

# Remote sensing of aerosols over the oceans using MSG/SEVIRI imagery

F. Thieuleux<sup>1,2</sup>, C. Moulin<sup>1</sup>, F. M. Bréon<sup>1</sup>, F. Maignan<sup>1</sup>, J. Poitou<sup>1</sup>, and D. Tanré<sup>2</sup>

<sup>1</sup>IPSL/LSCE, CEA-CNRS, Gif-sur-Yvette, France <sup>2</sup>LOA, CNRS-USTL, Villeneuve d'Ascq, France

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Abstract. The SEVIRI instrument on board Meteosat Second Generation (MSG) offers new capabilities to monitor aerosol transport over the Atlantic and the Mediterranean at high temporal and spatial resolutions, in particular, Saharan dust from North Africa, biomass-burning aerosols from subtropical Africa and pollution from Europe. An inversion technique was developed to estimate both aerosol optical thickness and Angström coefficients from SEVIRI measurements at 0.63 and 0.81  $\mu$ m. This method relies on an optimized set of aerosol models to ensure a fast processing of full-resolution MSG images and to allow the processing of long time series. SEVIRI images for slots 45, 49 and 53 (11:15, 12:15, 13:15 UT) were processed for June 2003. The retrieved optical thicknesses and Angström coefficients are in good agreement with AERONET in-situ measurements in the Atlantic and in the Mediterranean. Monthly mean maps of both parameters are compared to that obtained with the polar orbiting sensor POLDER for June 2003. There is a good consistency between the two monthly means in terms of optical thickness, but the Angström coefficients show significant differences in the Atlantic zone which is affected by dust transport. These differences may be explained by the lack of specific non-spherical dust models within the inversion. The preliminary results presented in this paper demonstrate, nevertheless, the potential of MSG/SEVIRI for the monitoring of aerosol optical properties at high frequencies over the Atlantic and the Mediterranean.

**Keywords.** Atmospheric composition and structure (Aerosols and particles; Instruments and techniques)

# 1 Introduction

Aerosols contribute significantly to the Earth's radiative budget, both directly and through interactions with cloud optical properties. Their concentrations must be monitored at the global scale to develop and evaluate climate models that take into account the impact of these airborne particles. Because of the strong variability of aerosol sources, distributions and optical properties, a monitoring at high spatial and temporal resolutions is needed. Satellites are well suited for this objective (Kaufman et al., 2001). Several space-borne sensors dedicated to aerosol studies have been launched during the last decade, such as POLDER in 1996 and 2003 or MODIS in 1999 and 2001. Daily images of the METEOSAT VIS band were used to monitor almost ten years of African dust transport over the Atlantic and the Mediterranean (Moulin et al., 1997a). The images were first calibrated (Moulin et al., 1996; Moulin and Schneider, 1999) and a Look-Up Table (LUT) of TOA reflectances from radiative transfer simulations was used to retrieve the dust optical thickness with an accuracy of 25% (Moulin et al., 1997b and c).

This method was applied to the entire set of noon images of the ISCCP-B2 archive between 1983 and 1997. This unequalled data set was used to generate the first monthly climatology of African dust loads over the Atlantic and the Mediterranean and to show that the intensity of dust export is closely related to climate parameters, such as the North Atlantic Oscillation (Moulin et al., 1997c; Chiapello and Moulin, 2002). The major shortcoming of this inversion method is that a single spectral band does not allow one to distinguish between various aerosol types and origins from their optical properties. As a consequence, the LUTs are based on a single aerosol model, whereas several types of aerosols are present in the marine atmosphere, each one with different optical properties.

The characteristics of the SEVIRI sensor on board MSG (this sensor will be hereafter referred to as MSG for simplicity), greatly improved from the previous Meteosat series, make it a well suited tool to study aerosols over the oceans and seas surrounding the African continent. Contrary to ME-TEOSAT, MSG has three narrow spectral bands in the solar spectrum (at 0.63, 0.81 and 1.64  $\mu$ m), in addition to the wide HRV band, and combines thus somewhat the advantages of a multi-spectral sensor, such as MODIS or POLDER, with

the frequency of METEOSAT measurements. In this paper, we make use of the two solar bands at 0.63 and 0.81  $\mu$ m to retrieve both the optical thickness and the spectral signature of the aerosols over the oceans. The results are evaluated against both sunphotometer measurements and another satellite product derived from a polar orbiting sensor.

# 2 Method

#### 2.1 Radiative transfer computations

Over a cloud-free area, the signal measured in the solar spectrum by a space-borne sensor is the sum of the light backscattered towards the satellite by the surface, the atmosphere molecules and the aerosols. The contribution of aerosols to MSG Top-Of-Atmosphere (TOA) reflectances is simulated in both 0.63 and 0.81  $\mu$ m bands with the radiative transfer model 6S (Vermote et al., 1997) for various geometries and 15 different aerosol optical thicknesses at 0.55  $\mu$ m ( $\tau$ ).  $\tau$  was varied from 0 for an aerosol-free atmosphere, i.e. with only molecular reflectance, to 2 for a very turbid atmosphere. The 6S model computes all orders of scattering by aerosols and molecules. It also accounts for gaseous absorption, which is mostly affected by O<sub>3</sub>, H<sub>2</sub>O and O<sub>2</sub> in the two MSG spectral bands. The spectral integration of reflectances over the satellite band is performed with a step of 2.5 nm.

Here we use the pre-defined US62 standard atmosphere, which includes an ozone column of 0.344 cm-atm and a water vapor column of  $1.42 \,\mathrm{g}\,\mathrm{cm}^{-2}$ . Note that a more accurate correction could be done by using daily meteorological fields from numerical global models, as within most operational processing (e.g. POLDER, MODIS, etc.). We used the other pre-defined atmospheres of 6S (i.e. tropical, midlatitude summer and winter, and high-latitude summer and winter) to verify that the use of the US62 model does not yield too much uncertainty to our retrievals. The mean error on the TOA reflectance at 0.81  $\mu$ m, and thus also on the optical thickness, is of about 5%, with a maximum of about 7% difference with the tropical model. Errors are smaller (about 3 and 5%, respectively) at 0.63  $\mu$ m. We also assume that the water is non reflecting in both bands, so that the surface contribution is only generated by specular reflection and affects the measurements only in directions close to the glint.

Aerosol optical properties over the Atlantic and the Mediterranean may vary widely depending on the size distribution and refractive index of the particles, which, in turn, depend on the source and nature of the aerosols. The Angström coefficient ( $\alpha$ ) is an optical property that can be used as a proxy for the aerosol size distribution. It is computed from the spectral variation of  $\tau$  between two wavelengths and is large (up to 2.5) for very small aerosols, like sulfates from polluted areas, and is low ( $\approx$ 0) for aerosols containing large particles as desert dust. Here we define  $\alpha$  from the ratio of the retrieved optical thicknesses at 0.63 and 0.81  $\mu$ m. Both insitu and satellite measurements in the Mediterranean have shown that the Angström coefficient

computed in the visible Spectrum varies roughly between 0 and 2 (Holben et al., 2001; Jamet et al., 2004). Fifteen aerosol models with an Angström coefficient ranging from -0.1 to 2.4 were thus used to compute TOA reflectances and to build the LUT. These 15 models include the 12 models developed for the atmospheric correction of SEAWIFS imagery (Gordon and Wang, 1994; McClain et al., 2000). They have a bimodal size distribution with variations of the respective proportions of the troposheric (small particles) and oceanic (large particles) modes of Shettle and Fenn (1979), and the relative humidity, which mostly modifies the mean radius of the two elementary modes. Since this set covers the Angström coefficient up to 1.5 only, it was complemented with three additional models of very small particles to cover the range 1.8–2.4. Table 1 provides details on each aerosol model.

A LUT of TOA reflectances at 0.55  $\mu$ m, as well as for the two MSG shortwave bands (0.63 and 0.81  $\mu$ m), was computed with 6S for the 15 aerosol models, for 15 aerosol optical thicknesses and for about 110 000 different measurement geometries: 48 values for both solar zenith angle  $\theta_S$  and viewing zenith angle  $\theta_V$  (varied between 0° and 90° using a Gaussian quadrature), and 49 values for the relative azimuth angle  $\Delta \phi$  between 0 and 180°. A useful property of the selected aerosol model set is that the simulated reflectances increase continuously from the largest to the smallest size distribution for a given optical thickness, regardless of the wavelength or the viewing geometry. As shown by Gordon and Wang (1994), this feature strongly speeds up the search for the most suitable aerosol model, because it is thus possible to interpolate the aerosol reflectance between two consecutive models without searching the whole set of models.

#### 2.2 Image processing

In terms of image processing, we followed approximately the method for METEOSAT (Moulin et al., 1997b). The first step is to convert the numerical counts to TOA reflectances using the calibration coefficients provided by EUMETSAT (Govaerts and Clerici, 2004). The next step is to remove cloud-contaminated pixels over the ocean using the spatial variability of the TOA reflectance at 0.81  $\mu$ m, with aerosol plumes being much more homogeneous than clouds. Pixels with a standard deviation of the TOA reflectance, computed on an area of  $3 \times 3$  pixels (centred on the considered pixel) greater than 0.0045, are considered as cloudy. This value has been chosen by trial and error, by visual examination of several true-colour composite images. In order to avoid any cloud contamination of our aerosol product, an additional test removes the pixels located next to a cloudy pixel (as detected by the test on the local variance). Note that the MSG resolution is 3 km at nadir (i.e. at the crossing of the equator and of the Greenwich meridian).

When it is a significant contribution to the signal, the Sun-glint prevents an accurate retrieval of the aerosol optical properties. The size of the contaminated area in the MSG imagery depends strongly on the wind speed and direction,

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**Table 1.** Characteristics of the aerosol models used in this study. They are all made of a mixing of the oceanic and tropospheric lognormal modes of Shettle and Fenn (1979), except for the W01, W02 and W03 models. These three models were derived from the tropospheric mode at a relative humidity of 0% by decreasing the mean radius from 0.027  $\mu$ m (original value) to 0.021, 0.017 and 0.015  $\mu$ m, respectively. Note that the Angström coefficient is computed between 0.63 and 0.81  $\mu$ m and that the contribution of both Oceanic and tropospheric modes are in number (not in volume or in surface).

Model	Туре	Relative Hum. (%)	Contribution (%) of the	Contribution (%) of the	Angström Coeff.	
			Oceanic mode	Tropospheric mode		
O99	Oceanic	99	100	0	-0.09	
M99	Maritime	99	1	99	0.08	
C99	Coastal	99	0.5	99.5	0.21	
M90	Maritime	90	1	99	0.22	
C90	Coastal	90	0.5	99.5	0.42	
M70	Maritime	70	1	99	0.43	
M50	Maritime	50	1	99	0.51	
C70	Coastal	70	0.5	99.5	0.66	
C50	Coastal	50	0.5	99.5	0.78	
T99	Tropospheric	99	0	100	1.29	
T90	Tropospheric	90	0	100	1.49	
T50	Tropospheric	50	0	100	1.61	
W01	Tropospheric	_	0	100	1.79	
W02	Tropospheric	_	0	100	1.97	
W03	Tropospheric	_	0	100	2.25	

a piece of information that is not available within our processing, contrary to what is done for the operational processing of sensors like POLDER and MODIS. In order to avoid any possible Sun-glint contamination, we discard all pixels whose line-of-sight makes an angle less than 30° with the specular direction. Finally, pixels with geometries which are too extreme ( $\theta_S$  or  $\theta_V$  greater than 75°) are also discarded because they are out of the validity range of the 6S radiative transfer model.

The last step of the image processing is the inversion itself. The LUT is first used to estimate the aerosol optical thickness that corresponds to the 0.81  $\mu$ m TOA reflectance for each of the 15 aerosol models, accounting for the actual observation geometry by linearly interpolating between the tabulated values of the three angles  $\theta_S$ ,  $\theta_V$  and  $\Delta\phi$ . The LUT is further used to interpolate the corresponding theoretical TOA reflectance at 0.63  $\mu$ m for each of the 15 aerosol models. The model that reproduces the best TOA reflectance at this shorter wavelength, which was not used to compute the optical thickness, is selected.

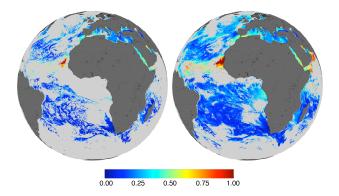
The final result of the inversion algorithm is thus this aerosol model, which is associated with a given Angström coefficient, and the aerosol optical depth extrapolated at 0.55  $\mu$ m, a standard wavelength for the inter-comparison of aerosol load products. This latter spectral extrapolation is done once again by interpolating on the geometry in the LUT the TOA reflectances at 0.55  $\mu$ m for the selected aerosol model. Note, however, that when the aerosol load is very low, the satellite measurement is too small to properly assess its spectral signature. In such a case (i.e. whenever the

retrieved optical thickness is lower than 0.07), the O99 model is utilized and a new optical thickness is derived. This model was selected because it is representative of marine aerosols encountered over the remote oceans, where the smallest optical thicknesses are mostly expected.

The results of this processing are maps of both the aerosol optical thickness and the Angström coefficient. Note that by default the optical thickness is computed at  $0.55 \,\mu$ m, but that it is possible to select another wavelength for comparison with other measurements. In that case, the optical thickness is then simply extrapolated from that at  $0.55 \,\mu$ m, using the retrieved Angström coefficient. The developed computer code is fast enough to enable the processing of several slots per day for long periods of time: it takes about 24h to process one month of MSG data with 9 slots per day on a standard mono-processor PC/Linux.

#### 3 Results

Figure 1 compares daily maps of aerosol optical thickness obtained using one single slot and combining the estimates from nine successive slots. The use of several images per day yields a significant improvement of the spatial coverage for three main reasons: 1) the area affected by the Sunglint (located in the Gulf of Guinea in the single slot image) moves from east to west during the day and does not affect all slots; 2) it fills a large fraction of the cloudy areas where the cloud coverage is fractional (mainly in the tropics), because clouds move away between successive acquisitions; 3) it enables to process pixels that are not sunlit on one particular

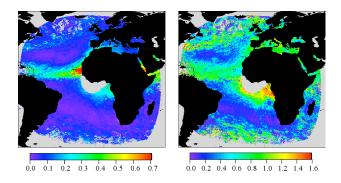


**Fig. 1.** Impact on the spatial coverage of the number of slots used to compute the daily aerosol optical thickness at 0.55  $\mu$ m for 7 June 2003 (METEOSAT-7 images). The left map has been computed with one single slot (12:15 UT), whereas the right map has been computed using nine slots (from 8:15 to 16:15 UT by step of 1). Continents are in dark grey, whereas masked marine pixels (either cloudy or contaminated by the Sun-glint) are grey.

slot (e.g. pixels on the edge of the Earth disk on the noon image). The daily map obtained using nine slots in Fig. 1 has a remarkable spatial coverage when compared to what is produced by orbiting sensors with comparable spatial resolution, such has MODIS (1 km) or POLDER (6 km). Such coverage clearly improves the quality of weekly or monthly mean aerosol products and facilitates the physical interpretation of interesting aerosol events. Such an example can be seen in Fig. 1, which shows both a dust plume from Northern Africa over the Atlantic, up to America, and a transport of biomass burning aerosols from southern Africa.

Note, however, that several areas in Fig. 1 within dustcontaminated regions (in the Persian Gulf or in the western part of the Atlantic dust plume) show significant differences when the daily mean is estimated from a single image or 9 slots. This is likely due to the fact that the aerosol optical thickness may vary significantly during the day in regions close to the sources or where the wind is strong. It is, however, also possible that part of this diurnal variation of the optical thickness is artificial and related to the wrong assumption of particle sphericity when computing aerosol optical properties throughout the Mie theory. It has indeed been shown that the effect of the non-sphericity strongly varies with the viewing geometry, and, in particular, during the day with the solar zenith angle (Masuda et al., 2002; Herman et al., 2005).

Figure 2 shows monthly mean maps of retrieved aerosol optical thickness and the Angström coefficient for June 2003, using only three slots per day instead of the nine slots planned for the final processing. Maps of the aerosol optical thickness show aerosol transport patterns which are consistent with those of the METEOSAT climatology (Moulin et al., 1997c). The dominant aerosol transports show up clearly as African dust over the Atlantic, the western Mediterranean, the Red Sea and the western Indian Ocean; as biomass burning aerosols south of the Gulf of Guinea. These two

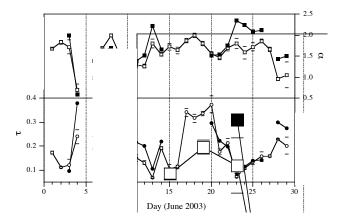


**Fig. 2.** Monthly mean maps of the aerosol optical thickness at 0.87  $\mu$ m (left) and the Angström coefficient (right) for June 2003, as retrieved from MSG/SEVIRI images using three slots per day (11:15, 12:15 and 13:15 UT). Monthly means for both parameters are computed as the arithmetic average of all clear-sky pixels of the month. MSG data are projected on a regular latitude-longitude grid, and the optical thickness is converted to a different wavelength to enable comparison with available products from POLDER (Fig. 6). The east-west dissymmetry in the images' coverage is because MSG was not yet at its operational location and was still at 10° W in June 2003. Note that only three slots per day were available at the time of this study because MSG was in its commissioning phase and we had restricted access to the data. Note also that 5 days (15–19 June) of MSG data were not available at the time of this study.

dominant aerosols are characterized by Angström coefficients of about 0.4–0.7 for mineral dust and of about 1.4–1.7 for biomass burning aerosols. Interestingly, the Angström coefficient map brings up additional information in regions where the optical thickness is small, such as the predominance of aerosols likely from pollution origin in the East Mediterranean downwind of eastern Europe or the plume of likely biomass burning aerosols coming from Amazonia over the southern Atlantic. Note that the zone contaminated by the Sun-glint does show up in these maps because three slots are used instead of 9 in Fig. 1.

#### 4 Validation using AERONET

Direct measurements of the aerosol optical thickness at several wavelengths in the visible and the near-infrared at coastal Mediterranean and Atlantic stations by the AERONET network (Holben et al., 1998) are used to validate  $\tau$  and  $\alpha$  retrieved from MSG measurements. Hourly level 2.0 (cloud-screened and quality-assured) sunphotometer measurements at six stations (see Table 2) were available for June 2003. Because Sun photometers do not always have a channel at 0.55  $\mu$ m,  $\tau$  was linearly interpolated between the two basic bands at 0.44 and 0.67  $\mu$ m, to ensure the compatibility with MSG estimates.  $\alpha$  was computed directly using the measured aerosol optical thicknesses at 0.67 and 0.87  $\mu$ m. The MSG product validated here is based on data acquired at 11:15, 12:15 and 13:15 UT. To reduce the possible noise due to the diurnal variability, AERONET

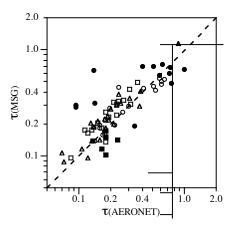


**Fig. 3.** Comparison of MSG/SEVIRI retrievals with AERONET measurements at the Forth Crete station in June 2003 for both aerosol optical thickness at 0.55  $\mu$ m (bottom) and the Angström coefficient  $\alpha$  (top). AERONET measurements are represented by opened symbols and MSG retrievals by filled symbols. The standard deviation on AERONET measurements is shown as dashes above and below each AERONET symbol. Note that 5 days (15–19 June) of MSG data were unavailable at the time of this study.

measurements taken between 10:00 and 14:00 UT only were averaged and compared to the satellite product.

The comparison between MSG retrievals and Sun photometer measurements is made by extracting  $\tau$  and  $\alpha$  over the closest  $3 \times 3$  pixel marine area distant by at least 5 kmfrom the coast, to avoid turbid water contamination. A total of 77 coincident measurements are available for the validation. Figure 3 shows the results of the comparison for the AERONET station located in Crete, in the eastern Mediterranean. As shown in Fig. 2, this region is of particular interest for aerosol studies because of the contrasted influences of Europe (pollution) and North Africa (desert dust). In addition, there is almost no cloud during summer over this region. In AERONET data,  $\tau$  varied between 0.1 and 0.4 at this station during June 2003, and  $\alpha$  varied between 0.5 and 2.0. Figure 3 shows a good agreement between MSG retrievals and ground-based measurements. The sensitivity of the MSG Angström coefficient is sufficient to allow differentiation between African dust (e.g.  $\alpha$ =0.5 on day 4) from other aerosols, even if it seems overestimated during the last week of June 2003. The agreement is also good for aerosol optical thickness, and the largest difference occurs for the African dust event of day 4. This suggests that, even if the Angström coefficient is properly retrieved on that particular day, the very specific optical properties of mineral particles (e.g. Moulin et al., 2001), which are not accounted for in our set of aerosol models, may introduce some error in the retrieval of  $\tau$ .

Figure 4 summarizes the validation results for the aerosol optical thickness at the six AERONET stations. There is no significant bias in the satellite product and 85% of the estimates are within  $\pm 30\%$  of the Sun photometer measurements. The comparison over the various sites indicates that



**Fig. 4.** Comparison of the 77 aerosol optical thicknesses at  $0.55 \,\mu$ m retrieved from coincident MSG and AERONET measurements. Open and filled symbols are respectively for Mediterranean (56 measurements; circles for Erdemli, squares for Nes Ziona and triangles for Forth Crete) and Atlantic (21 measurements; circles for Dahkla, squares for Azores and triangles for Sal island) stations (see Table 1). Note that logarithmic axes have been used to evidence small values of  $\tau$ . The slope of the linear regression performed on all data is 1.0, with a correlation coefficient of 0.83.

the product is of somewhat lower quality for the sites that are directly affected by the main dust plume from Northern Africa, which seems to confirm that the product accuracy is lower for dust. Note that the point with the largest discrepancy ( $\tau$ =0.17 for AERONET and 0.62 for MSG) is from the Dahkla data set and seems to correspond to the MSG image of a dust plume that was transported over the ocean but not over the coastal AERONET station. As usual with this kind of validation based on AERONET measurements, it is likely that a careful visual examination of individual images would enable one to remove the worst cases in Fig. 4, and thus to improve the comparison.

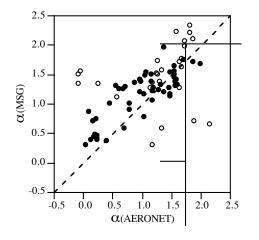
Figure 5 shows a similar comparison for the Angström coefficient. There is no significant bias but there is a much larger dispersion than for  $\tau$ . In addition to the unsuitable optical properties for mineral particles, this relative inaccuracy is mostly apparent for data with low optical thicknesses ( $\tau < 0.15$ ) when the low signal makes it difficult to assess its spectral signature. When the data set is restricted to large optical thicknesses, there is a better agreement between the surface measurement and satellite estimates of the Angström coefficient, as shown in Fig. 5.

# 5 Comparison with POLDER-2

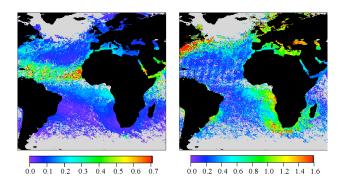
The series of POLDER (Polarization and Directionality of the Earth Reflectance) sensors are dedicated to the observation of climate-relevant parameters, which include a comprehensive set of aerosol optical properties. This instrument is particularly well suited for aerosol retrievals, thanks to its multidirectional and polarization capabilities (Deschamps et al., 1994). POLDER measurements provide a strong

Sal Island	Dahkla	Azores	Forth Crete	Erdemli	Nes Ziona	
Long.	22.93° W	15.95° W	28.63° W	25.28° E	34.25° E	34.79° E
Lat.	16.73° N	23.72° N	38.53° N	35.33° N	36.56° N	31.92° N
Country	Capo Verde	Morocco	Portugal	Greece	Turkey	Israel
P.I.	D. Tanré	H. Benchekroun	B. Holben	M. Drakakis	B. Holben	B. Holben

Table 2. AERONET stations available in June 2003 to validate MSG aerosol retrievals.



**Fig. 5.** Comparison of the 77 aerosol Angström coefficients retrieved from coincident MSG and AERONET measurements. Open and filled circles are, respectively, for data with  $\tau < 0.15$  (26 measurements) and  $\tau > 0.15$  (51 measurements). The slope of the linear regression performed on all data is 1.02, with a correlation coefficient of 0.58.



**Fig. 6.** Monthly mean maps of the aerosol optical thickness at 0.87  $\mu$ m (left) and the Angstrôm coefficient (right) for June 2003 as retrieved from POLDER-2 images.

constraint on the aerosol phase function and its spectral signature, which can be used to estimate the aerosol size distribution and non-sphericity (Deuzé et al., 2000; Herman et al., 2005), and thus to improve the accuracy of both  $\tau$  and  $\alpha$ . POLDER-1 and -2 were flown on board the Japanese satellites ADEOS-1 and -2, and provided data from November 1996 to June 1997, and from April 2003 to October 2003,

respectively. A third similar instrument, named PARASOL, has been successfully launched on board a micro-satellite in December 2004. Note that these POLDER instruments have all been flown on board polar orbiting satellites, which allow one to observe a given pixel once a day at best, whereas MSG observes the same pixel every 15 min during daytime. In contrast, POLDER provides global observations and can avoid the use of Sun-glint-contaminated measurements, thanks to its directional capability.

Briefly, the POLDER retrieval algorithm is based on measurements of both total and polarized light at 0.67 and 0.87  $\mu$ m. The best aerosol model is found by comparing these measurements to a large set of bimodal aerosol models made of various proportions of fine and coarse modes, including spherical and non-spherical particles. This set of aerosol models is by far more sophisticated and likely realistic than the one used in this study. MSG measurements are more limited than those of POLDER in terms of aerosol information content for a given pixel, and therefore do not permit one to distinguish such a large range of aerosol models.

Figure 6 shows POLDER monthly mean maps of both the aerosol optical thickness and the Angström coefficient for June 2003. These maps can be directly compared to those shown in Fig. 2. The optical thicknesses obtained with both sensors are remarkably similar. The transport of African dust over the north tropical Atlantic, the Red Sea and the northern Indian Ocean is clearly depicted, as well as the pollution plume from North America over the North Atlantic. The two sensors retrieve similar optical thicknesses of biomass burning aerosols retrieved over the Guinea Gulf. The monthly mean image of optical thickness from MSG shows much less small-scale structures - most probably noise - than that of POLDER. Similarly, the coverage by MSG is much better. This observation confirms that the use of several images per day strongly improves the quality of the aerosol monitoring, even though Fig. 2 has been generated using only three slots per day while many more are available. The gap in coverage at the image center, south of the coast of Senegal, is a consequence of the Sun-glint contamination and would be filled when earlier and later slots are used.

The differences between the two sensors in Figs. 2 and 6 are much more important for the Angström coefficient, particularly for the dust plume over the tropical Atlantic that is characterized by values of  $\alpha$  of about 0.8 for MSG and

of about 0.2 for POLDER. Even if the POLDER map is particularly noisy in the dust zone, it seems that the MSG Angström coefficients are overestimated in a region where values of 0.2-0.5 are often reported (Moulin et al., 1997b). Such a difference has still to be understood, but it might be due once again to the lack of realistic models for mineral particles in the LUT. Other regions show significant differences between the two sensor products, such as around South Africa or Madagascar, where the O99 aerosol model is imposed in the inversion because digital counts are too small to ensure the accuracy of the results. The two sensors are, on the contrary, in relatively good agreement with the pollution plume from North America and with the biomass aerosols in the Guinea Gulf. The overall agreement on  $\alpha$ , however, remains relatively poor between MSG and POLDER. A possible explanation is that the MSG calibration is still preliminary and a small error on the reflectance at 0.63  $\mu$ m, for example, will not greatly affect the retrieval of the optical thickness but will modify significantly the Angström coefficient.

#### 6 Conclusions

An inversion method has been developed to estimate the aerosol optical thickness and the Angström coefficient over the ocean from MSG/SEVIRI measurements at 0.63 and 0.81  $\mu$ m. This is achieved using a standard "best–match" comparison of MSG reflectances to a set of pre-computed simulated TOA reflectances. For efficiency purposes, TOA reflectances were computed for a set of only 15 aerosol models that enables a rapid search of the best solution. We applied this inversion technique to slots 45, 49 and 53 (11:15, 12:15, 13:15 UT) of all available days of June 2003.

Monthly mean maps of both parameters exhibit clearly the dominant aerosols, i.e. mineral and biomass burning aerosols, as well as the influence of pollution from Europe over the Mediterranean. A comparison with the AERONET in-situ measurements performed in the tropical Atlantic and in the Mediterranean shows that MSG retrievals have a good accuracy in terms of optical thickness, while the Angström coefficient shows some dispersion. The comparison with POLDER products for June 2003 confirms that the retrieved optical thickness is of good quality, whereas the two sensor estimates of the Angström coefficient show significant differences. Potential explanations are the uncertainties on the MSG calibration at this early stage of its lifetime and the lack of a realistic model for mineral dust in the simulated reflectances.

The preliminary results presented here, however, demonstrate the strong potential of MSG for the characterization of marine aerosols, in particular, because several images per day are available. Such a high frequency of observations make MSG aerosol products very complementary to those of the polar orbiting sensor, such as POLDER and MODIS, even though the latter are better suited to retrieve detailed information on the aerosol size distribution. Thanks to the efficiency of the inversion code, we aim at processing nine MSG slots per day in full resolution over several months. A potential improvement would come from the use of the MSG 1.64  $\mu$ m band to provide additional information and better constrain the aerosol model, in particular, for large mineral particles. Another possible improvement of the method could be to use ancillary data from meteorological models to refine the Sunglint removal through the wind speed and the atmospheric gaseous absorption estimate through ozone and water vapor content.

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