

Letter to the Editor

Low-frequency electric field fluctuations and field-aligned electron beams around the edge of an auroral acceleration region

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Abstract. Electron beams narrowly collimated to the magnetic field line were observed continuously from a downward current region to an auroral acceleration region (i.e., upward current region). They were well correlated with low-frequency electric field fluctuations in the auroral acceleration region as well as in the adjacent downward current region. Magnetic field fluctuations were found only in the downward current region. The analysis suggests that static field-aligned electric fields are not fully responsible for the field-aligned electron acceleration; the ac electric field, presumably associated with Alfvénic fluctuations, should also be involved in the acceleration of ionospheric electrons.

Key words. Ionosphere (particle acceleration) – Magnetospheric physics (auroral phenomena; magnetosphere-ionosphere interactions)

1 Introduction

Upgoing, counterstreaming, and downgoing field-aligned electron beams (FEB) are often associated with upflowing ions with conical pitch angle distribution (i.e., ion conics) over the auroral region (Sharp et al., 1980; Collin et al., 1982; Miyake et al., 1998). FEB are also found with upflowing ion beams in the auroral acceleration region (Hultqvist et al., 1991; Lundin and Eliasson, 1991; Yoshioka et al., 2000).

Upward acceleration of ionospheric electrons by a downward electrostatic field (Gorney et al., 1985) associated with a downward field-aligned current was presented from recent FAST observations (Carlson et al., 1998). The field-aligned electrostatic field is developed along the magnetic field line in order to supply the charge carrier to the current. This is even true for the downward current regions (Elphic et al., 1998).

An alternative explanation for the generation of FEB was provided by Viking observations (Hultqvist et al., 1991). Low-frequency electric field fluctuations (LEF) are found to be

well correlated with FEB. The LEF are supposed to propagate down from the magnetosphere (Lundin et al., 1990) and, therefore, to be Alfvénic. The possible parallel component of LEF (Goertz and Boswell, 1979) may generate FEB.

We present here an event of FEB extended from the downward current region to the adjacent auroral acceleration region (i.e., upward current region). The event analyzed here is the same as the example of co-existence of FEB and upflowing ion beams presented in Yoshioka et al. (2000). This report focuses on electromagnetic aspects of the event in terms of the generation mechanism of FEB.

2 Observations

Data presented here come from the polar-orbiting Akebono (Exos-D) satellite (Oya and Tsuruda, 1990; Tsuruda and Oya, 1991). The satellite was launched with an initial apogee and perigee, respectively, of 10482 km and 272 km in 1989. The Low Energy Particle (LEP) instrument on board (Mukai et al., 1990) was designed to observe energy-pitch angle distributions of auroral electrons and ions. We also use electric field (Hayakawa et al., 1990) and magnetic field (Fukunishi et al., 1990) data in this study.

Figure 1 shows an Akebono equatorward pass through the auroral oval around 4.6 MLT on 7 December 1989. Counterstreaming FEB with energies below 1 keV are observed for most of the interval. Discrete electron precipitation begins at 6:42:30 and the field-aligned current (FAC) derived from magnetic field data changes direction from downward to upward. Ion conics, with low-energy cutoff, are observed in the downward FAC and upflowing ion beams with a field-aligned peak of flux beginning at 6:43:00 in the upward FAC. The FEB show a rapid variation but are continuously observed into the edge of the ion beam region. A statistical study on the co-existence of FEB and ion beams is presented in Yoshioka et al. (2000).

Intense LEF, perpendicular to B_0 (the Earth's magnetic field), accompany the FEB. We take Z as upward direction

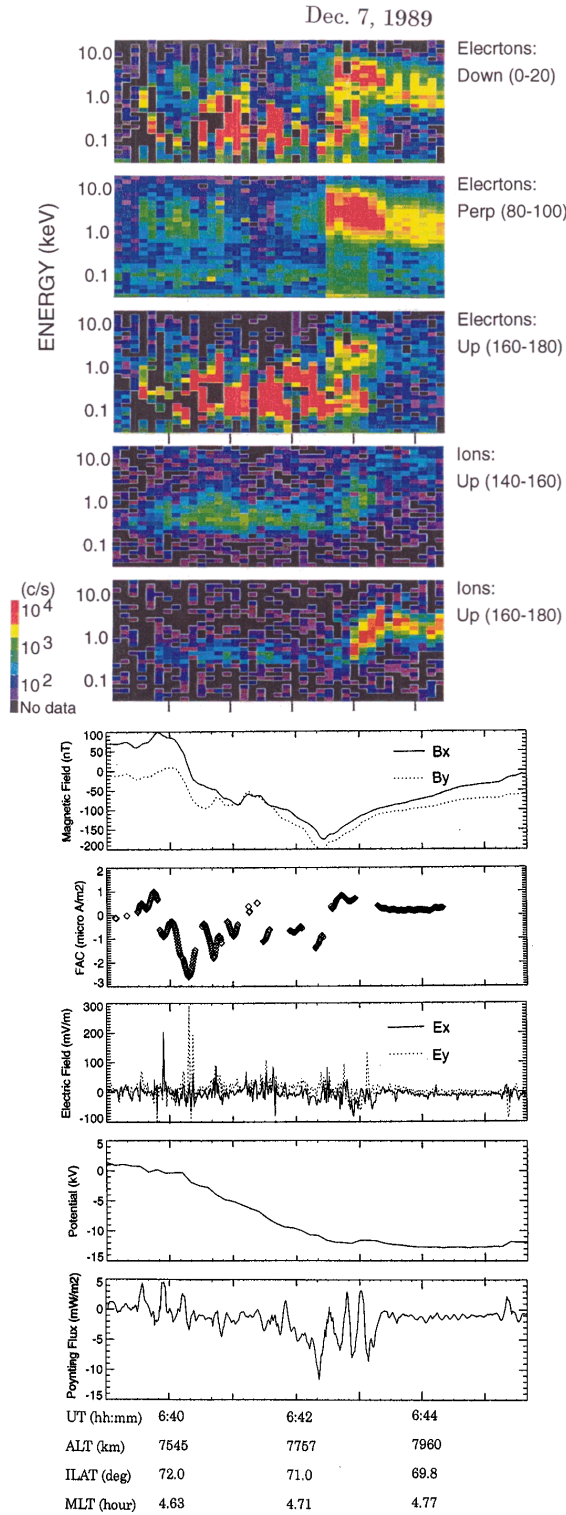


Fig. 1. Overview of a field-aligned electron event observed with Akebono on 7 December 1989. From top to bottom: Energy-time diagrams of downgoing (pitch angle of 0–20°), perpendicular (80–100°), and upgoing (160–180°) electrons, and of upgoing (140–160° and 160–180°) ions. Two components of magnetic field. Field-aligned current derived from the magnetic field observation. Two components of electric field. Electric potential calculated by the integration of the observed electric field along the satellite path. Poynting flux along the geomagnetic field line.

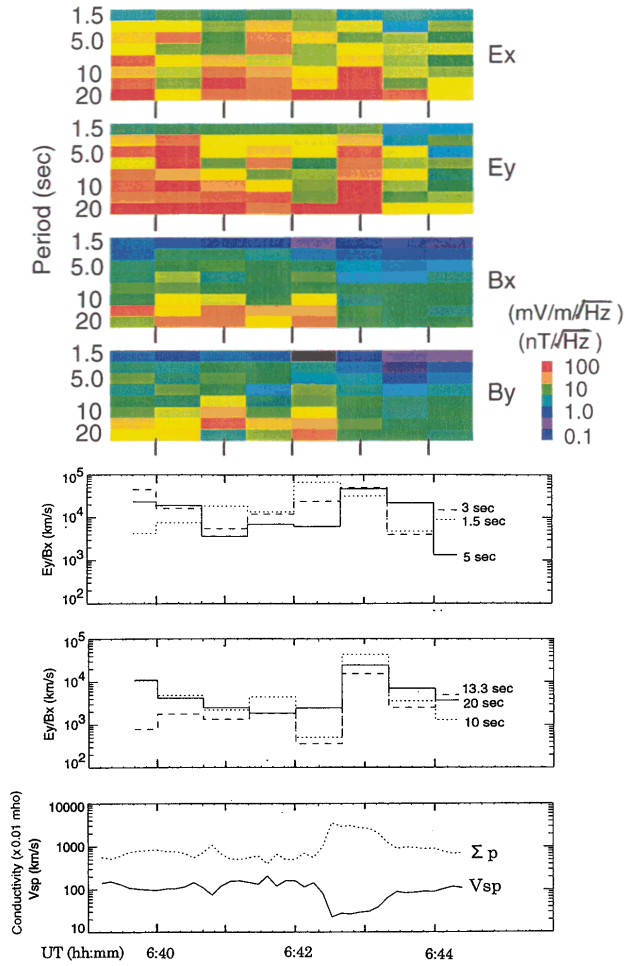


Fig. 2. From top to bottom: Spectral density of the two components of electric field and magnetic field measured by Akebono satellite. E/B ratios for fluctuations of 6 wave periods. Height-integrated Pederson conductivity and $V_{sp} = 1/\mu_0 \Sigma_p$.

along the magnetic field line. X and Y denote here the east (-west) and north(-south) direction, respectively. The FEB terminated at 6:43:16 and the intensity of LEF decreased then. The electric potential is calculated from the integration of an electric field component along the satellite path.

Akebono measures the electric field in two dimensions perpendicular to the spin axis. We assume no parallel electric field (i. e., $E_z = 0$) to derive the E_X and E_Y . Even when there is a substantial parallel field accelerating electrons, it is just a small fraction of the total electric field and does not significantly affect E_X and E_Y .

8-sec averaged Poynting flux in the direction of B_0 are fluctuated and there is a large-scale negative peak around the current reversal. The negative peak implies a downward flux associated with the dc current circuit of the high-latitude downward and low-latitude upward current system.

3 Analysis on LEF

It is known that there are two origins of low-frequency electric fluctuations measured by a moving probe in the auroral zone: static structure of FAC and Alfvén waves (see, for example, Sugiura et al., 1982; Gurnett et al., 1984; Weimer et al., 1985; Knudsen et al., 1990; Nagatsuma et al., 1996; Aikio et al., 1996). When we take the electric perturbation E and magnetic perturbation B perpendicular each other, there is a difference in the ratio of E to B between the two cases. If the fluctuations are due to static structures, then

$$\frac{E}{B} = \frac{1}{\mu_0 \Sigma_p} + \frac{k^2}{\mu_0 a} \quad (1)$$

where Σ_p is the height-integrated Pederson conductivity, k is the characteristic spatial wave number, and $a = J_{\parallel}/V_{\parallel}$ (field-aligned conductivity). Alfvén waves give the relation

$$\frac{E}{B} = V_A \quad (2)$$

When upgoing and downgoing waves are interfering, then the ratio can be different from V_A .

Figure 2 shows the results of spectral analysis of the electric and magnetic field fluctuations. As pointed out in Fig. 1, LEF continue beyond the current reversal at 6:43:00 while magnetic field fluctuations stop before 6:43:00.

We take the ratio of E_Y (north) to B_X (east) here. V_{sp} ($= 1/\mu_0 \Sigma_p$) stands for the static structures of FAC with no field-aligned potential drop and is derived from precipitating electrons (Robinson et al., 1987).

The observed ratio is generally higher for fast fluctuations. V_A is estimated to be 10^4 – 10^5 km/s at the altitude of 7000–8000 km. The ratio is smaller for static structures. Knudsen et al. (1990) reported that fast fluctuations are more Alfvénic and slow fluctuations are due to static structures.

The observed ratio is about 10^3 km/s in the downward current region and about 10^4 – 10^5 km/s in the upward current region. V_{sp} is 100–200 km/s in the downward current region and 20–100 km/s in the upward current region. The E/B ratio in the downward FAC is close to V_{sp} . Since a in Eq. (1) can be assumed from the observations (J_{\parallel} and the energy of FEB) and $k = 2\pi/V_s T$ where V_s is the satellite velocity and T is the wave period, the second term of the right-hand of Eq. (1) is estimated to be about 10^4 km/s for a period of 10 sec. Therefore, taking into account the effect of field-aligned potential drop below the satellite, the observed magnetic field fluctuations can be basically interpreted in terms of static current structure. The FAC, derived from MGF data in Fig. 1, actually contains large fluctuations. Elphic et al. (1998) reported that the spatial scale of static structure of downward FAC is smaller than that of the upward FAC.

We also examined the phase difference between the electric and magnetic fluctuations. No phase difference is expected for pure static structures regardless of the presence of the parallel potential drop and for Alfvén waves traveling toward the field line direction. The difference should be 180° for Alfvén waves traveling toward the inverse direction of the

field line. It can be different from the ideal case for both the static structures superposed by Alfvén waves and the interfering downcoming/upgoing Alfvén waves. The phase difference scatters over all the wave periods analyzed here in both the upward and downward current regions.

The E/B ratio is increased in the upward current region, while V_{sp} is decreased there. Although the potential drop below the satellite may be able to account for the E/B ratio, the large deviation of the phase difference there can be only attributed to the presence of Alfvén waves. The large fluctuations of Poynting flux (the bottom of Fig. 1), especially at the edge of the auroral acceleration region, can also result from the downcoming and upgoing Alfvén waves interfering with one another. Nagatsuma et al. (1996) reported large fluctuations of Poynting flux due to the interfering Alfvén waves often found around the poleward region of the auroral oval.

4 Discussion

LEF are well correlated with FEB in the downward current region as well as in the upward current region. Magnetic field fluctuations are only found in the downward current region. The analysis of the E/B ratio suggests that the low-frequency magnetic fluctuations in the downward current region are mostly due to static structure of the FAC while the LEF in the upward current region come mostly from Alfvén waves. Yoshioka et al. (2000) argued that wave ac electric field is most likely to accelerate ionospheric electrons upward below the auroral acceleration regions.

The origin of the downward component of the counterstreaming FEB in downward current regions is still an open question. Carlson et al. (1998) suggested that the upward FEB originated in the conjugate auroral region coming down along the field line. Since some pitch angle scattering is expected during the travel from the other hemisphere, the electron beams originating from the other hemisphere can hardly explain the narrow pitch angle distribution of the downward component, almost equivalent to that of the upward component. The continuous extension of FEB from the downward current to the upward current regions and the deviation of the phase difference from both the pure static structures and traveling Alfvén waves indicate that parallel ac electric field associated with Alfvén waves, not only field-aligned electrostatic field, may be responsible for some part of field-aligned electron acceleration and reflection in the downward current region.

If the reflection of FEB, on the same auroral field line just above the satellite really takes place, there should be some relationship between the downward and upward components since both the components have the same origin. Figure 3a compares the upward and downward energies of counterstreaming electrons. The filled diamonds represent data in the upward FAC after 6:42:30. The minimum energies (right) are almost identical (correlation coefficient: 0.88) while the maximum (left) show rather a large scatter (correlation coefficient: 0.45).

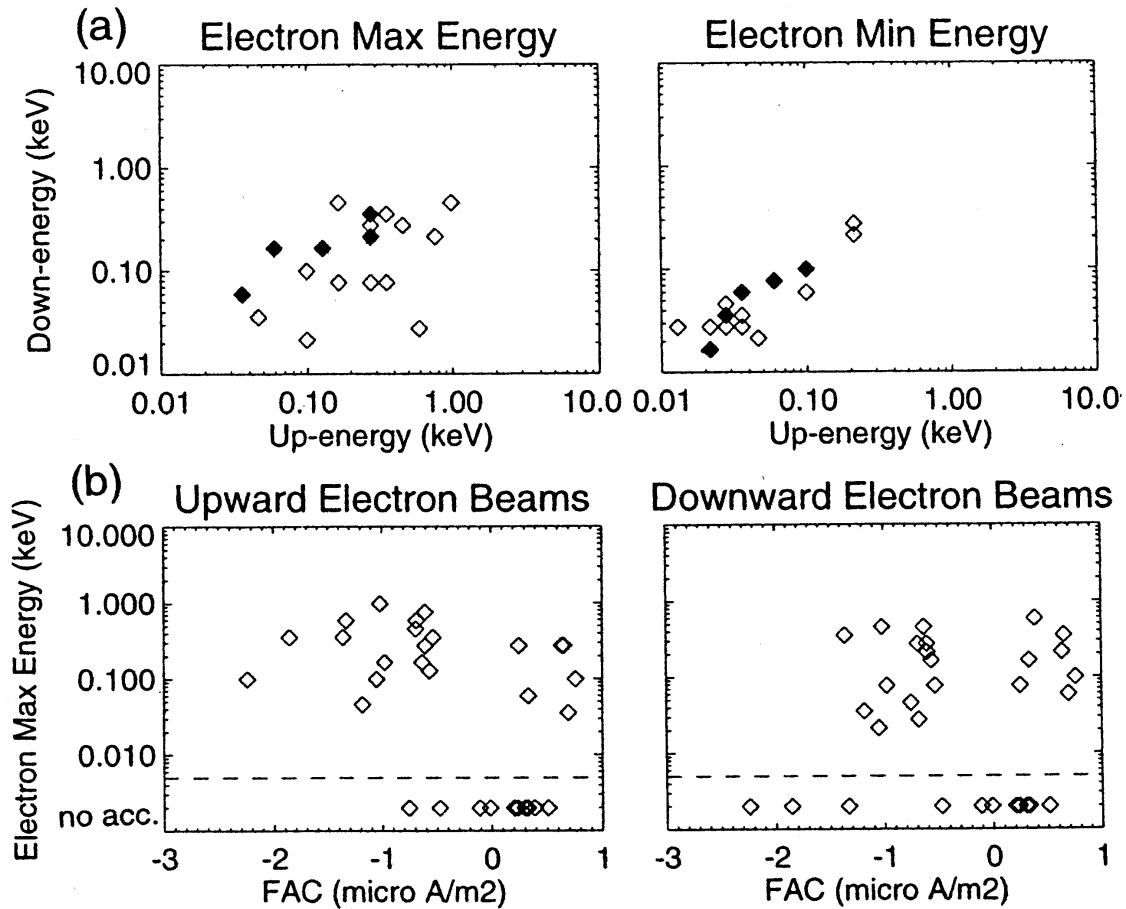


Fig. 3. (a) Comparison of upward and downward energies of counterstreaming electrons. The filled diamonds stand for the data after 6:42:30. (b) Relationship between the intensity of field-aligned current and the maximum energy of upward (left panel) and downward (right) field-aligned electron beams.

In the upward current region, a parallel ac field is expected below the electrostatic auroral acceleration region. A strong upward dc field in the auroral acceleration region completely reflects the ionospheric electrons accelerated upward by the ac field and symmetrical counterstreaming FEB are formed (Yoshioka et al., 2000).

Parallel ac field above the satellite is expected to reflect the upgoing FEB in the downward current region. Since the net current is downward, the ac field at high altitude is assumed to be smaller than the downward dc field at low altitude. The ac field reflects only the lower-energy portion of the FEB in this case and the downgoing FEB is not just a mirror image of the upgoing FEB. The minimum energies of both components of FEB are quite similar.

Figure 3b shows the relation between the magnitude of field-aligned current (FAC) and the maximum energy of FEB. Elphic et al. (1998) reported the current-voltage relation of downward FAC. The event analyzed here supports the relation with a large scatter. The acceleration modified by a superposed ac field may result in the large scatter.

The lower energy cutoff of ion conics is rather stable (the forth panel of Fig. 1) while the maximum energy of FEB

is quite variable. The lower energy cutoff of ion conics corresponds to the downward potential drop below the satellite (Carlson et al., 1998). The inconsistency of the lower cutoff energy of ion conics and the upward energy of FEB also agrees with the idea of the coexistence of both ac and downward dc electric fields on the same field line. The electrons are accelerated by both the ac and dc fields, while the ions are accelerated only by the dc field (Hultqvist et al., 1991).

The Alfvén waves probably propagate down from the magnetosphere and are reflected back at the ionosphere. Ivchenko et al. (1999), however, reported that the motion of discrete electron precipitation, which modifies the ionospheric conductance and current, leads to the launch of Alfvén waves upward from the ionosphere. We do not know the motion of the discrete electron precipitation region in the event, but it is possible that a part of upward-propagating waves are generated by the modification of the conductance around the edge of the intense precipitation.

It has also been suggested that electromagnetic ion cyclotron waves generate FEB (Temerin et al., 1986), but they accelerate electrons downward. The ion cyclotron waves are excited by electron beams of inverted-V precipitations (Er-

landson and Zanetti, 1998). Therefore, it is difficult to account for the upward acceleration of electrons below the auroral acceleration regions and downward acceleration in the diffuse precipitation regions.

In summary, the low-frequency electric field fluctuations observed around the edge of an auroral acceleration region are Alfvénic, suggesting that the parallel component of the wave electric field in the confined region may be responsible for the generation of FEB. The FEB in the adjacent downward current region are, on the other hand, basically accelerated upward by the downward field-aligned electrostatic field below the satellite as proposed by FAST observations. The origin of the associated downward FEB there is still an open question. One possibility is that the same ac acceleration mechanism as that around the edge of auroral acceleration regions is also operative, superposed over the acceleration by field-aligned static potential drop.

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