-D_s)/2$ ,  $DA_s=(D_s-D_n)/2$ . The DS component then describes the symmetric AO with the QBO and SAO, which tend to peak at the equator, while  $DA_n=-DA_s$  describes to first order the dominant antisymmetric AO with opposite phase in the two hemispheres, which vanishes at the equator.

Applying Fourier analysis, the hemispherically symmetric and anti-symmetric spectra are derived to identify the AO, QBO, and long-term signatures in the lower stratosphere (10 to 30 km). The relevant spectral features are then synthesized (filtered) to reveal the oscillations embedded in the data.

#### 3 Zonal wind and temperature spectra

Mayr et al. (2007), referred to as MetAl, analyzed the NCEP data for the years 1965 to 2005 and 1975 to 1995. Covering a time span of 20 years, the latter data set is unique because the 30-month QBO dominates, and we employ it also in the present analysis. For the 40-year time span, we chose a somewhat different data sample from 1962 to 2002; it supports the earlier findings to produce results that are virtually identical.

In our discussion of the zonal wind spectra, for conciseness, we refer to the Figs. 1 and 3 in MetAl, which were



**Fig. 1.** Amplitude spectra for the temperatures obtained from the NCEP/NCAR R-1 data for 40- and 20-year time spans from 1962 to 2002 and 1975 to 1995, respectively, plotted versus Fourier harmonic, h, and altitude. The corresponding frequencies are h/40 and h/20 (cycles per year) and the periods are 40/h and 20/h (years). Shown are the symmetric components at the equator (**a**, **c**) and at the pole (**b**, **d**), with contour intervals of 0.05 and 0.1 K, respectively, given at the lower right of the figures. To suppress clutter, the lowest four contours are eliminated. Note the spectral features describing the 12-month AO and the dominant QBO with periods of 28.2 (a, b) and 30 (c, d) months, which are prominent also in the zonal wind data (MetAL). With arrows are identified the signatures of the 5-year semi-decadal (SD) oscillations, which modulate the AO and QBO, and they are more pronounced at the pole than at the equator. For the 20-year time span (c, d), where the 30-month QBO dominates, the SD signatures modulating the symmetric AO are larger than those for the 40-year time span (a, b).

derived, respectively, for the 40- and 20-year NCEP data sets. Applying Fourier analysis, the amplitudes of the spectra are presented in terms of discrete harmonics, h, and are plotted versus altitude from 10 to 30 km. For the 20-year data sample (MetAl, Fig. 3), the frequencies are then given by h/20 (cycles per year), and the corresponding periods are 20/h (years). This format is perhaps unusual, but it is convenient for the present purposes where we interpret the results in terms of non-linear interactions. Since the product between complex amplitudes  $Aexp[i\omega_A t] \cdot Bexp[i\omega_B t]$  produces  $Cexp[i(\omega_A \pm \omega_B)t]$ , the non-linear interactions can be identified in the spectra simply from the additions and subtractions of integers, i.e.  $\omega_A \pm \omega_B \rightarrow h_A \pm h_B$ . Although the derived spectra of the NCEP data may appear complicated, their interpretations prove to be straightforward.

In Fig. 3a (MetAl) for the symmetric zonal wind component at the equator, large amplitude signatures are apparent that describe the QBO, which is dominated by the period of 30 months at h=8. And the sharp feature at h=20 represents the symmetric 12-month AO. Of interest for the present analysis are the spectral side lobes that describe the 5-year SD modulations as indicated with solid arrows. The feature at h=16 (20–4) produces the SD modulation of the AO, and the ones at h=4 and 12 (8±4) modulate the QBO. These SD modulations are believed to originate from a SD oscillation at h=4, which is anti-symmetric as shown in Fig. 3b (MetAl) at 20 degrees latitude. Such an anti-symmetric SD oscillation could interact, through non-linear processes, with the dominant anti-symmetric AO at h=20 to generate the symmetric side lobe at h=(20-4); and it could interact with the symmetric and anti-symmetric QBO at h=8 to produce the side lobes at  $h=(8\pm4)$  as indicated. The distinct anti-symmetric spectral features at h=12 and 28 (20 $\pm$ 8), identified with dashed arrows, represent periods of about 20 and 8.6 months, respectively, which are produced by the nonlinear interaction between the dominant symmetric QBO (h=8) and the strong anti-symmetric AO (h=20).

For the analyzed 40-year time span in Fig. 1a of MetAl, the QBO is dominated by the period of about 28.2 months at h=17, and the side lobe at h=9 (17–8) is the signature of the SD modulation. The sharp feature at h=40 represents the symmetric 12-month AO, and the signature of its SD modulation appears at h=32 (40–8). As seen also in the above-discussed spectra, these modulations originate in the anti-symmetric SD oscillation at h=8 as shown in Fig. 1b (MetAl) for 20 degrees latitude. The SD signatures, however, are weaker for this 40-year data sample where the dominant period of the QBO is not 30 but 28.2 months.

Commensurate with the above-discussed QBO and AO zonal wind signatures at the equator (Figs. 1 and 3 of MetAl), the corresponding temperature signatures are also pronounced as seen from Fig. 1 of the present paper. The dominant periods of the temperature QBO are identical to those for the zonal winds. In Fig. 1a, for the 40-year time span, the spectral side lobes are identified at h=9 and 25

 $(17\pm8)$ , which modulate the QBO (*h*=17) with the SD periodicity. At the pole (Fig. 1b), the symmetric QBO signature is also prominent (note the factor of 2 larger contour intervals), and the SD side lobes are evident as well. In that figure, the signature of the symmetric AO (*h*=40) is evidently much larger than at the equator, and around 30 km a side lobe appears to indicate that a stronger SD modulation is developing at higher altitudes. In the anti-symmetric component at high latitudes (not shown), the AO is very large as expected; but there are no discernable signatures of the QBO and SD oscillation.

For the shorter time span of 20 years, with dominant 30-month QBO in Fig. 1c, the spectral features essentially mimic those evident in Fig. 1a. However, the SD modulation of the AO is now apparent also at the equator, as seen from the side lobe at h=16 (20–4). And in the symmetric component at the pole (Fig. 1d), the SD signatures for the AO modulation are more pronounced. The anti-symmetric component at the pole (not shown) again does not reveal a QBO signature, and the SD oscillation is not excited.

In an analysis of Total Ozone Mapping Spectrometer (TOMS) ozone measurements for the years 1978 to 1991, a subset of the 40-year time span in our analysis, Tung and Yang (1994) also showed that the 30-month QBO dominates, and Yang and Tung (1994) found the same in ozone data collected from Dobson stations for the years 1957 to 1986. The 60-month (SD) oscillation is also evident in their spectra. In these papers, the authors furthermore show that the data contain pronounced oscillations with periods of 20 and 8.6 months, which were interpreted to be the signatures of non-linear interactions between the QBO and AO (Hamilton, 1989). Hamilton (1995) extended the ozone spectral analysis of Tung and Yang (1994) by separating the symmetric and anti-symmetric components and found, as in the present paper, that the oscillations with periods of about 20 and 8.4 months can be tied to the QBO interacting with the AO.

In partial support of our interpretation, we present in Fig. 2 the results from a 2-D run of the mechanistic Numerical Spectral Model, which generates the 30-month QBO (Mayr et al., 2003). This model employs the Doppler Spread Parameterization (DSP) for small-scale gravity waves (Hines, 1997a, b), which produces the momentum forcing and the eddy diffusivity, K, to generate the QBO. Numerical experiments show that the amplitude and period of the QBO vary with the momentum forcing, and with K in particular, and the parameters for the present solution were chosen from the proposed range of the DSP. In Fig. 2a, the symmetric component of the zonal wind spectrum for 30 model years is shown at 4° near the equator. The QBO signature evidently dominates and is confined to the sharp spectral feature at h=12 for a period of 30 months. And in the anti-symmetric component at 20° latitude (Fig. 2b), the 12-month AO dominates at h=30. As seen from Fig. 2b, the model generates the anti-symmetric 5-year SD oscillation at h=6, which interacts apparently with the AO at h=30 to generate the symmetric



**Fig. 2.** Spectra from a 2-D mechanistic model (Mayr et al., 2003), which generates the 30-month QBO with SD oscillations. From a 30-year model run, the numerical results are presented for the zonal winds at low latitudes (**a**, **b**), and for the temperature variations near the poles (**c**, **d**). All the spectral features can be identified to provide qualitative understanding.



**Fig. 3.** Synthesized or spectrally filtered zonal winds describe the AO at 5 degrees near the equator for the hemispherically symmetric and anti-symmetric components. For the time span 1962 to 2002 (a, b), the harmonic h=40 is synthesized with the SD features at  $h=40\pm8$  (given at the left bottom) to describe the SD modulation of the AO from 10 to 30 km plotted versus time from 1980 to 1990. Although the feature at h=48 (40+8) does not appear in the spectrum for the chosen contour intervals, it is accounted for in the synthesis. The symmetric AO (a) slowly propagates down (~3 km/month) at altitudes above 15 km, and a pronounced SD modulation is evident, causing the AO amplitude to vary by almost 4 m/s near 25 km. The anti-symmetric component (b), in contrast, displays virtually no phase progression with height (i.e. high propagation velocity), and no SD modulation. For the time span 1975 to 1995 (c), the harmonics h=20 and  $20\pm4$  are employed. The SD modulation in (c) is in phase with (a), and it is significantly larger in magnitude. In the anti-symmetric component (not presented) there is again no SD modulation evident, and no phase progression.

side lobes at h=24 and 36 (30±6). Due to the non-linear interaction between QBO and AO, the periods of 20 and 8.6 months are produced. The spectral features that characterize the zonal winds at low latitudes are also excited in the temperature variations of the polar region (Fig. 2c and d),

which are generated in this 2-D model by adiabatic heating and cooling associated with the meridional circulation. Without topography and the related atmospheric variations across the globe, however, this model does not generate the symmetric AO zonal winds near the equator seen in the NCEP data at h=20 (Fig. 3a of MetAl). The model also does not reproduce the anti-symmetric QBO at h=8 (Fig. 3b of MetAl), which can interact with the anti-symmetric SD oscillation at h=8 to generate in Fig. 3a of MetAl the side lobes at  $h=(8\pm4)$  that modulate the symmetric QBO in the real world.

# 4 Semi-decadal (SD) modulation of the annual oscillation (AO)

The filtered AO zonal winds at 5° N latitude near the equator are shown in Fig. 3. For the 40-year time span from 1962 to 2002, we synthesize the harmonics h=40 with  $h=40\pm8$  to describe the 5-year semi-decadal (SD) modulations of the symmetric (Fig. 3a) and anti-symmetric (Fig. 3b) components. Although the symmetric signature at h=48 does not appear in the spectrum (Fig. 1a of MetAl) for the chosen contour intervals, it is accounted for in the analysis. In Fig. 3a, a pronounced SD modulation is evident, which peaks at about 25 km as seen also in the spectrum. At this altitude, the SD oscillation causes the zonal winds to be modulated by almost 4 m/s, with the largest amplitudes occurring between 1983 and 1984. Although the results are shown only for the years from 1980 to 1990, the displayed pattern of the SD modulation continues to repeat over the entire time span analyzed. The zonal wind pattern in Fig. 3a shows that the symmetric AO slowly propagates down to about 17 km, and the phase velocity is about 3 km/month. Slow propagation velocities are characteristic of wave-driven equatorial oscillations, like the QBO with about 1.3 km/month. For the anti-symmetric component (Fig. 3b), in contrast, the AO displays virtually no phase progression with altitude (i.e. large propagation velocity), and the SD modulation is also not evident. Wave interactions in that case are relatively ineffective.

For the shorter time span of 20 years (1975 to 1995), where the 30-month QBO dominates, the filtered SD modulation of the AO is shown in Fig. 3c. Consistent with the differences in the spectra (Figs. 1a and 3a of MetAl), the SD modulation is larger causing the zonal winds to vary by almost 6 m/s at 25 km. The peak amplitudes again occur between the years 1983 and 1984 to show that the modulations for the two time spans are essentially in phase. The anti-symmetric component (not presented) again does not reveal any phase progression with altitude or a significant SD modulation, and over the entire altitude range this 12-month AO is in phase with that from the longer time span.

As seen from Fig. 3d of MetAl, the symmetric AO zonal winds peak at the equator where the SD modulation accounts for amplitude differences of about 3 m/s over 5 years. Outside the tropics, the AO sharply increases and peaks at about 60 degrees latitude where the winds reach amplitudes of about 14 m/s without being modulated much by the SD oscillation. A close inspection of the zonal winds at this latitude reveals that the weak SD modulation also seems to vary erratically since the patterns for the 40- and 20-year time spans

are out of phase. This further supports the picture that, for the zonal winds, the SD modulation of the AO is confined to equatorial latitudes.

In the temperature variations, the situation is completely different, and this is shown in Fig. 4a and b, where we present at the equator and pole the height variations of the filtered symmetric AO temperature variations for the time span from 1975 to 1995. As in Fig. 3, the oscillations are generated with the spectral features h=20 and  $20\pm4$ . At the equator (Fig. 4a), with contour intervals of 0.5 K, a weak SD modulation of about 0.5 K is apparent above 20 km where the amplitude of the AO itself is comparatively small. At the pole (Fig. 4b), with contour intervals of 1 K, the AO amplitude is close to 8 K, and the SD modulation is then barely discernable although it is much larger than at the equator. The background of the large polar AO amplitude is evidently hiding the SD modulation, and in Fig. 4c and d, we present an analysis that addresses this problem.

In this analysis, we portray the modulation patterns specifically and separate from the oscillation that is being modulated. The approach is illustrated by applying it to the time series of the large symmetric AO at polar latitudes, which is presented at 31 km in Fig. 4c. Figure 4c shows the filtered variations of the AO, with positive phase (solid) and negative phase (dashed), which describe the SD modulation at this altitude. This reveals that the modulation causes the AO to vary from about 5.5 to 7.5 K. The envelope of this modulation (solid line) is computed from the maxima of the positive and negative AO temperature excursions, and the dotted line at about 6.5 K represents the time average of the computed envelope. The deviations of this envelope from the mean, about  $\pm 1$  K, thus describe the magnitude of the modulation. Applying this procedure, we then construct the contour plot (Fig. 4d) that displays the SD modulation of the temperature AO presented in Fig. 4b. It shows that the symmetric AO at 30 km is modulated by about 2 K, with the maximum (positive) occurring between 1982 and 1983 and the minimum (negative) at 1985. The modulation is also shown propagating down and thus displays in concise form the SD pattern, unaffected by the large magnitude of the AO amplitude.

The above analysis is applied in Fig. 5 to describe the symmetric SD modulations of the temperature data. For the years 1962 to 2002, the harmonic h=40 is synthesized together with  $h=40\pm 8$  for the equator (Fig. 5a) and pole (Fig. 5b). In Fig. 5a, the SD modulation at 30 km produces variations of about  $\pm 0.25$  K, amounting to temperature differences of more than 0.5 K, with the maximum occurring after 1983. The modulation varies with altitude and produces smaller variations below 20 km. At the pole (Fig. 5b), the SD is larger to yield values of  $\pm 0.40$  K near 25 km, revealing a pattern of height variation that is completely different from that displayed at low latitudes (Fig. 5a).

For the shorter time span from 1975 to 1995, where the 30-month QBO dominates, the harmonics analyzed are h=20 with  $h=20\pm 4$ , as in Fig. 4, and the results are presented in



**Fig. 4.** Similar to Fig. 3c but for the synthesized symmetric temperature variations at the equator (**a**). The SD modulation above 25 km is about 0.5 K, and it is distinguishable since the AO amplitude itself is relatively small. At the pole (**b**), however, the amplitude of about 7 K is relatively large, and the magnitude of the modulation is not discernable. When the background oscillation is large, its weak but measurable amplitude modulation must be extracted separately, and the analysis for that is discussed with the subsequent figures (c, d). (**c**) With reduced temperature excursion, the envelope of the AO (solid line) describes the SD modulation (varying from about 5.4 to 7.6 K), and the dotted line at 6.5 K represents the time average. The deviations of the envelope from the mean are applied to construct the contour plot in (**d**), which describes in concise form the SD modulation of the AO presented in (b). Without being masked by the large AO amplitude that is effectively hiding the SD in (b), the contour plot in (d) shows that the modulation amounts to variations of about 2 K at 30 km, with the maximum occurring between 1982 and 1983 and the minimum at 1985. It also shows that the magnitude of the SD modulation is largest at 30 km and propagates down to lower altitudes.

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**Fig. 5.** Applying the analysis discussed in Fig. 4c and d, the synthesized (filtered) SD modulations of the 12-month AO are presented for the equator and pole based on data covering the years 1962 to 2002 (a, b) and 1975 to 1995 (c, d). The contour interval is 0.1 K. (a) For the 40-year time span at the equator, the SD temperature modulation produces variations of about  $0.5 \text{ K} (\pm 0.25 \text{ K})$  at 30 km, and somewhat smaller and more isolated variations close to 20 km. At the pole (b), the modulation amplitude is larger to generate variations of about 0.8 K near 25 km. For the shorter time span of 20 years, the SD temperature modulation at the equator (c) reveals a pattern that resembles the one shown in (a). The amplitude and phase at 30 km almost agree, and the modulation is also shifted to earlier times at lower altitudes. However, the modulation pattern at the pole (d) is significantly different from that displayed in (b), and its magnitude is much larger at altitudes above 20 km.

Fig. 5c for the equator and in Fig. 5d for the pole (identical to Fig. 4d). At the equator, the SD pattern in Fig. 5c resembles the one for the longer time span in Fig. 5a. The modulations almost agree in amplitude and phase at 30 km, and the height variations in both cases produce patterns in which the maxima occur after 1980 at altitudes below 20 km. For the pole, however, the displayed SD modulation in Fig. 5d is qualitatively different from that shown in Fig. 5b. Unlike the pattern for the longer time span, the SD produces relatively large temperature variations of about 2 K at 30 km, but the modulations decay more rapidly at lower altitudes.

# 5 Semi-decadal (SD) modulation of the quasi-biennial oscillation (QBO)

Considering the spectral features that describe the SD modulations of the QBO zonal winds in Fig. 1a of MetAl, we present with Fig. 6 the synthesized variations for the years 1962 to 2002 near the equator. For this 40-year time span, with dominant QBO of 28.2 months, the harmonic h=17 is synthesized with  $h=17\pm8$  to produce in Fig. 6a the hemispherically symmetric oscillations plotted versus altitude from 15 to 30 km, applying contours intervals of 2 m/s. The resulting alternating wind velocities are close to 15 m/s and show a pronounced SD modulation. To reveal that modulation explicitly, we present in Fig. 6b the results from our analysis above discussed. Instead of computing the AO modulation as in Fig. 4d, the analysis is applied to the QBO to produce the contour plot in Fig. 6b. This shows that the SD oscillation modulates the QBO at altitudes close to 24 km by about  $\pm 2.5$  m/s to produce variations of about 5 m/s at that altitude. Though smaller in magnitude, the variations are also significant at 30 km. For comparison, we present in Fig. 6c the anti-symmetric component of the QBO wind velocities at 5 degrees latitude. With contour interval of 0.1 m/s, the amplitudes are small compared to those shown in Fig. 6a. Moreover, the displayed pattern reveals that there is virtually no phase progression with altitude in the anti-symmetric component of the QBO, much like the AO presented in Fig. 3b. The QBO results, like those for the AO, thus emphasize again the close connection between hemispherical symmetry and the efficient wave forcing with slow phase progression at equatorial latitudes. The analysis results for the 20-year time span (not presented), where the 30-month QBO dominates, reveal that the amplitudes are larger, close to 20 m/s at around 26 km. But the SD modulation of the QBO is somewhat weaker. At 24 km, the SD modulation causes the amplitudes to vary by about 4 m/s, and the modulation maximum occurs close to 1981 where it also occurs in Fig. 6b. However, at higher altitudes where the effect is weaker, the phase of the SD modulation for the 20-year time span is almost opposite to that shown in Fig. 6b.

For the QBO temperature variations, we present in Fig. 7 the SD modulations derived from the 40-year data span cov-

ering the years 1962 to 2002. Employing the harmonics h=9, 17, 25 as in Fig. 6, the synthesized or filtered variations for the symmetric component are shown at the equator in Fig. 7a. The amplitudes are about 1.5 K near 25 km where the temperature QBO peaks. The relatively small SD modulation is then presented in Fig. 7b, which reveals variations between minimum and maximum amplitudes as large as 1 K at 30 km and smaller values of about 0.6 K at about 18 km altitude. In the polar region (Fig. 7c), the SD modulation at 18 km produces amplitude differences of about 1 K as seen also from Fig. 7d. Unlike the QBO temperature variations at the equator, the SD modulated oscillations extend down to 10 km and below.

To reveal the latitudinal temperature variations, Fig. 8a shows near 18 km a synthesis of the harmonics h=9, 17, 25,which represent the symmetric and anti-symmetric (or total) spectral features that characterize the SD modulation for the 40-year time span covering the years 1962 to 2002. Around the equator, the maximum and minimum temperature amplitudes are, respectively, close to 1.2 K and 0.6 K to produce a SD modulation of about 0.6 K. At the poles in the Northern and Southern Hemispheres, the SD temperature signatures are different to indicate that the anti-symmetric component is important. As seen at the poles in the two hemispheres, symmetry is also on display for example in the negative and positive phase patterns during the years 1981 and 1982, respectively, which are opposite in phase to the equatorial oscillations. For the years 1975 to 1995 in Fig. 8b, the harmonics h=4, 8, 12 are employed to describe the SD modulation of the 30-month QBO that dominates in this case. Figure 8a and b shows that the maxima and minima in the QBO temperature variations around the equator are essentially in phase. In the polar regions, the temperature patterns also reveal similarities, but the differences are larger.

#### 6 Summary and conclusion

An initial analysis of the NCEP zonal wind and temperature data in the stratosphere revealed a pronounced component of the AO, which is hemispherically symmetric and is modulated by a 5-year semi-decadal (SD) oscillation at equatorial latitudes as reported in Mayr et al. (2007, referred to as MetAl). Observational evidence indicates that the SD oscillation is generated by the 30-month QBO. In the present paper, we describe more fully the SD modulations that appear in the symmetric AO and QBO and extend in the temperature to the polar regions.

As in MetAl, the data analysis is applied to time spans of 40 years (1962 to 2002) and 20 years (1975 to 1995), with the latter covering a period where the 30-month QBO dominates. For these time spans, the amplitude spectra of the symmetric zonal winds in Figs. 1a and 3a of MetAl reveal pronounced SD signatures that modulate the AO and QBO. These features originate apparently in the SD oscillations of



**Fig. 6.** Synthesized height variations for the QBO zonal winds and their SD modulations near the equator for the years 1962 to 2002. The harmonics h=17 with  $17\pm8$  describe the SD modulation of the dominant 28.2-month QBO. For the symmetric component (**a**), the contour intervals are 2 m/s. The maximum amplitudes are close to 15 m/s, and the SD modulation is apparent. Applying the analysis discussed with Fig. 4c and d, but for the QBO instead of AO, the computed SD modulation is presented in (**b**). This shows that the wind amplitudes vary over a period of 5 years by as much as 5 m/s at about 24 km, with the maximum occurring close to the year 1981. At higher altitudes, the modulation is weaker but still significant. For the small anti-symmetric component near the equator with contour intervals of 0.1 m/s (**c**), the QBO displays virtually no phase progression with height, like the AO in Fig. 3b.

the anti-symmetric wind component evident in Figs. 1b and 3b of MetAl. As shown in MetAl (Figs. 1c, 3d), the SD modulated symmetric AO wind component peaks at the equator. Outside the tropics, the winds then increase to reach the primary maximum at about 60° latitude; but the SD modulations at that latitude are small. The temperature spectra (Fig. 1) also contain pronounced signatures of the 12-month AO and QBO. In contrast to the zonal winds, however, the symmetric temperatures exhibit SD oscillations at higher latitudes and in the polar regions (Fig. 1b and d), where the spectral signatures are about a factor of two larger than those at the equator (note the differences in contour intervals). In the anti-symmetric



**Fig. 7.** Synthesized height variations of the QBO temperature variations and their SD modulations for the years 1962 to 2002. The harmonics h=17 with  $17\pm8$  are employed. For the equator, we present in (a) the QBO height variations and in (b) their SD modulations, which reveal amplitude differences exceeding 0.5 K at about 18 km and still larger values at 30 km. Analogous to (a, b), the QBO temperature variations at the pole are shown in (c) and their embedded SD modulations in (d), which exhibit amplitude variations of about 1 K near 20 km.



**Fig. 8.** Latitudinal variations of the synthesized QBO temperature variations near 18 km altitude, describing the SD modulation for the time spans 1962 to 2002 (a) and 1975 to 1995 (b). In the latter shorter time span, the 30-month QBO dominates. The symmetric and anti-symmetric (total) spectral features are employed for the harmonics h=9, 17, 25 (a) and h=4, 8, 12 (b), and the contour interval is 0.3 K. Around the equator, the magnitude of the SD modulation accounts for amplitude differences of about 0.5 K over a period of 5 years. At the poles, the differences are about a factor of two larger. While the polar regions in the two hemispheres show differences to indicate that the anti-symmetric component is important, symmetry is also on display for example during the years 1981, 1982 and 1986, 1987, where the oscillations are in phase in both hemispheres.

temperature component of the polar region (not presented), the 12-month AO is very large, but there are no signatures of the QBO and SD oscillation.

The pivotal role of wave forcing evident in the NCEP data is illustrated in Fig. 3, where we present the synthesis for the harmonics that describe the SD modulations of the 12-month AO. At altitudes above 17 km, the symmetric AO slowly propagates down with a phase speed of about 3 km/month, and the SD modulation accounts for amplitude variations as large as 5 m/s. In the anti-symmetric component, in contrast, there is virtually no phase progression evi-

dent owing to the lack of effective wave forcing in that case, and the SD modulation is weak. These characteristics of the NCEP data are in qualitative agreement with the model results from a solar cycle (SC) study (Mayr et al., 2005), which shows that the modulated symmetric AO zonal winds propagate down through the stratosphere with a velocity of about 3 km/month, thus carrying the SC signature to lower altitudes. The anti-symmetric AO component from the model also does not reveal a phase progression with altitude, and not a SC modulation either. In the NCEP data, a relatively weak SD modulation is seen in the symmetric AO temperature variations at equatorial latitudes (Fig. 5a and c). But the SD modulation is sizable in the polar regions (Fig. 5b and d), where it can account for amplitude variations between 1 and 2 K in the lower stratosphere. The high-latitude temperature signatures are apparently generated by the meridional circulation and planetary waves, which redistribute the energy globally to focus it partially onto the polar regions. Although the temperature signatures in the data and model results are larger at the poles, we believe that they originate in the equatorial region where most of the energy resides and where the wave forcing is most effective in amplifying the oscillations.

This picture of the observed symmetric AO emerges also from the NCEP results for the QBO. The symmetric QBO zonal wind oscillation near the equator (Fig. 6), with velocities close to 15 m/s, slowly propagates down at a rate of about 1.2 km/month, and the SD modulation peaks at about 24 km where the 5-year amplitude variation amounts to about 4 m/s. As in Fig. 3b for the AO, the anti-symmetric QBO in Fig. 6c does not reveal a phase progression with altitude, absent efficient wave forcing, and for the same reason, the amplitudes are small (note the small contour interval of 0.1 m/s). In the temperature variations of the QBO (Figs. 7 and 8), like those for the AO (Fig. 5), the symmetric SD modulations are larger at the poles than at the equator. This is again consistent with our contention that the energy, redistributed by the meridional winds and planetary waves, is partially focused onto the polar regions so as to increase the magnitude of the temperature oscillations, which is shown also in the model results (Mayr et al., 2000).

It was demonstrated with a mechanistic 2-D model (Mayr et al., 2000, 2003) that the SD oscillation in the stratosphere can be generated by a 30-month QBO interacting with the dominant anti-symmetric 12-month AO, and gravity wave node filtering was proposed for the mechanism involved. As shown in Fig. 2, such a model reproduces qualitatively the major spectral features seen in the NCEP data. Mayr et al. (2000) argued that the 30-month QBO appears to be unique in that it is optimally synchronized by the SAO that acts as pacemaker. This synchronization apparently accounts for the stability of the 30-month QBO, evident in the sharp spectral features in Fig. 2, and it may account for the observation that this QBO periodicity and its SD spin-off are important features of the stratospheric climatology.

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